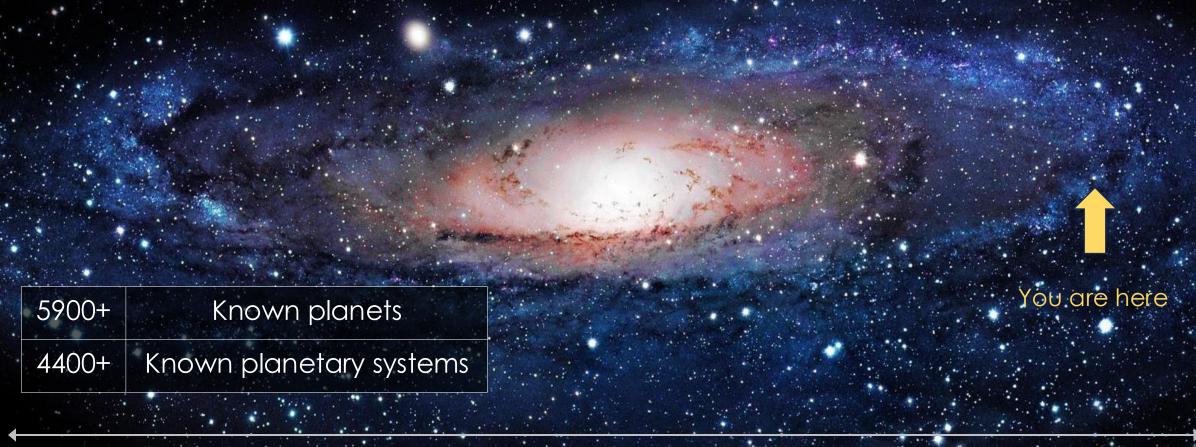


GIOVANNA TINETTI (KCL), ARIEL TEAM, BSSL TEAM



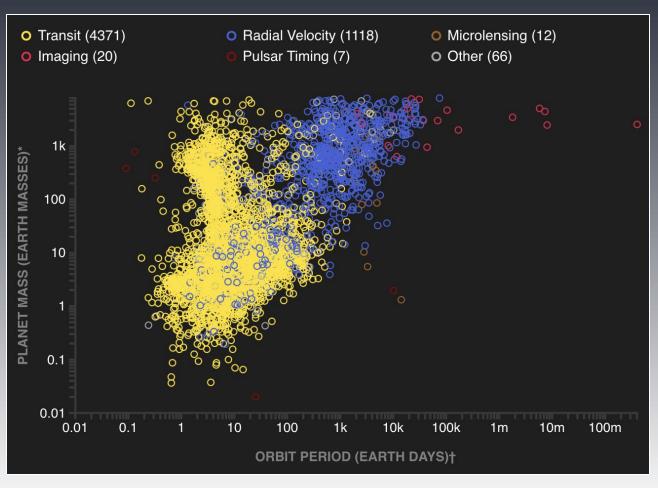
Planets are around most stars

There are at least as many planets as stars ... billions in our galaxy



Planets are very diverse

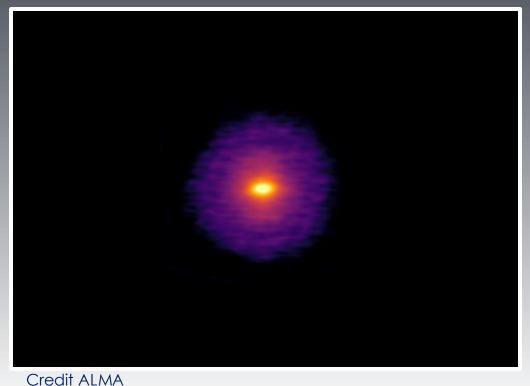
Exoplanets show greater diversity than their Solar siblings, why?



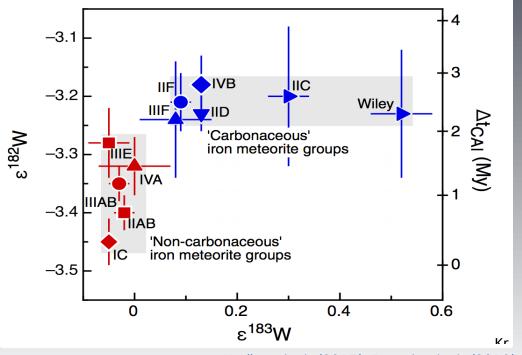
Planets are very diverse

Diversity must come from formation & evolution processes

☆ Structure and chemistry of discs from ALMA + direct imaging observations in NIR, MIR



★ Isotopic ratios from meteorites + in-situ measurements in Solar System

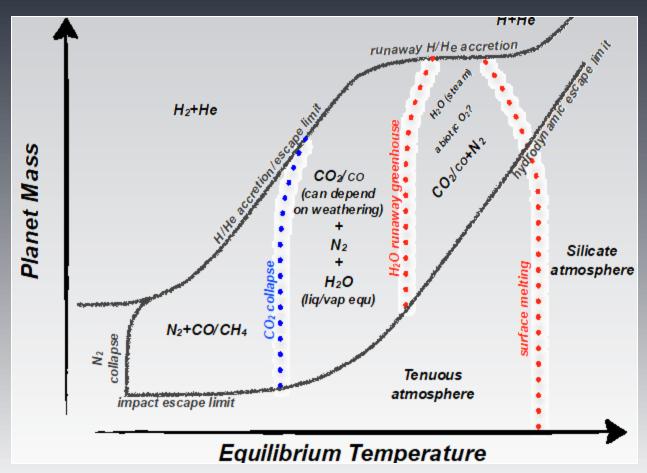


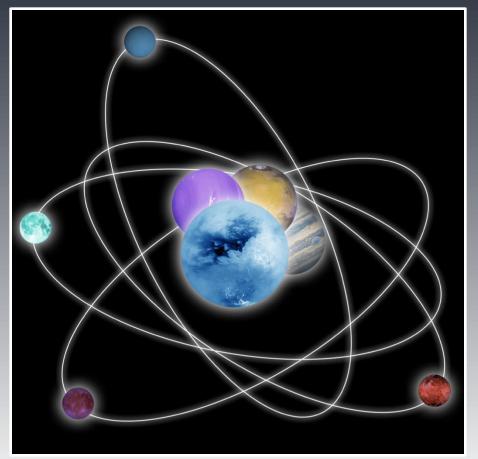
Kruijer et al. (2017); Desch et al. (2018)

Sagan workshop – 2025

Chemical diversity?

Correlation with other key parameters? Where are the transitions?

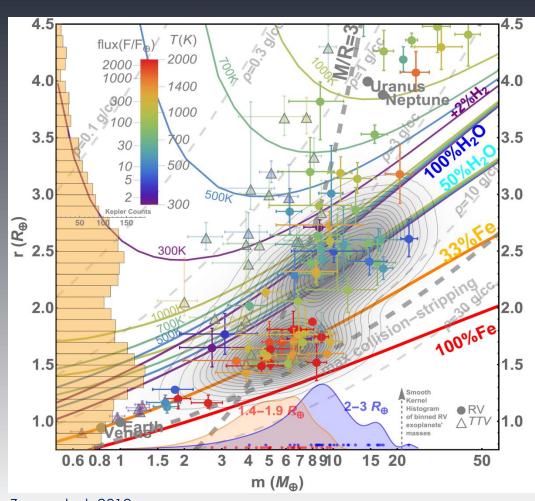


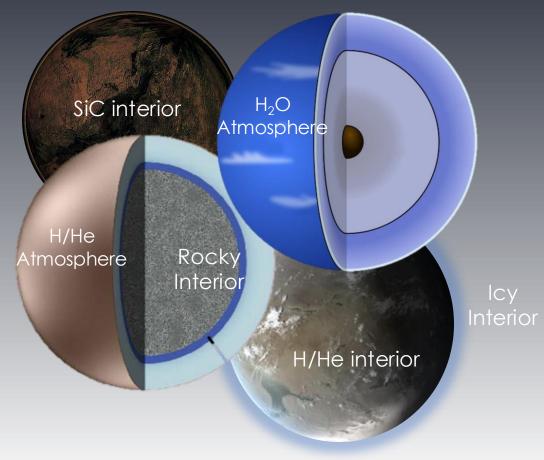


Forget and Leconte 2013

Planetary bulk composition

Many solutions compatible with the density data

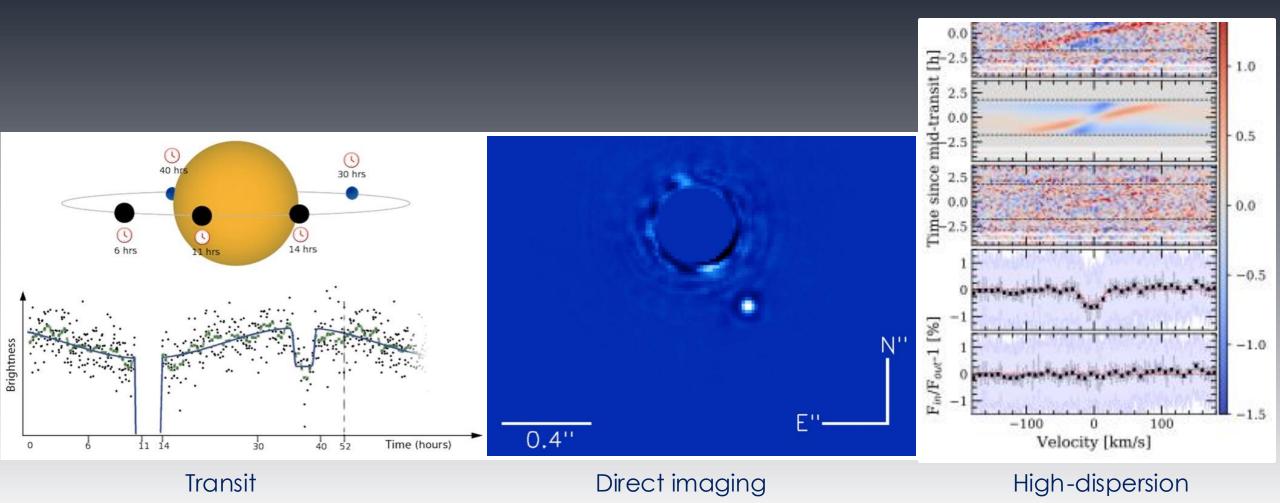




Zeng et al. 2019

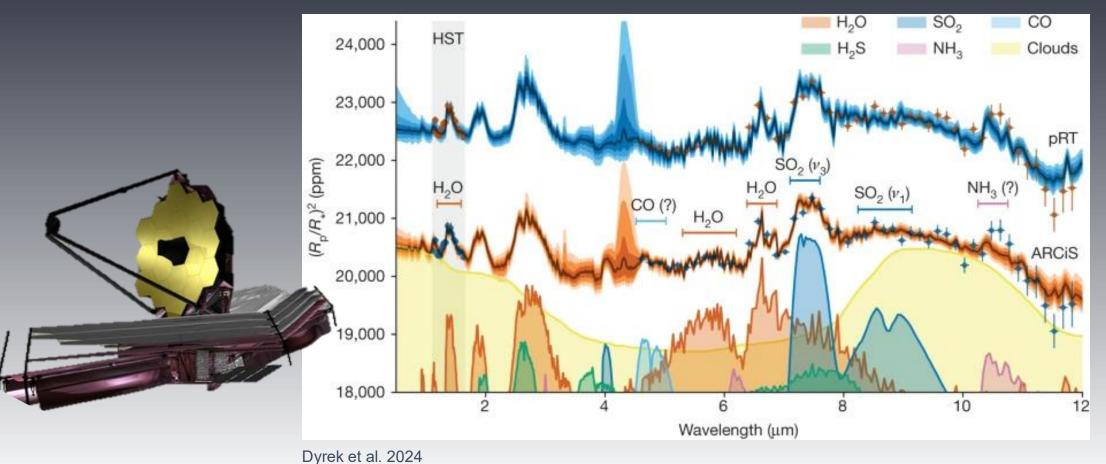
Atmospheric composition

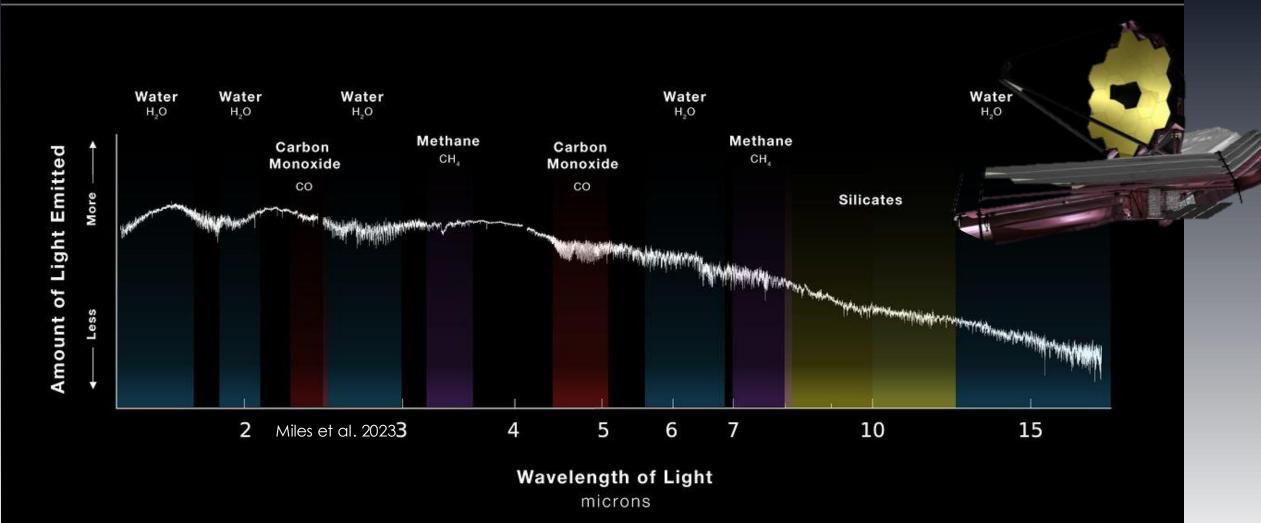
No more science fiction! Many different techniques...



Transit spectroscopy

WASP 107b



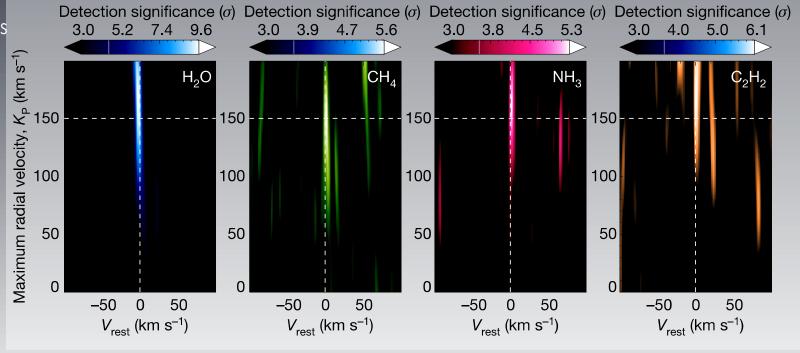


Exo-atmospheres in HD

High-spectral resolution from the ground

Many atomic and ionic species in gaseous planets' atmospheres

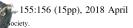




Giacobbe et al. 2021

Population studies: H-rich atmospheres

Beyond individual planets:
data for 70+ giant exoplanets analysed with Bayesian statistics



A. Tsiaras ¹, I. P. Waldmann ¹, T. Zingales ^{1,2}, M. Rocchetto ¹, G. Morello ¹, M. Damiano ^{1,2}, [5], L. K. McKemmish ¹, J. Tennyson ¹, and S. N. Yurchenko ¹, Department of Physics & Astronomy, University College London, Gower Street, WC1E6BT London, UK; and ²INAF—Osservatorio Astronomico di Palermo, Piazza del Parlamento ¹, 1-90134 Palermo,

³ Department of Physics, Section of Astrophysics, Astronomy and Mechanics, Aristotle University of Thessaloniki,

THE ASTRONOMICAL JOURNAL, 169:32 (16pp), 2025 January © 2024. The Author(s). Published by the American Astronomical Society.

OPEN ACCESS

A Comprehensive Analysis of Spitzer 4.5 μ m Phase Curves of Hot Jupiters

A Population Study of Gaseous Exoplanets

Lisa Dang^{1,21}, Taylor J. Bell^{2,3}, Ying (Zoe) Shu¹, Nicolas B. Cowan^{4,5}, Jacob L. Bea Eliza M.-R. Kempton⁷, Megan Weiner Mansfield^{8,22}, Emily Rauscher⁹, Vivien Parmentier Kevin B. Stevenson¹¹, Mark Swain¹², Laura Kreidberg¹³, Tiffany Kataria¹², Jean-Michel De Jonathan J. Fortney¹⁶, Nikole K. Lewis¹⁷, Michael Line¹⁸, Caroline Morley¹⁹, an Trottier Institute for Research on Exoplanets and Département de Physique, Université de Montréal, 1375 Avenue Thérèse-L

Exploring the Ability of HST WFC3 G141 to Uncover Trends in Populations of Exoplan Homogeneous Transmission Survey of 70 Gaseous Planets

BILLY EDWARDS,^{1,2,*} QUENTIN CHANGEAT,^{2,3,4,*} ANGELOS TSIARAS,^{5,2} KAI HOU YIP,² AHMED I MICHELLE F. BIEGER,⁶ AMÉLIE GRESSIER,^{7,8,9} SHO SHIBATA,¹⁰ NOUR SKAF,^{9,11,2} JEROEN BOU MASAHIRO IKOMA,^{14,15} OLIVIA VENOT,¹⁶ INGO WALDMANN,² PIERRE-OLIVIER LAGAGE,¹

https://doi.org/10.3847/1538-3881/aaaf75



THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES, 276:70 (55pp), 2025 February © 2025. The Author(s). Published by the American Astronomical Society.

OPEN ACCESS

OPEN ACCESS



Arianna Saba ®, Alexandra Thompson ®, Kai Hou Yip ®, Sushuang Ma ®, Angelos Tsiaras ®, Ahmed Faris Al-Refaie ®, and Giovanna Tinetti ®

Department of Physics and Astronomy, University College London, Gower Street, WC1E 6BT London, UK Received 2024 April 23; revised 2024 October 25; accepted 2024 October 27; published 2025 February 6

rossmark

THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES, 260:3 (49pp), 2022 May © 2022. The Author(s), Published by the American Astronomical Society.

https://doi.org/10.3847/1538-4365/ac5cc2



Five Key Exoplanet Questions Answered via the Analysis of 25 Hot-Jupiter Atmospheres in Eclipse

Q. Changeat^{1,11}, B. Edwards^{1,2,11}, A. F. Al-Refaie¹, A. Tsiaras^{1,3}, J. W. Skinner⁴, J. Y. K. Cho⁵, K. H. Yip¹, L. Anisman¹, M. Ikoma^{6,7}, M. F. Bieger⁸, O. Venot⁹, S. Shibata¹⁰, I. P. Waldmann¹, and G. Tinetti¹, Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, UK; quentin.changeat.18@ucl.ac.uk

¹ Department of Physics and Astronomy, University College London, Gower Street, London WC1E 6BT, UK; quentin.changeat.18@ucl.ac.uk

² AIM, CEA, CNRS, Université Paris-Saclay, Université de Paris, F-91191 Gif-sur-Yvette, France

³ INAF, Osservatorio Astrofisico di Arcetri, Largo E. Fermi 5, I-50125 Firenze, Italy

⁴ School of Physics and Astronomy, Queen Mary University of London, Mile End Road, London E1 4NS, UK

⁵ Center for Computational Astrophysics, Flatiron Institute, 162 Fifth Avenue, New York, NY 10010, USA

⁶ Division of Science, National Astronomical Observatory of Japan, 2-21-1 Osawa, Mitaka, Tokyo 181-8588, Japan

⁷ Department of Astronomical Science, The Graduate University for Advanced Studies (SOKENDAI), 2-21-1 Osawa, Mitaka, Tokyo 181-8588, Japan

⁸ College of Engineering, Mathematics and Physical Sciences, University of Exeter, North Park Road, Exeter, UK

⁹ Université de Paris and Univ Paris Est Creteil, CNRS, LISA, F-75013 Paris, France

O University of Zurich, Winterthurerstr. 190, 8057 Zurich, Switzerland

Received 2021 August 9; revised 2022 January 31; accepted 2022 February 28; published 2022 April 25

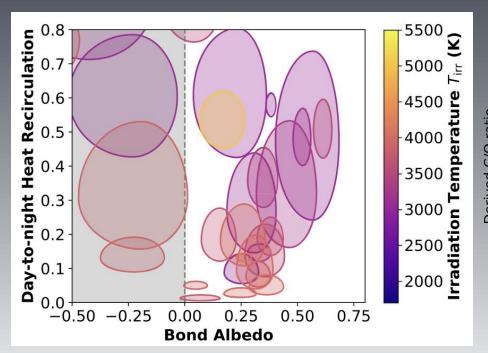
Sagan workshop – 2025

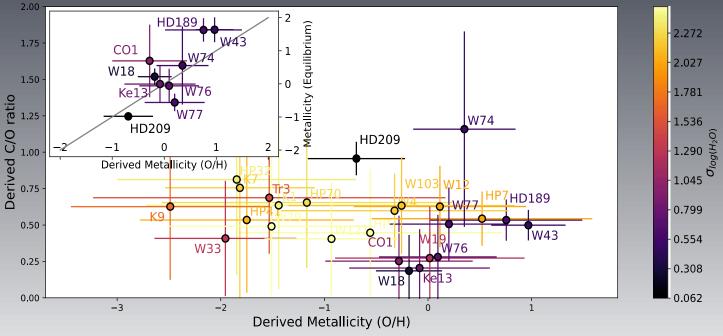
11

Population studies: H-rich atmospheres



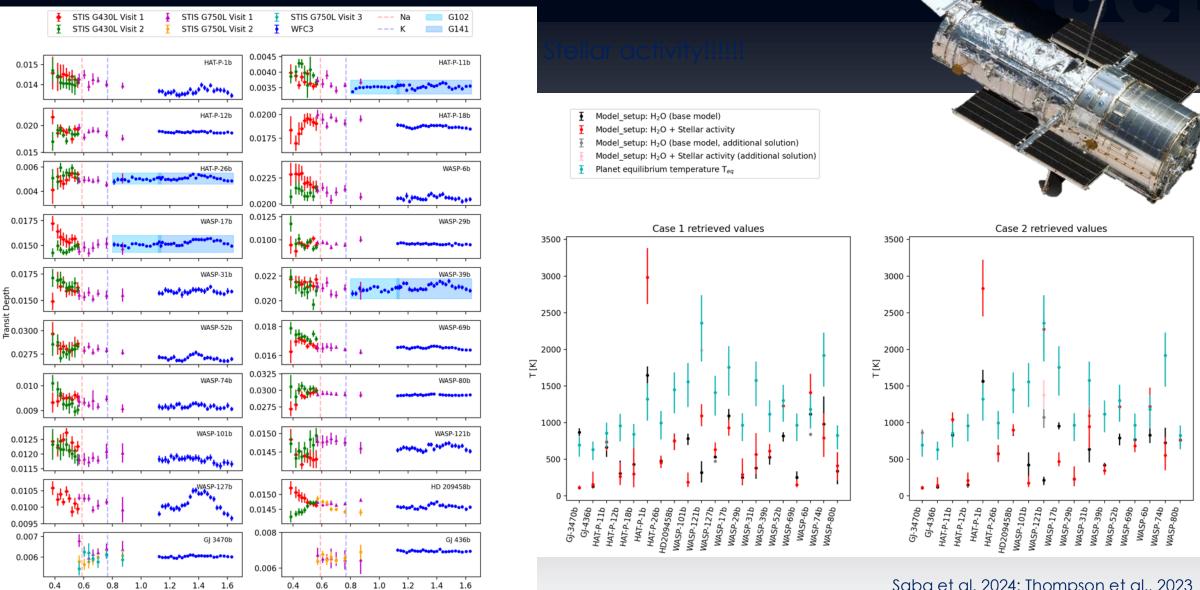
Beyond individual planets: data for 70+ giant exoplanets analysed with Bayesian statistics





Edwards et al. 2022; Tsiaras et al. 2018; Changeat et al. 2022; Dang et al. 2024

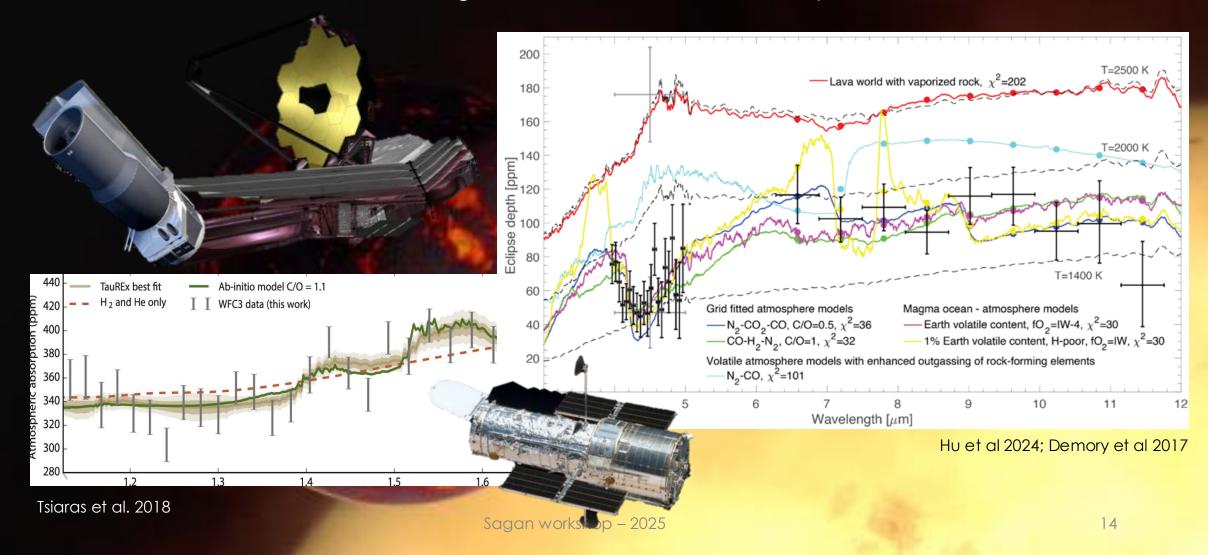
Population studies: H-rich atmospheres



55 Cnc e

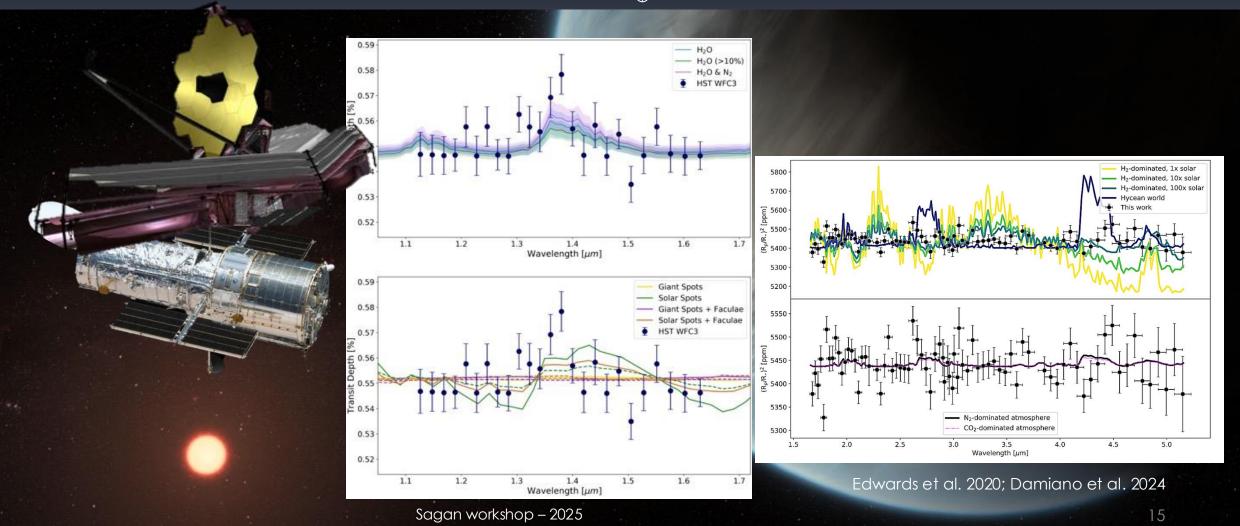


2000 K hot! Magma ocean? There is an atmosphere...



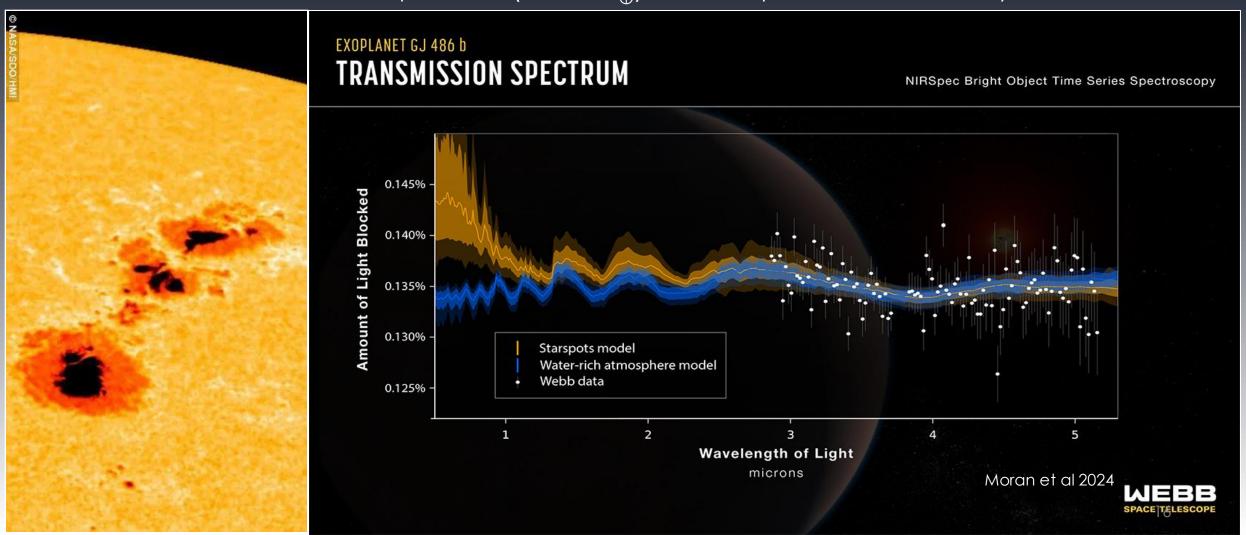
LHS 1140b

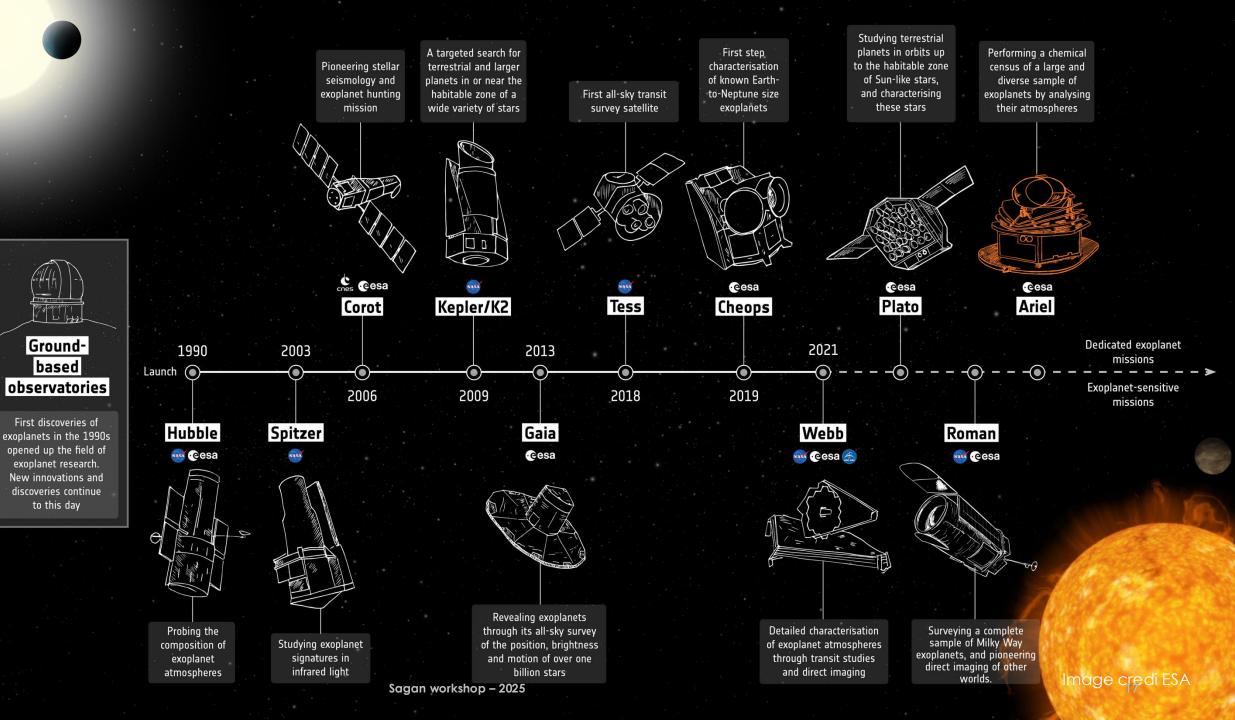
T ~ 235 K; super-Earth (R = 1.7 R_{\oplus}) An atmosphere? Stellar activity?



GJ 486b

T ~ 700 K; super-Earth (R = 1.3 R_{\oplus}) An atmosphere? Stellar activity?





Ariel: a chemical survey

- Adopted as ESA M4 in Nov. 2020
- Launch to L2 in 2029
- 1m-class telescope
- Simultaneous coverage 0.5-7.8 μm
- ~1000 exoplanets observed
- Rocky + gaseous; 300-3000K;
- stars A-M



Ariel Definition Study Report – Tinetti et al. 2021, arXiv:2104.04824











































Ariel spacecraft & payload

Payload delivered by a consortium of 16 ESA countries, NASA, JAXA, CSA



Ariel spacecraft & payload

Payload integrated at RAL in Didcot. Spacecraft from Airbus















































Ariel payload consortium



600+ scientists and engineers from 16 ESA countries + NASA, JAXA, and CSA



A mission is more than flying hardware... there are people



































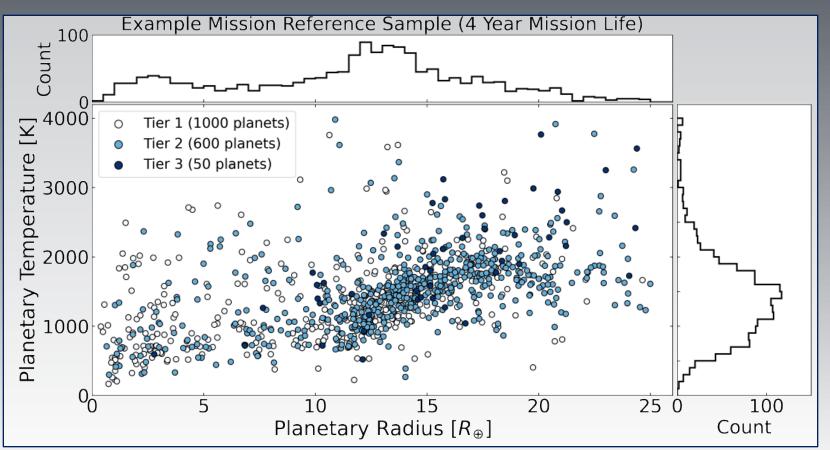




Ariel target candidates (MCS)



Ariel Mission Candidates Sample (MCS) available on Github





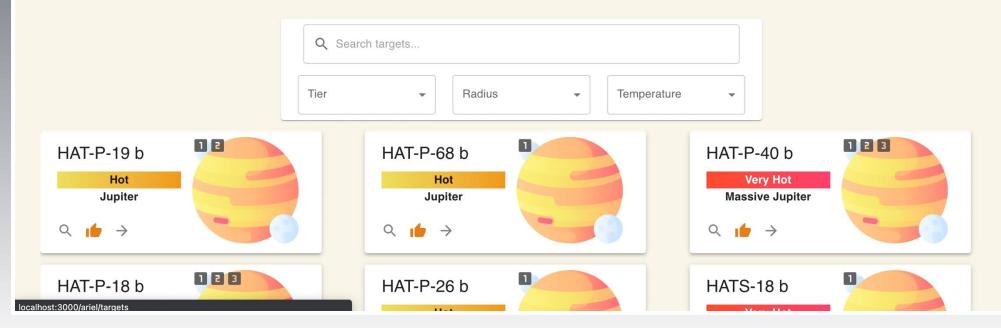
Edwards et al. 2022

Ariel target candidates

Ariel I

Catalogue Available soon through a new interactive, well maintained website

Ariel Target Candidate List

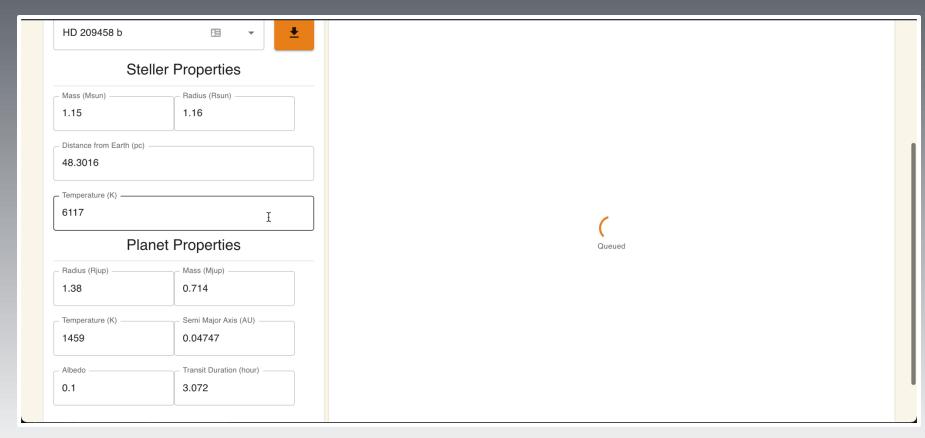


Credit: Al-Refaie & Nambiyath Govindan

Ariel target candidates

Arel 1

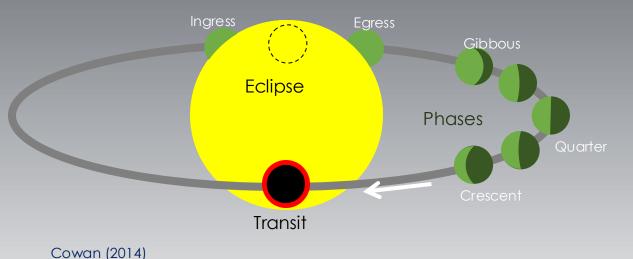
Catalogue Available soon through a new interactive, well maintained website

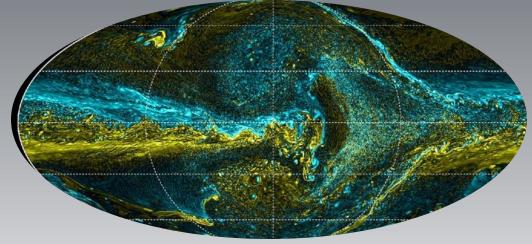


Mugnai et al, 2022

Planets are 4D complex objects

Variability in space and time: phase-curves & repeated observations



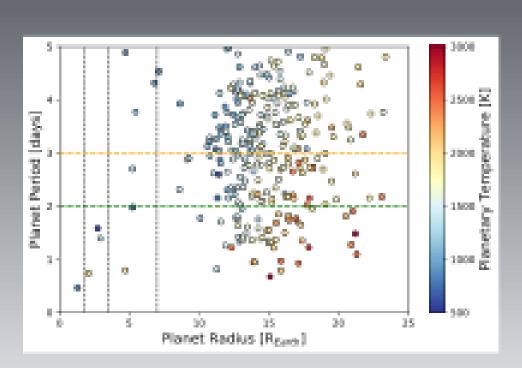


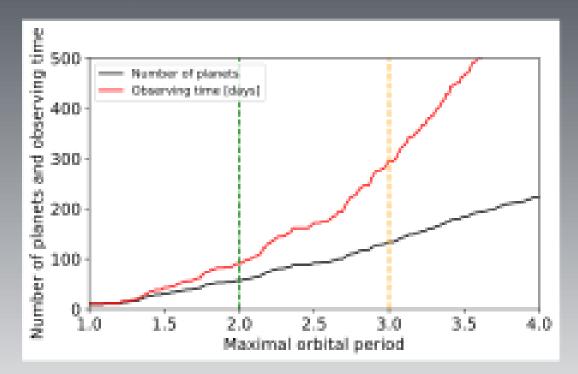
Skinner & Cho, 2022

Planets are 4D complex objects

S Ariel

Variability in space and time: More phase-curves....





Charnay, Dang, Cowan et al 2025

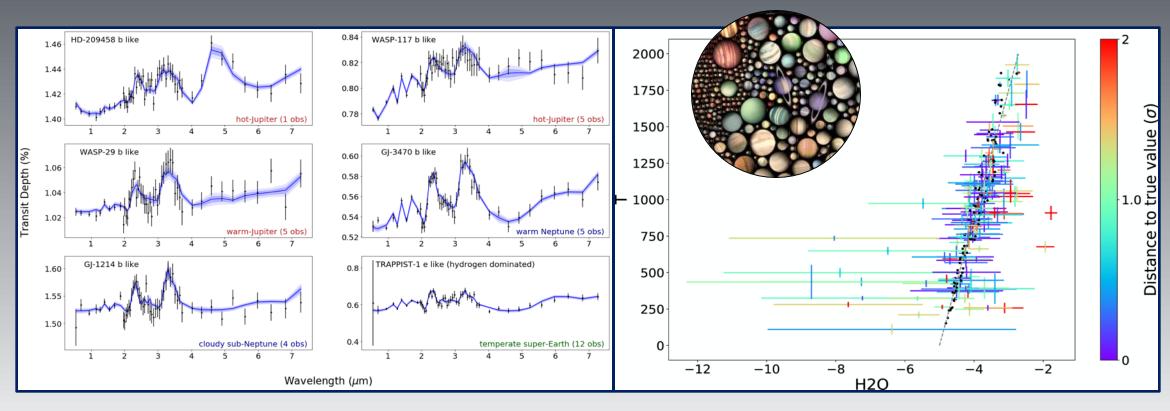
Conclusion

All three of our analyses lead to similar conclusions. By dedicating $\sim 20-25\%$ of the lifetime of the Ariel mission, we could observe $\sim 100-150$ exoplanets if imposing constraints on phase curve

Chemical survey



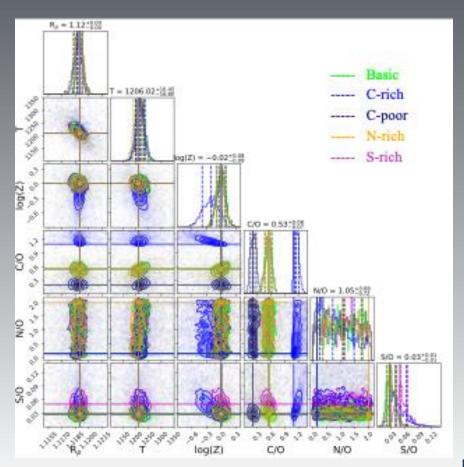
Searching for chemical and cloud transitions

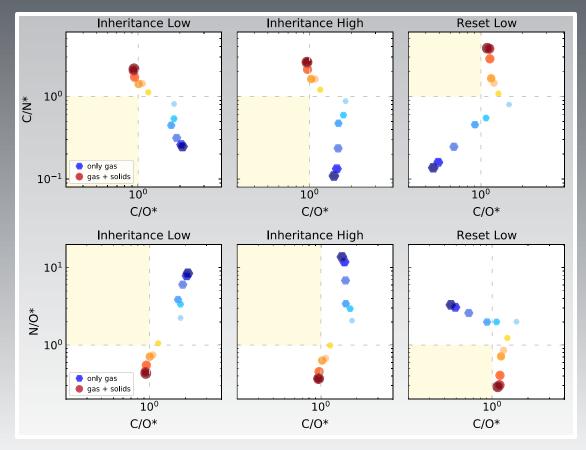


Changeat et al. 2020; see also Mugnai et al 2022, Bocchieri et al., 2024; Ma et al. in prep.

Link to planet formation

Ariel ability to detect elemental ratios in giant planets' atmospheres: beyond C/O





Fang et al. 2023 Sagan workshop – 2025

Pacetti et al 2022; Turrini et al. 2021

Focus on Ariel targets: stars







D L. Magrini¹, D C. Danielski^{2,3},

🗓 A. Brucalassi¹, 📵 M. Tsantaki¹, 📵

🗓 M. Van der Swaelmen¹. 🏻 📵 S. G. Soເ

Bruno¹¹ and 🔟 G. Casali^{12,13}

Contents

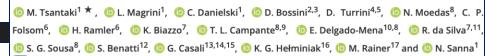
The homogeneous Ariel stellar catalogue currently includes the following parameters:

- Identification parameters
- Photometric properties
- Kinematics properties (galactic positions, velocities, parallaxes)
- Effective temperature, surface gravity and [Fe/H]
- vsin(i), vmicro
- Mass*, Radius, age
- Abundances of C, N, O for 181 stars

A&A 663, A161 (2022) Ariel stellar characterisation I. Homogeneous stellar parameters of 187 FGK pla validation of the method A&A, 697, A102 (2025) I. Chemical abundances of carbon, nitrogen, and oxygen for 181 planet-host FGK dwarf stars *.** © R. da Silva^{1,2}, C. Danielski^{3,4}, © E. Delgado Mena⁵, © L. Magrini³, D. Turrini⁶, K. Biazzo¹, M. Tsantaki³, © M. Rainer⁷, © K. G. Helminiak⁸, © S. Benatti⁹, V. Adibekyan⁵, © N. Sanna³, © S. Sousa⁵, © G. Casali 10.11,12 and © M. Van der Swaelmen³ A&A, 697, A102 (2025)

Ariel stellar characterisation

III. Fast rotators and new FGK stars in the Ariel mission candidate sample



Spectropolarimetric characterisation of exoplanet host stars in preparation of the *Ariel* mission

Magnetic environment of HD 63433

S. Bellotti^{1,2}, D. Evensberget¹, A. A. Vidotto¹, A. Lavail², T. Lüftinger³, G. A. J. Hussain³, J. Morin⁴, P. Petit², S. Boro Saikia⁵, C. Danielski⁶, and G. Micela⁷

Focus on Ariel targets: stars



Targets monitoring is being prioritised to maximise the science return of Ariel

From: Science Mission Office hubblereview@stsci.edu
Subiect: HST Cycle 32 Phase I Notification Snapshot Letter

Date: 8 July 2024 at 18:05

To: Dr. Sudeshna Boro Saikia sudeshna.boro.saikia@univie.ac.at

Cc: HST17794@stsci.edu, Giovanna Tinetti g.tinetti@ucl.ac.uk, Manuel Guedel manuel.guedel@univie.ac.at, Kristina Kislyakova kristina.kislyakova@univie.ac.at, Simon Schleich simon.schleich@univie.ac.at, Gwenael Van Looveren gwenael.van.looveren@univie.ac.at, Franz Kerschbaum franz.kerschbaum@univie.ac.at, Andrea Bocchieri andrea.bocchieri@uniroma1.it, Lorenzo Mugnai lorenzo.mugnai@uniroma1.it, Yamila Miguel ymiguel@strw.leidenuniv.nl, Aline Vidotto vidotto@strw.leidenuniv.nl, Jiri Zak jirizak1@seznam.cz, Donna Rodgers-Lee dlee@cp.dias.ie, Theresa Lueftinger theresa.rank-lueftinger@esa.int, Ignazio Pillitteri ignazio.pillitteri@inaf.it, Sarah Casewell slc25@leicester.ac.uk, Billy Edwards b.edwards@sron.nl, Krisztian Vida vidakris@konkoly.hu, Luca Fossati luca.fossati@oeaw.ac.at, Stefano Bellotti sbellotti@irap.omp.eu, Olivia Venot olivia.venot@lisa.ipsl.fr, Antonio Maggio antonio.maggio@inaf.it, Antonio Garcia Munoz antonio.garciamunoz@cea.fr, Carol Rodriguez crodriguez@stsci.edu

Sudeshna Boro Saikia University of Vienna

AUT

Jul 08, 2024

Dear Dr. Boro Saikia.

We are pleased to inform you that your Hubble Space Telescope Cycle 32 proposal

Title: FUV flux of nearby exoplanet host stars in the Ariel target list

ID: 17794

has been approved for Hubble Space Telescope Cycle 32 and Cycle 33 Snapshot observations, following detailed consideration by the Cycle 32 Peer Review Panels and final review by the STScl Director.

The allocations approved for your program in Phase I are:

37 Snapshot Targets in Cycle 32

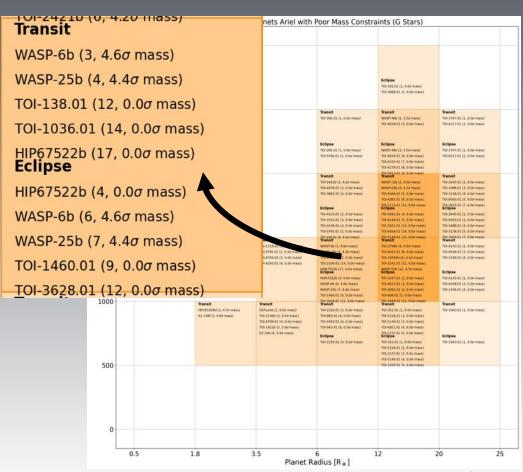


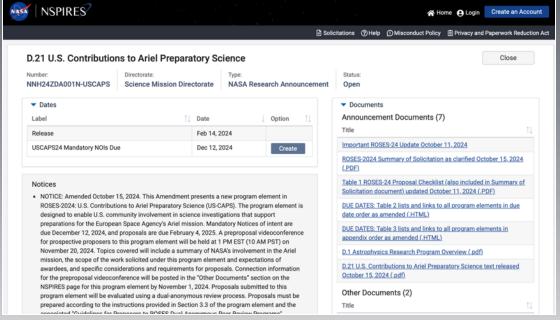


Focus on Ariel targets: masses



Targets monitoring is being prioritised to maximise the science return of Ariel







ExoClock: target ephemerides+ exo



2000+ participants from 70+ countries (77% amateurs). New paper submitted ©



ExoClock Project IV: A homogenous catalogue of 620 exoplanet ephemerides for the Ariel space mission

A. Kokori, ¹ A. Tsiaras, ² G. Pantelidou, ³ A. Jones, ^{4,5} A. Siakas, ⁶ B. Edwards, ^{7,1} G. Tinetti, ¹ P. Batsela, ² L. Bewersdorff, ^{4,8} A. Bocchieri, ⁶ R. A. Buckland, ^{9,10,5} A. R. Capildeo, ^{4,11} R. Casali, ⁴ S. R.-L. Futcher, ^{4,12,13,14} D. Gakis, ^{15,16} G. Grivas, ³ A. Iliadou, ³ Y. Jongen, ^{4,17} G. Lekkas, ¹⁸ F. Libotte, ^{4,19,20,21} P. Matassa, ⁴ V. Michalaki, ⁶ L. V. Mugnai, ^{22,23,1} A. Nastasi, ^{24,25} N. I. Paschalis, ⁴ C. Pereira, ^{4,26} A. Popowicz, ²⁷ E. Poultourtzidis, ³ D. Rees, ²⁸ C. Sidiropoulos, ²⁹ F. Walter, ^{4,30,31} A. Wünsche, ³² M. Á. Álava-Amat, ^{4,33} M. V. Crow, ^{4,5,10} S. Dawes, ^{4,5,10} C. Falco, ^{34,35} A. O. Kovacs, ^{4,36} C. Lopresti, ^{4,37} A. Marchini, ³⁸ B. E. Martin, ^{4,36} R. Naves, ^{4,19,39} M. Raetz, ^{4,40,41} R. Roth, ⁴² D. Stouraitis, ⁴ J.-P. Vignes, ⁴ K. Agabi, ⁶ N. A-Thano, ^{43,44} L. ABE, ⁶ R. Abraham, ^{4,45,46}



Ariel Databases

A new paper and Database about spectroscopic, chemical and cloud data for Ariel

RAS Techniques and Instruments



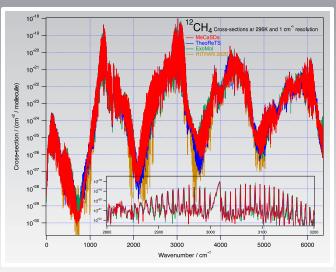
RASTAI 3, 636–690 (2024) Advance Access publication 2024 September 19

https://doi.org/10.1093/rasti/rzae039

Data availability and requirements relevant for the *Ariel* space mission and other exoplanet atmosphere applications

Katy L. Chubb , 1,2 Séverine Robert, Clara Sousa-Silva, 4,5 Sergei N. Yurchenko , 6 Nicole F. Allard , Vincent Boudon, Jeanna Buldyreva, Benjamin Bultel, Athena Coustenis, Aleksandra Foltynowicz, Iouli E. Gordon , 13 Robert J. Hargreaves , 13 Christiane Helling, 14,15 Christian Hill, Helgi Rafn Hrodmarsson, Tijs Karman, Helena Lecoq-Molinos, 14,15,19 Alessandra Migliorini , 20 Michaël Rey, I Cyril Richard, Ibrahim Sadiek, Frédéric Schmidt, Andrei Sokolov, Stefania Stefani, Jonathan Tennyson , Olivia Venot , Tsam O. M. Wright, Rosa Arenales-Lope, Jonathan K. Barstow , Andrea Bocchieri, Nathalie Carrasco, Dwaipayan Dubey, Oleg Egorov, Antonio García Muñoz, Ehsan (Sam) Gharib-Nezhad, Leonardos Gkouvelis, Fabian Grübel, Patrick Gerard Joseph Irwin , Antonín Knížek, David A. Lewis, Matt G. Lodge , Sushuang Ma, Zita Martins, Karan Molaverdikhani , Sushuang Ma, Gita Martins, Giovanna Rinaldi, Giuseppe Morello, Andrei Nikitin, Emilie Panek, Miriam Rengel, Giovanna Rinaldi, Giuseppe Morello, Giovanna Tinetti, Tim A. Van Kempen, Jingxuan Yang , and Tiziano Zingales, Giovanna Tinetti, Tim A. Van Kempen, Jingxuan Yang , and Tiziano Zingales, Andrei Nikitin, Tim A. Van Kempen, Jingxuan Yang , so Jingxuan Zingales, Jingx

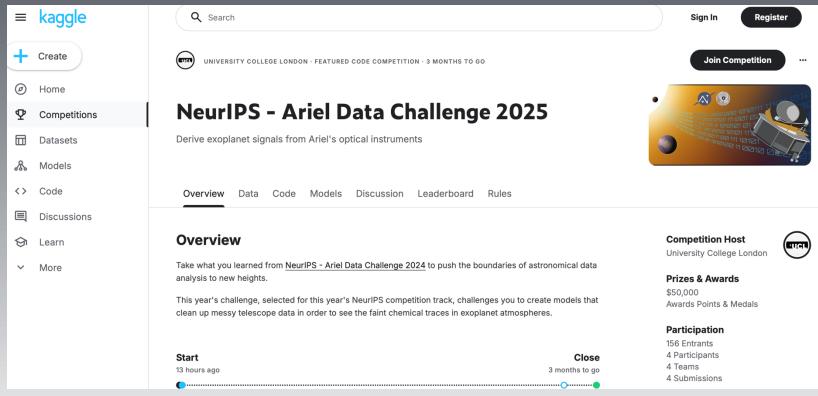




Ariel Data Challenges 2025



ADC 2025: just started!!!!!!





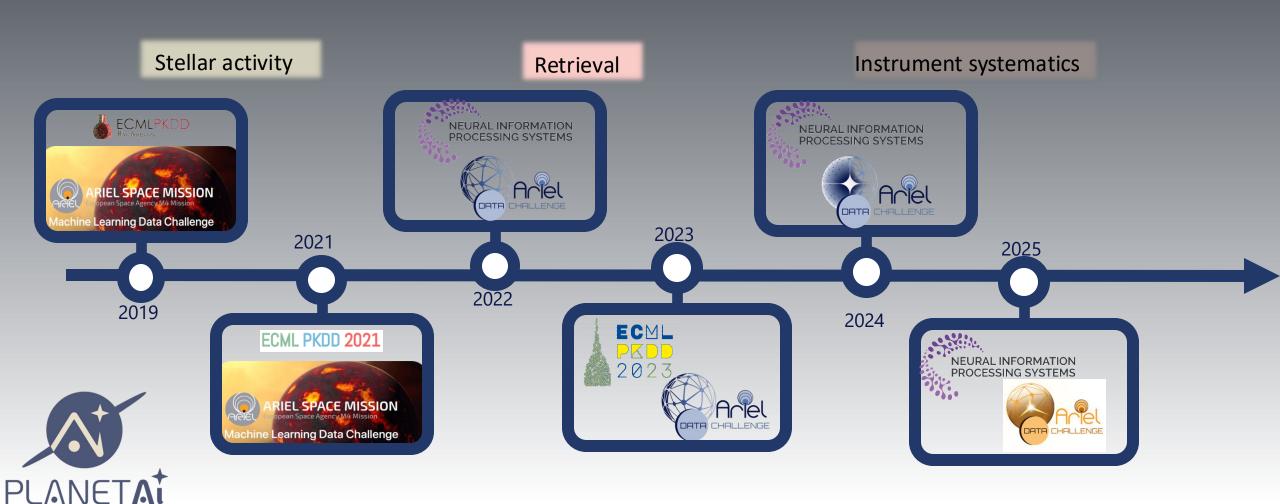


A mission is more than flying hardware.... There is XAI @

6 years of Ariel Data Challenges



A huge global success. ADCs yearly planned to support ground segment activities



Ariel Data Challenge 2024





77 Countries

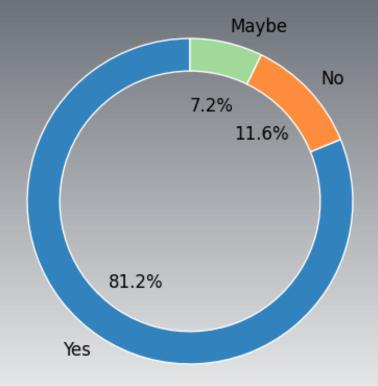


6000+ Entrants



23,000+ Submissions

Is your work/study related to ML/AI?





Ariel Data Challenges



All the data and codes are open sourced to drive progress

JOURNAL ARTICLE

ESA-Ariel Data Challenge NeurIPS 2022: introduction to exo-atmospheric studies and presentation of the Atmospheric Big Challenge (ABC) Database 3

Quentin Changeat ™, Kai Hou Yip Author Notes

RAS Techniques and Instruments, Volume 2, Issue 1, January 2023, Pages 45-61,

Pub

Proceedings of Machine Learning Research 220:1-17, 2023

NeurIPS 2022 Competition Track

Lessons Learned from Ariel Data Challenge 2022 Inferring Physical Properties of Exoplanets From **Next-Generation Telescopes**

Reproducing Bayesian Posterior Distributions for Exoplanet Atmospheric Parameter Retrievals with a Machine Learning Surrogate Model

Eyup B. $Unlu^{1[0000-0002-6683-6463]}$, Roy T. Forestano $^{1[0000-0002-0355-2076]}$, Konstantin T. Matchev¹[0000-0003-4182-9096], and Katia $Matcheva^{1[0000-0003-3074-998X]}$

Institute for Fundamental Theory, Physics Department, University of Florida. Gainesville FL 32653, USA

KAI.HOU.YIP@UCL.AC.UP IN.CHANGEAT.18@UCL.AC.UK NGO.WALDMANN@UCL.AC.UK

EYUP.UNLU@UFL.EDU

Taru 2024 - Arie

OPEN ACCESS

Searching for Novel Chemistry in Exoplanetary Atmospheres Using Machine Learning for Anomaly Detection

Roy T. Forestano^{2,1}, Konstantin T. Matchev^{2,1}, Katia Matcheva^{2,1}, and Eyup B. Unlu^{2,1} Published 2023 November 16 • © 2023. The Author(s). Published by the American Astronomical Society.

> Lessons Learned from the 1st ARIEL Machine Learning Challenge: Correcting Transiting Exoplanet Light Curves for Stellar Spots

NIKOLAOS NIKOLAOU D, INGO P. WALDMANN D, ANGELOS TSIARAS D, MARIO MORVAN D, BILLY EDWARDS D, KAI HOU YIP 0, GIOVANNA TINETTI 0, SUBHAJIT SARKAR, JAMES M. DAWSON, VADIM BORISOV, GJERGJI KASNECI, MATEL PETROMÓ 4 TOMA Z STEDIČNIK 4 TAREK AT LIDATOT 5,6 RACHEL LOUISE BALEV D 6 MICHAEL GRANITZER D,7 MIRKO BUNSE , 10 AND

Simulation-based Inference for Exoplanet Atmospheric Retrieval: Insights from winning the Ariel Data Challenge 2023 using Normalizing Flows

G Predicting Exoplanetary Features with a Residual Model for Uniform and Gaussian Distributions

Andrew Sweet

Assemi Group, Inc., Fresno CA 93704, USA asweet@assemigroup.com

Ariel Data Policy

A very open approach: foundation of good rigorous science and reproducibility

Science Demonstration Phase

Data will be released immediately after processing and quality control

Nominal Science Operations Phase

- Tier 1 data public immediately after quality control is completed;
- Tier 2, 3 data public 6 months after quality control is completed;
- Tier 4 data public 1 year after quality control is completed.

Complementary Science data

- 5%-10% time available for other science, allocated through ESA calls
- Proprietary to the proposers for 6 months

Yes! We're OPEN

Strong commitment to open-source software, Explainable AI solutions





@arieltelescope

European Space Agency M4 Mission













Latest News & Updates



READ - Ariel undergoes the UK's first space acoustic tests! STORY HERE!



WATCH Ariel feature on BBC Click's Satellite Testing Centre segment HERE!

Consortium & Agency Sites







Ariel Open Conference 2026



@ ESA ECSAT, 17-19 March 2026. Ariel special issue to be published by RASTI











































Mauve ready for launch



Blue Skies Space announces Mauve launch scheduled for



Blue Skies Space Roadmap





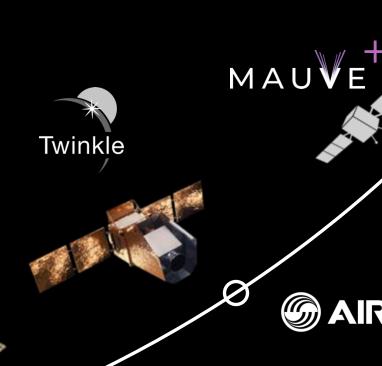
Built by industry experts



High heritage approach



Delivered in under three years





RadioLuna



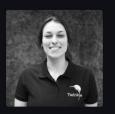


Meet the Team





Dr Arianna SabaScience
Performance Analyst



Sarah Harvey Junior Software Engineer



Sharafina Razin Junior Science Programme Manager



James McLaren Software Consultant



Tailong Zhang Intern



Gabriele Galletta Intern



Prof Giorgio Savini Science and Instrumentation



Dr Dan Brown Adviser



Dr Fabio Favata Senior Adviser



Prof Jonathan Tennyson Chair



Dr Marcell Tessenyi Chief Executive Officer



Prof Giovanna Tinetti Chief Scientist



Phillip Windred Chief Operating Officer



Richard Archer Strategic Partnerships



Benjamin Wilcock Senior Science Programme Manager



lan Stotesbury Lead Systems Engineer



Rachel Grant Senior Software Engineer



Lawrence Bradley Software Engineer



Dr Fatemeh Zahra MajidiMauve Project
Scientist

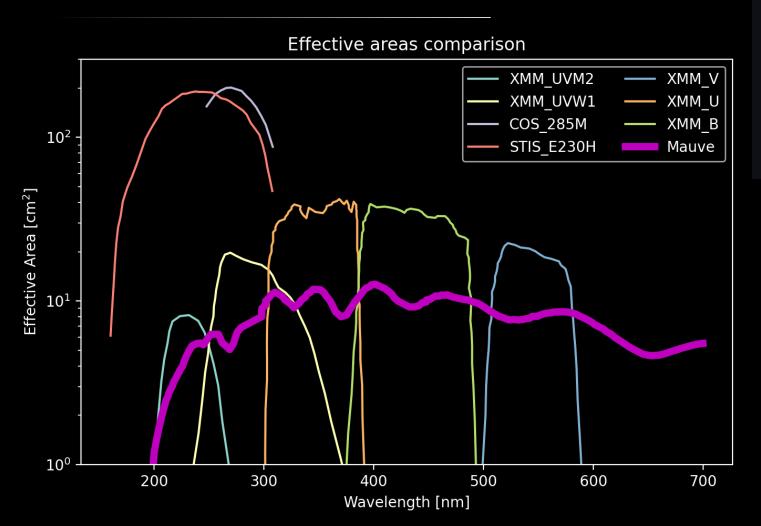


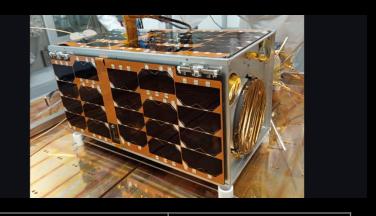
Yoga Barrathwaj Raman Mohan Marketing & Business Operations



Dr Parul Janagal Science Outreach

Comparison of MAUVE with known facilities





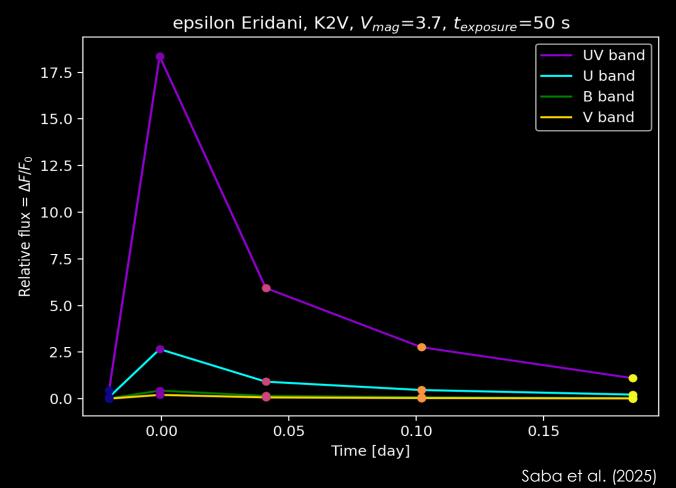
Wavelength range	200 - 700 nm (NUV + Visible)
Resolution	R = 20 - 65
Telescope	13 cm Cassegrain
Spectrometer	2-mirror grating spectrograph with CMOS linear array detector
Launch	October this year
Orbit	LEO polar – 500 km
Field of View	~105'' full cone

Epsilon Eridani, K2V, V=3.7



Light curves showing the relative flux increase over the course of the flare event.

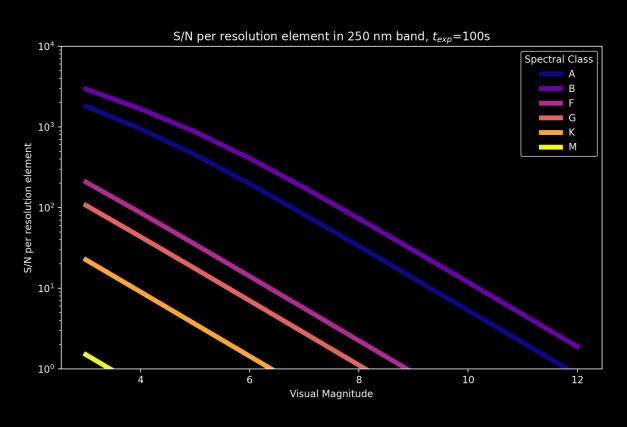
Different wavelength bands selected, highest contrast in the UV.

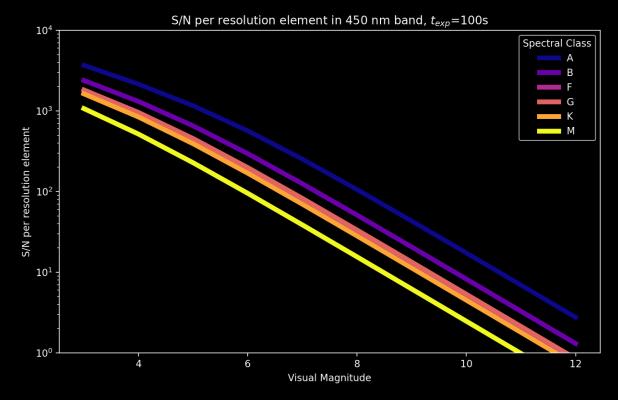




Mauve sensitivity

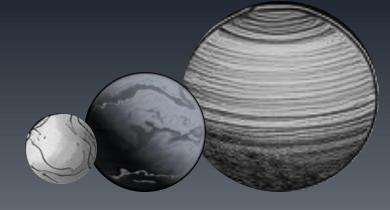
General performance capabilities of the observatory





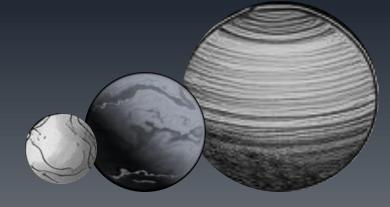
Saba et al. (2025)

Conclusions



- Planets are ubiquitous in our galaxy, they are very diverse
- We have a host of questions to answer in the next decade:
 - ? How do planets form? How do planets evolve?
 - ? How diverse are exoplanets chemically?
 - ? Does chemical diversity correlate with other parameters?
- Population studies of exoplanets are key to unlocking their statistical properties.

Conclusions

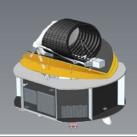


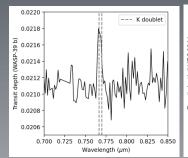
- An increasing number of exoplanet atmospheres are/have been observed from space and the ground.
- With Webb, Roman and Ariel, population-based studies of exoplanet atmospheres and their stellar hosts, offer an unparalleled opportunity to uncover chemical trends and how they correlate with stellar/planetary/orbital parameters.
- We advocate an ecosystem of small and moderate size facilities from space and the ground to complement largest facilities
- This information is key to guide our understanding of formation and evolution mechanisms of the planets in our galaxy.

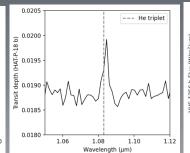
Synergy Ariel-Webb

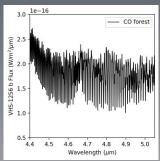


Complementarity and synergy between Ariel and Webb will be transformational



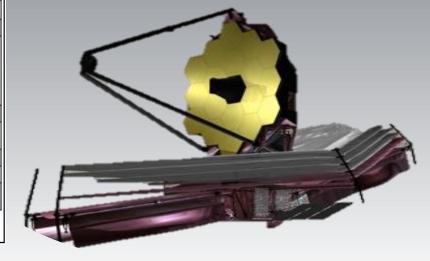






0.033 Ariel Ariel **JWST** PRISM 0.032 Diameter 0.9 m 6.5 m 0.031 NIRISS: 0.8 - 2.9 0.5 - 7.8PRISM: 0.6 - 5.3 Grisms: 1 - 5.2 NIRCam: 2.5 - 5 MIRI-LRS: 5 - 12 들 0.029 MIRI-MRS: 5 - 28 Transit I 15 to 100 Up to 3000 General 0.027 observatory **Nplanets** ~1000 100+ (primary) Launch 2029 2021 0.025 4 years 5 years (primary) Fuel for 20 0.024 Wavelength (µm)

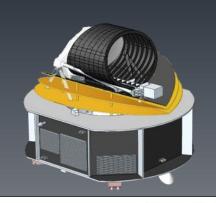




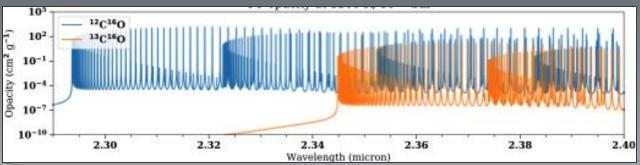
Changeat, Lagage et al. in prep

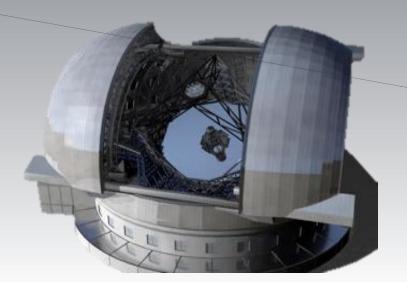
Synergies ground

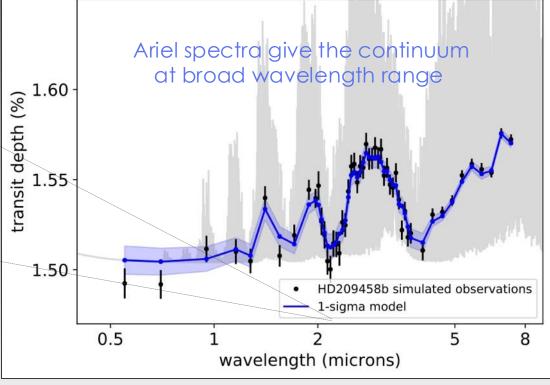
HIGHLY COMPLEMENTARY TO LARGE, GROUND-BASED FACILITIES









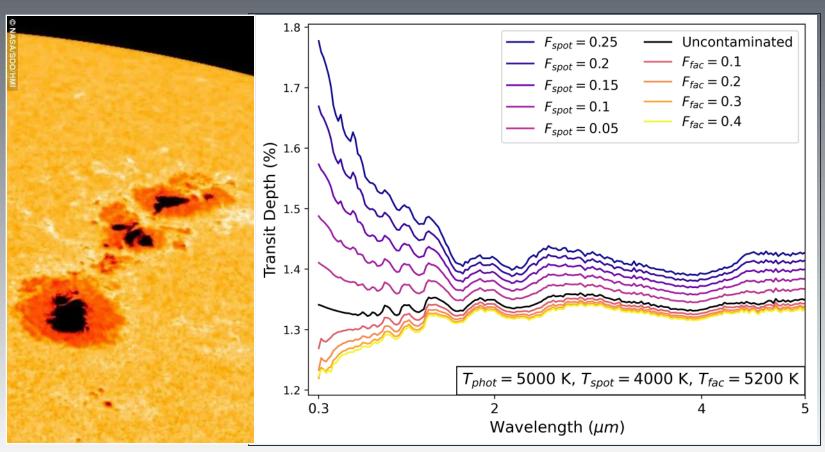


Ariel ELT synergy WG

Stellar contamination



Ariel Tier 3 observations will help understanding impact of contamination through time



Thompson et al. 2025

Ariel 4-Tier approach

INDIVIDUAL PLANETS & POPULATION ANALYSIS

- What fraction of planets have clouds?
- Have small planets still retained H/He?
- Colour-colour diagrams
- Refinement of orbital/planet parameters in IR

SURVEY - TIER 1

DEEP SURVEY — TIER 2

BENCHMARK - TIER 3

~ 50-100

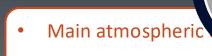


- Phase-curves
- Tailored observations



~ 1000 PLANETS

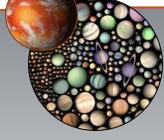
- Atmospheric circulation
- Spatial & temporal variability





- Thermal structure
- Cloud characterization



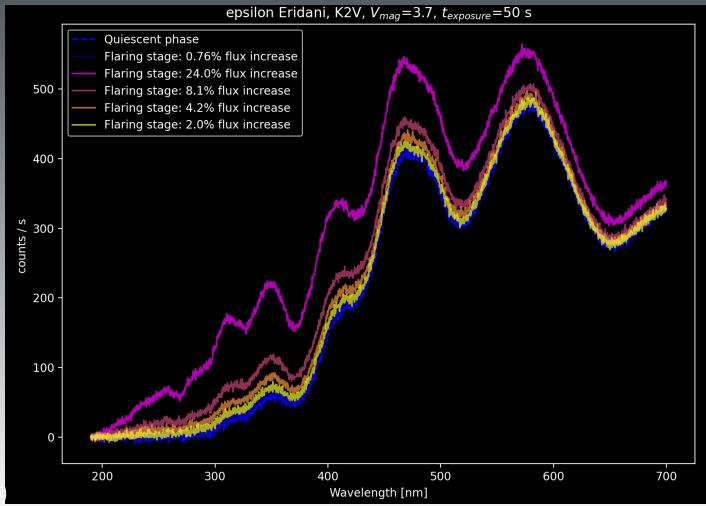


53

Epsilon Eridani, K2V, V=3.7



- Simulated Mauve spectra at different stages of the flare event.
- Quiescence, flare peak and decay stages are clearly distinguishable from one another.
- High contrast across the wavelength range accessible to Mauve, and specifically in the UV



aba et al. (2025)