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EXOPLANET DETECTION LIMITS

WOW!
OCCURRENCE
RATES!

NEW
METHOD!

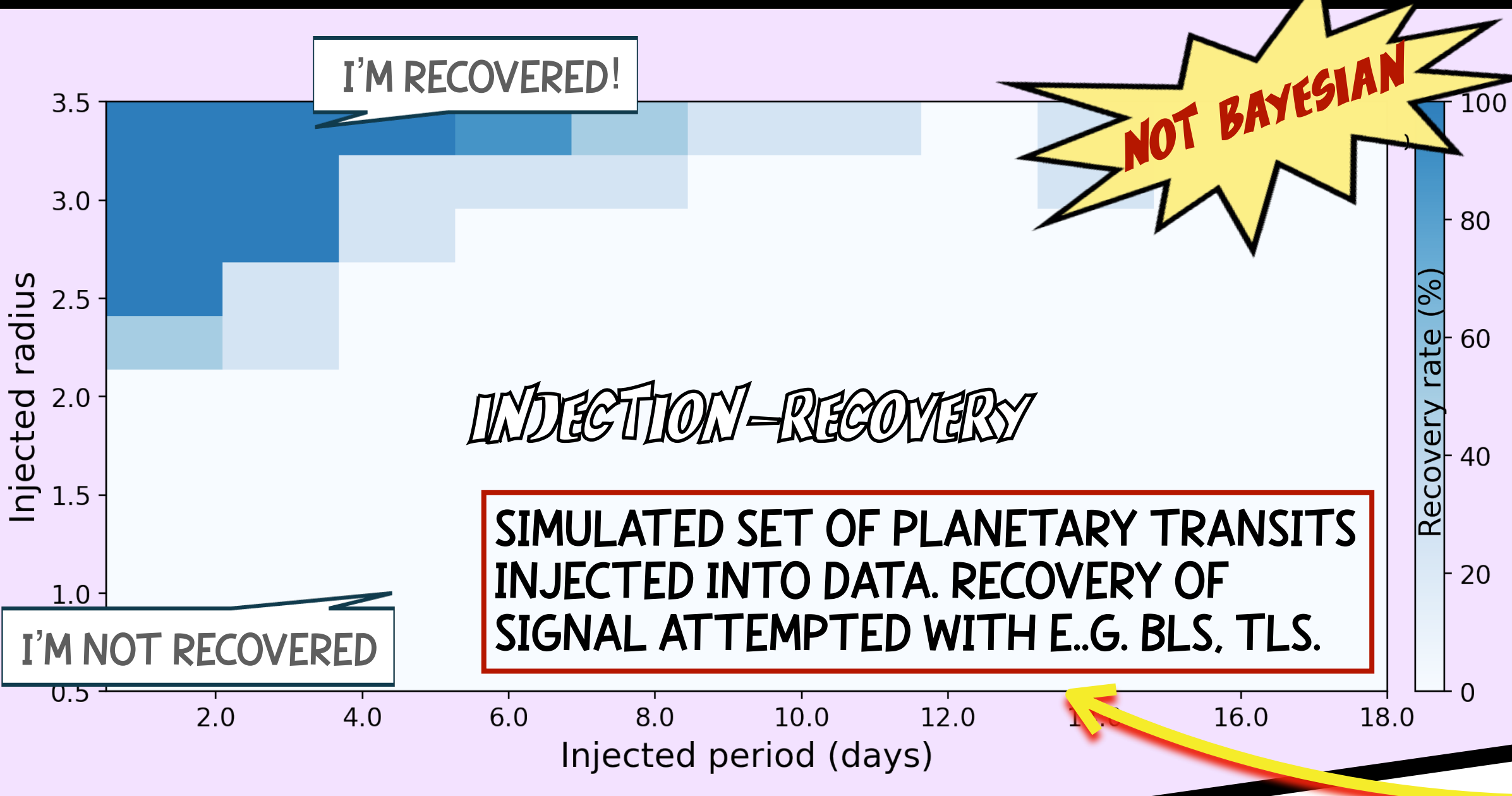
WE WANT TO MEASURE
THE OCCURRENCE RATES
OF EXOPLANETS AROUND
LATE-TYPE M DWARFS

WE WANT TO FIND WHERE OUR DETECTION LIMIT LIES FOR A GIVEN PHOTOMETRIC DATASET

THIS WILL HELP INFORM OCCURRENCE RATES BY ASSESSING OUR COMPLETENESS USING
DETECTION LIMITS

WE ARE ASKING THE QUESTION
“WHAT IS COMPATIBLE WITH THE DATA?”

WHAT REMAINS FORMALLY UNDETECTED?
THIS IS NOT INJECTION RECOVERY



PROBLEM 1
ONLY TESTS FOR
A VERY SPECIFIC
SET OF PLANET
PARAMETERS.

PROBLEM 2
EACH RECOVERY IS
COMPUTATIONALLY
INTENSE: COARSE
GRIDS USED

COMPUTATION TIME ~ 5 DAYS

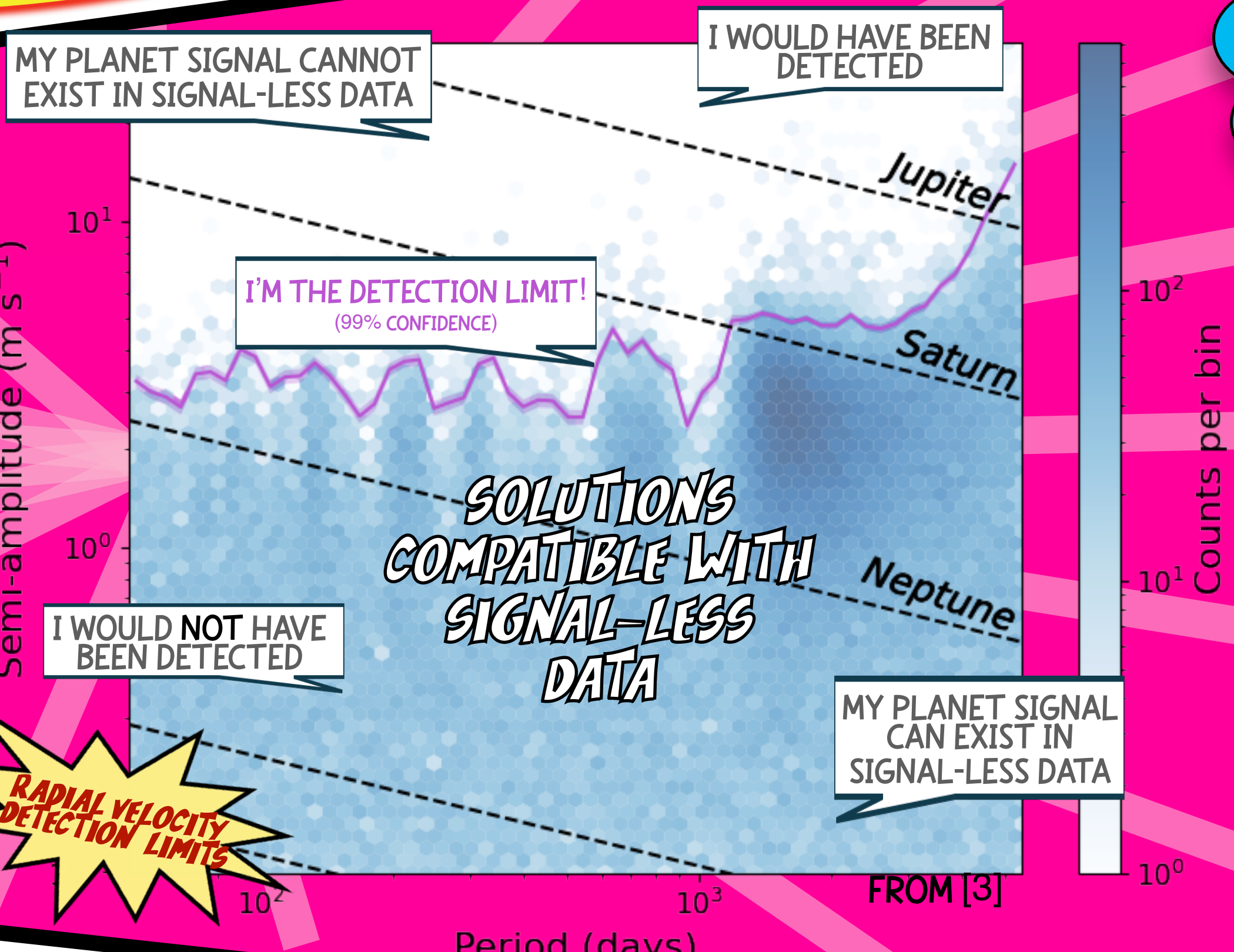
THIS HAS BEEN DONE
FOR RADIAL VELOCITY
DATA
USING KIMA [2]

KIMA IS A PYTHON
PACKAGE FOR THE
DETECTION AND
CHARACTERISATION
OF EXOPLANET
SIGNALS IN IN RADIAL
VELOCITY DATA

THE APPROACH IS MORE COMPLEX FOR APPLICATIONS TO
TRANSITS.

PROBLEM 3 UNLIKE RV DATA, THERE IS NOT A SIGNAL
ACROSS THE ENTIRE ORBIT OF THE PLANET.

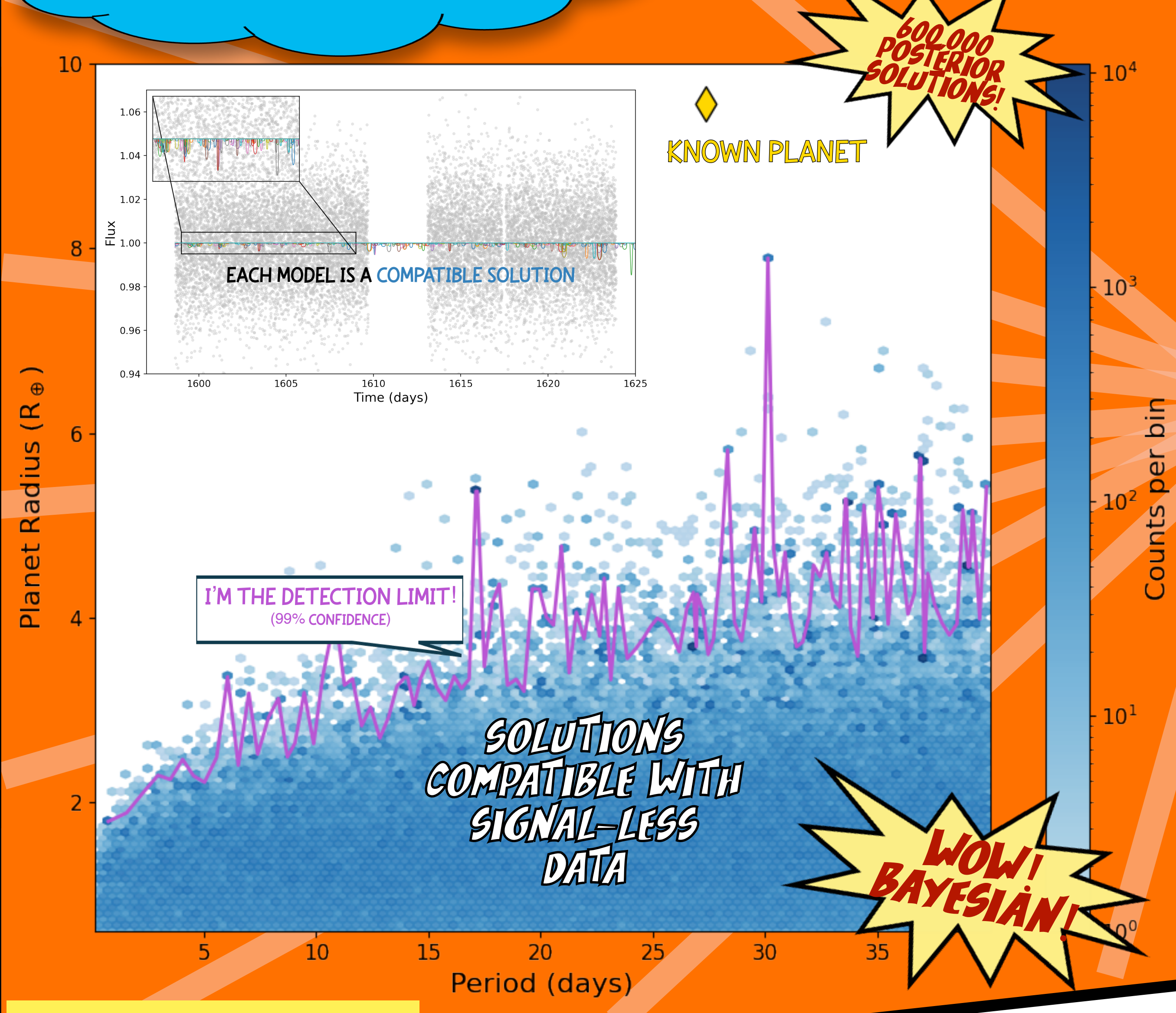
PROBLEM 4 HOLES IN THE DATASET HAVE BIG
IMPACTS ON EXOPLANET DETECTIONS.



FOR A GIVEN PERIOD + SEMI-AMPLITUDE UNDER THE PURPLE LINE, THE RV SIGNAL PRODUCED COULD EXIST IN THE DATA.

IT'S “HIDDEN” IN THE NOISE AND REMAINS FORMALLY UNDETECTED. EVERYTHING ABOVE THE PURPLE LINE WOULD HAVE BEEN DETECTED.

HOW ARE WE DOING IT
FOR PHOTOMETRY?



COMPUTATION TIME ~ < 5 HOURS

METHOD

BATMAN [4] + DYNESTY [5] PYTHON PACKAGES.

PRIORS ON MODEL PARAMETERS ARE EITHER LARGE + UNIFORM ($t_0, P, R_p/R_*, \omega$) OR INFORMED FROM THE CURRENT TRANSITING EXOPLANET POPULATION (e, i)

TENS OF THOUSANDS OF MODELS ARE PROPOSED. MODELS ARE EITHER REJECTED IF NOT COMPATIBLE WITH SIGNAL-LESS DATA, OR ACCEPTED AS A COMPATIBLE SOLUTION.

ADVANTAGES
COMPARED TO INJECTION
RECOVERY

- 1 BAYESIAN
- 2 NON-DISCRETE SOLUTION GRID
- 3 4 MORE PARAMETERS DRAWN FROM DISTRIBUTIONS
- 4 100s – 1000s MORE SOLUTIONS
- 5 ~20 × FASTER

SCOTT TRIAUD &
DAVIES, IN PREP

COMING
2025 / 26!

[1] DÍVORA-PAJARES, M. AND POZUELOS, F. J., “MATRIX: MULTI-PHASE TRANSITS RECOVERY FROM INJECTED EXOPLANETS”

[2] FARIA, J. P., SANTOS, N. C., FIGUEIRA, P., AND BREWER, B. J., “KIMA: EXOPLANET DETECTION IN RADIAL

[3] BAYCROFT, T. ET AL., “BEBOP VII: SOPHIE DISCOVERY OF BEBOP-3B, A CIRCUMINARY GIANT PLANET ON AN ECCENTRIC ORBIT”

[4] KREIDBERG, L., “BATMAN: BASIC TRANSIT MODEL CALCULATION IN PYTHON”

[5] SPEAGLE, J. S., “DYNESTY: A DYNAMIC NESTED SAMPLING PACKAGE FOR ESTIMATING



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