# Discovery space and science with the PLACID stellar coronagraph

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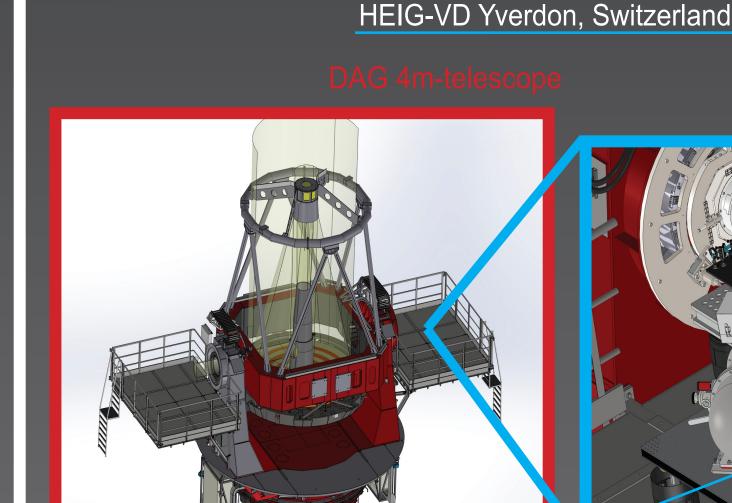
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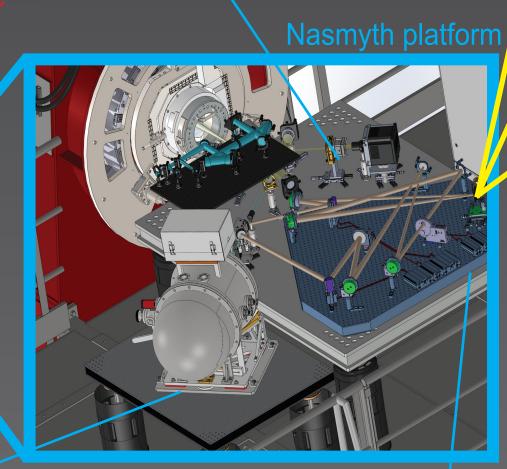
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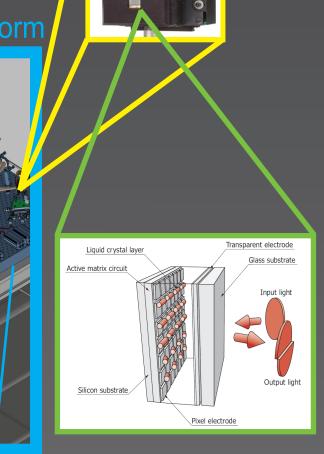


### Introduction

- PLACID (Programmable Liquid-crystal Active Coronagraph Imager for the DAG telescope) is a coronagraphic instrument, providing adaptive high-contrast imaging capbilities from H to Ks band
- PLACID uses a pixelated Spatial Light Modulator (SLM) to generate coronagraphic focal-plane masks (FPMs) for the first time on a telescope (4 m, DAG observatory, Erzurum, Turkey)
- Instrument delivered in March '24, first light expected by end of 2024
- Remote reconfiguration on-demand to adapt to e.g. observing conditions, multiple star coronagraphy, correcting aberrations, segmented primary mirrors (ELTs, HWO, ...)







Doğu Anadolu Gözlemevi

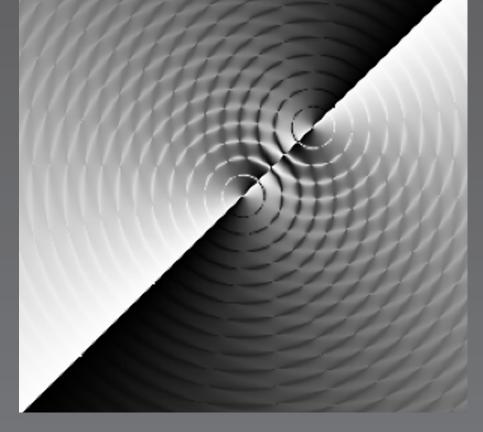
(DAG, Eastern Anatolia

Spatial Light Modulator (SLM)

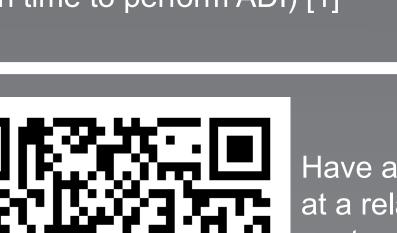
Observatory)

DIRAC NIR camera: developed by Macquarie University Sydney, Australia

PLACID: active (SLM-based) coronagraph instrument jointly developed at University of Bern and HEIG-VD Yverdon, Switzerland



Vortex FPM (n=2) programmed for a binary star (can be rotated in time to perform ADI) [1]

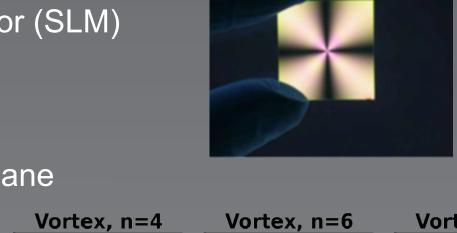


Have a look at a related poster on simulations of an SLMbased coronagraph!

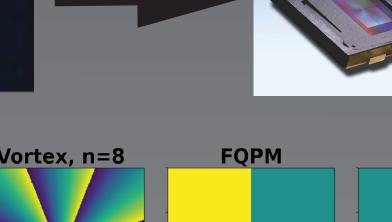
## The SLM as an FPM

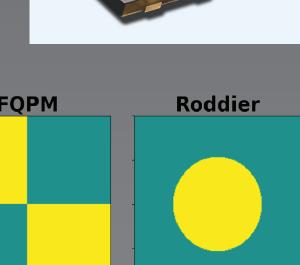
- Liquid Crystal on Silicon (LCOS, Meadowlark) Spatial Light Modulator (SLM) used to program adaptive FPM
- Micron sized pixels provide sampling of (>10 px per λ/D units)
   ~10 μm pixels provides exquisite spatial sampling in the focal plane
- Simulations: sampling of 10 px per λ/D sufficient in most cases
- Scalar phase shift applied:
- Chromatic, requires linearly polarized light

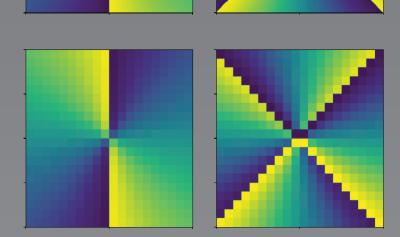
Figure: commonly used FPMs with left column: 10 px per λ/D and right column: 100 px per λ/D



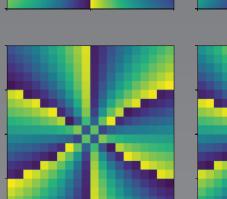
TROIA: AO-system developed by

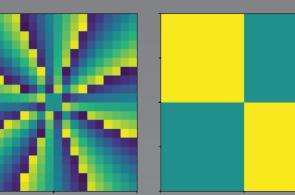


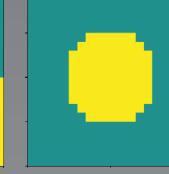




Vortex, n=2





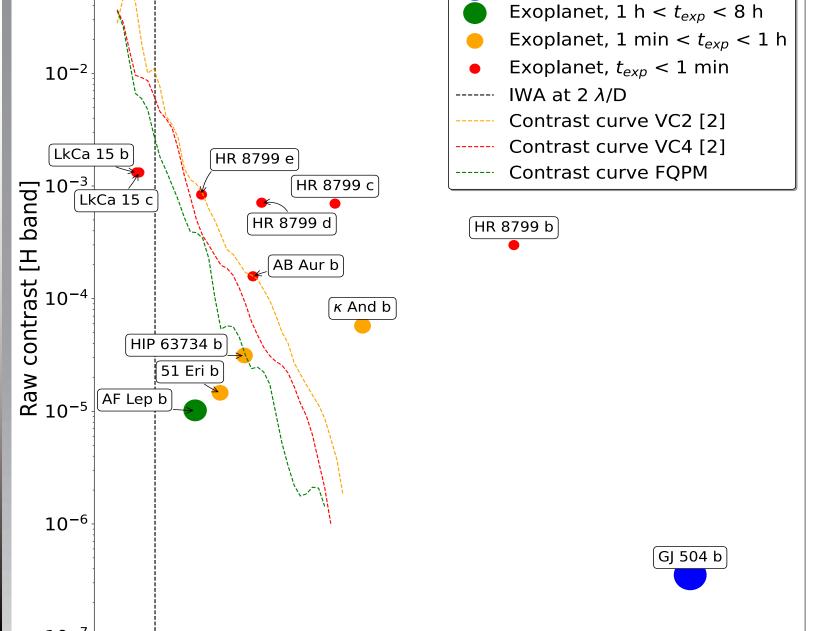


## PLACID targets and discovery space

2.5

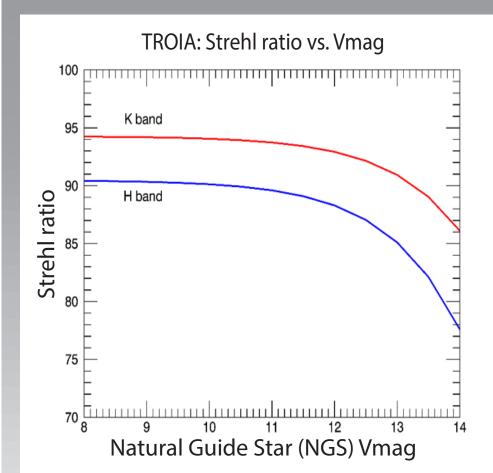
Known Directly Imaged Exoplanets/Candidates

Exoplanet,  $t_{exp} \ge 8h$ Exoplanet,  $1 h < t_{exp} < 8 h$ Exoplanet,  $1 \min < t_{exp} < 1 h$ Exoplanet,  $t_{exp} < 1 \min$ 

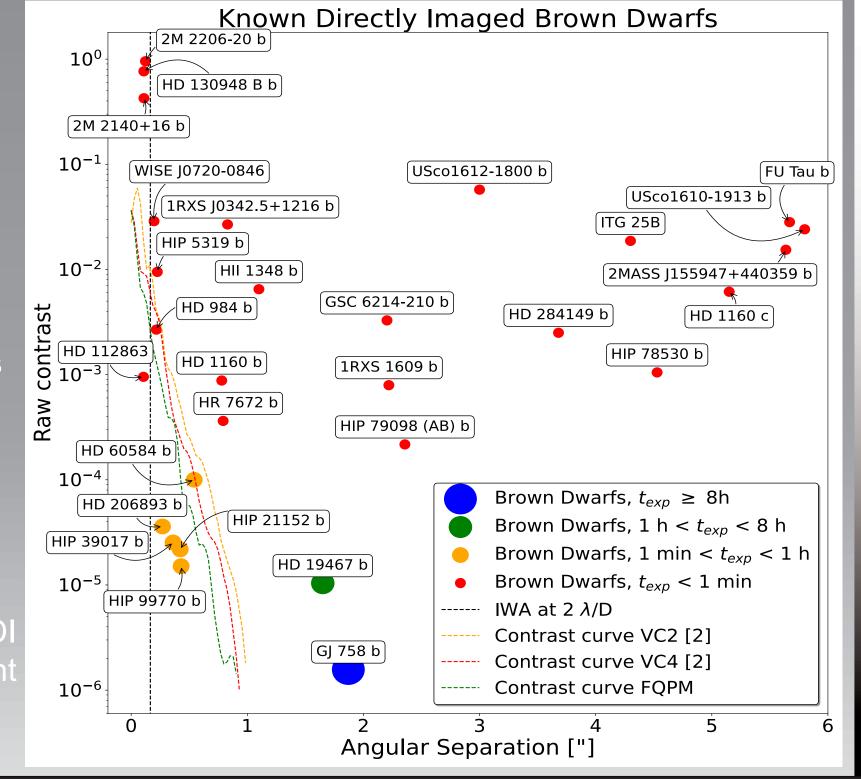


Angular Separation ["]

- Known exoplanets / candidates, Brown Dwarfs, circumstellar disks, binaries/triples
- Gaia, TESS, PLATO direct-imaging follow-ups in the North
- PLACID observational constraints:
  - ∘ Site: DEC: ≥ -24°
  - ∘ TROIA AO guide star: V ≤ 13 mag
  - On-sky FOV: 16" x 9.6"



- Plots: lab contrast curves
- Exposure time t<sub>exp</sub>
- required for SNR = 5
   without coronagraph
- no speckles
- Post-processing, ADI, CDI
   → factor 10 improvement



#### Outlook

- Obtaining on-sky contrast curves for PLACID
- Upgrading Exposure Time Calculator
- Setting up data reduction pipeline

## References

[1] Jonas G. Kühn et al. SLM-based digital adaptive coronagraphy: current status and capabilities. Advances in Optical and Mechanical Technologies for Telescopes and Instrumentation III, volume 10706 of SPIE Conference Series, page 107062N, July 2018. doi: 10.1117/12.2312554.

[2] Jonas G. Kühn et al. SLM-based Active Focal-Plane Coronagraphy: Status and future on-sky prospects. Advances in Optical and Mechanical Technologies for Telescopes and Instrumentation IV, 114511S, Proceedings of SPIE 11451, February 2021. doi:10.1117/12.2562579









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**Swiss Confederation** 

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