

# Detecting False Positives with Oxygen: A Feasibility Study



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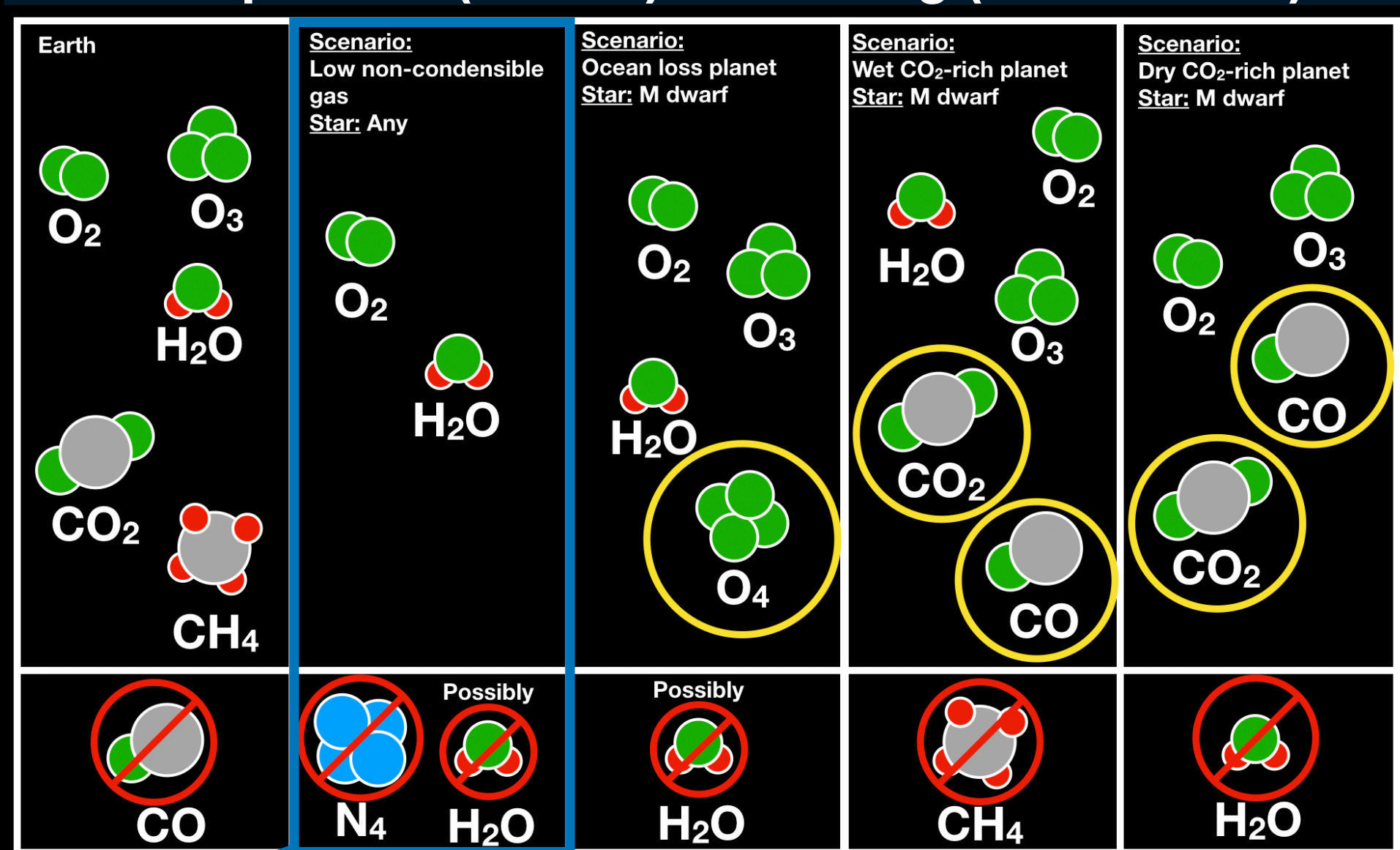


We are exploring how to use ground-based high-resolution spectroscopy to distinguish between **abiotic O<sub>2</sub>** formed from H<sub>2</sub>O photolysis in the upper atmosphere, and well-mixed **biological O<sub>2</sub>**.

## O<sub>2</sub> as a Biosignature

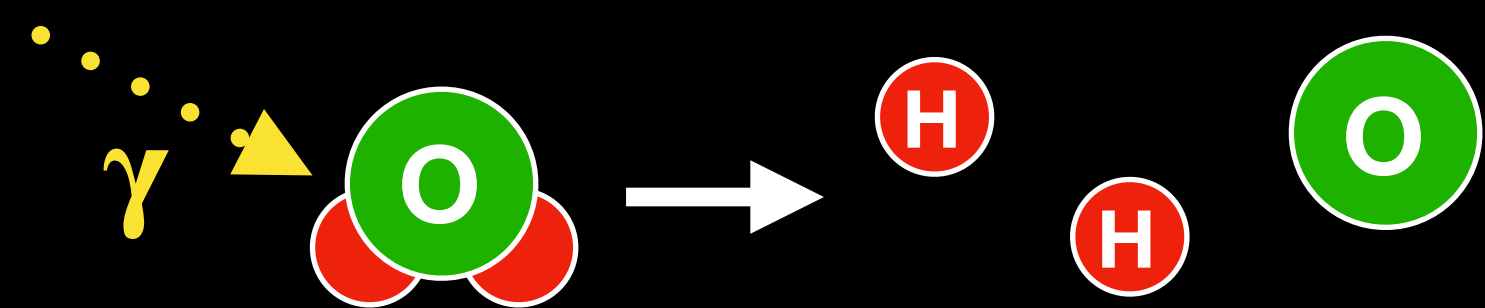
- Oxygen is a strong biosignature
- However, there are several abiotic mechanisms which can produce O<sub>2</sub>
- The abiotic production of O<sub>2</sub> via H<sub>2</sub>O photolysis may be particularly difficult to distinguish from biological O<sub>2</sub>
- We present the prospects for using ground-based high-resolution spectroscopy to reconcile this **false positive** (below) mechanism

Many false positives for biological O<sub>2</sub> are known, but most reveal their presence via additional molecules that are present (circled) or missing (crossed out).



Adapted from Meadows et al. 2018b

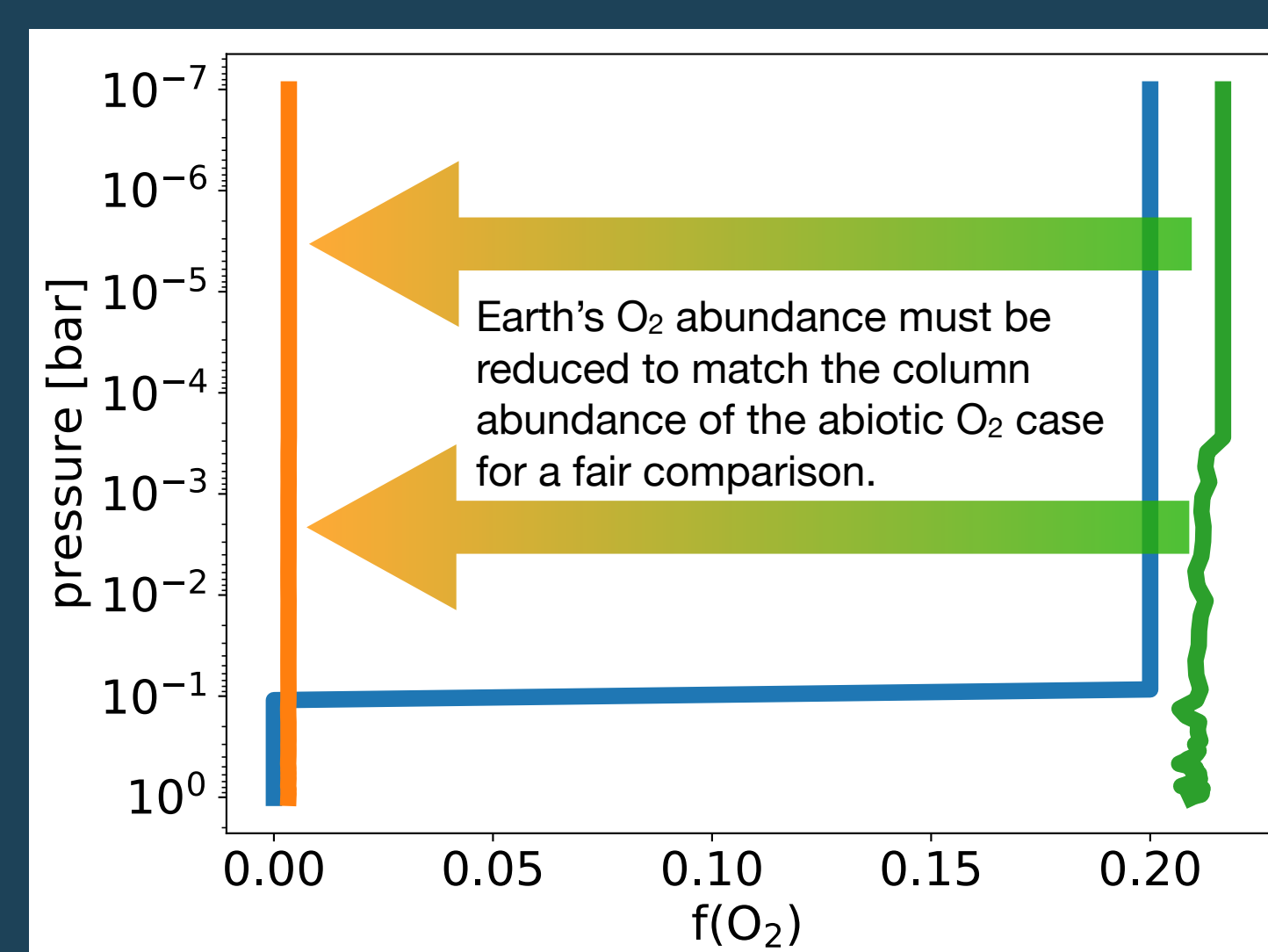
## Abiotic O<sub>2</sub> Production



- Lowering the non-condensable gas (e.g. N<sub>2</sub>) inventory allows more H<sub>2</sub>O into the upper atmosphere (Wordsworth and Pierrehumbert 2014)
- H<sub>2</sub>O is photolyzed and H escapes to space
- O<sub>2</sub> builds up in the upper atmosphere**
- Possible for planets around all stellar types

## O<sub>2</sub> Profiles

To compare the spectrum of a photochemically produced oxygen profile to the spectrum of a well-mixed biologically produced oxygen profile, we must ensure that the column abundances of these two profiles are equal. This requires a well-mixed profile with  $f(\text{O}_2) = 0.33\%$ .



## Test Cases:

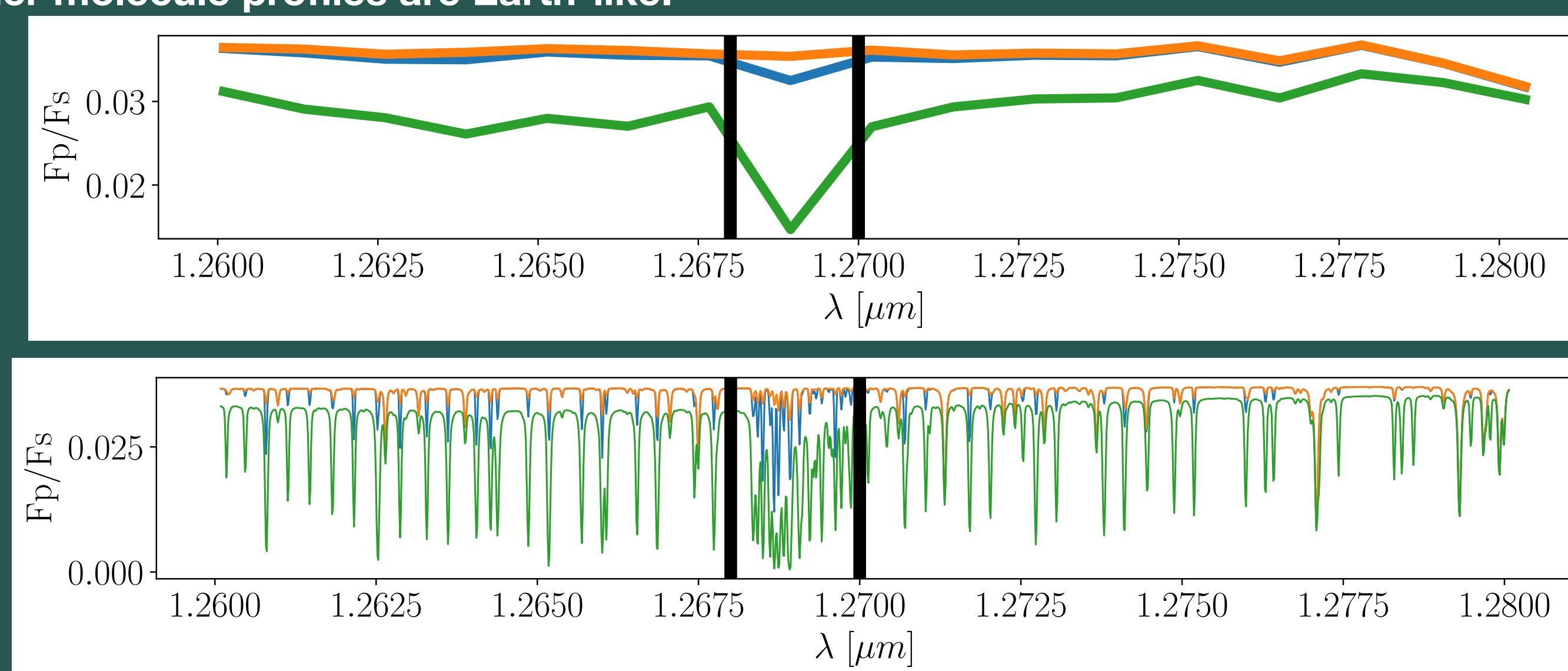
**Abiotic O<sub>2</sub>**  
(Photochemistry)

**Biological O<sub>2</sub>**  
With same column abundance as abiotic case  
(Photosynthetic Life)

**Earth**

## 1.27 μm O<sub>2</sub> Band

We create low (R=1,000) and high (R=100,000) resolution spectra from these oxygen profiles using SMART, our one-dimensional, multi-scattering, multi-stream line-by-line atmospheric modeling algorithm. Below is the result of simulating TRAPPIST-1e around TRAPPIST-1. All other molecule profiles are Earth-like.



## Branch Ratios

Because the Q branch of this spectroscopic band is more sensitive to changes in the oxygen profile than the P and R branches, we hypothesize that the Q branch flux will saturate before the others. Thus, the ratio of the fluxes in the Q branch to the P and R branches may be sufficient for probing the distribution of oxygen in the atmosphere. Below we show the calculated ratios for each of the above scenarios.

$$\text{Ratio} = \frac{Q_{\text{flux}}}{P_{\text{flux}} + R_{\text{flux}}}$$

	Abiotic O <sub>2</sub>	Biological O <sub>2</sub>	Earth
Low-res	0.054	0.058	0.030
High-res	0.096	0.101	0.068

## Ground-Based High-Res Spectroscopy

- Potential solution for solving one of the trickiest false positives
- Current space-based instruments are not capable of the resolution needed to distinguish this false positive (R ≥ 100,000)
- The upcoming extremely large (30-40m) ground-based observatories will carry the high-resolution instruments needed for using this method for distinguishing this false positive

## Conclusions

- We may be able to distinguish evenly mixed biological O<sub>2</sub> from the abiotic production and buildup of oxygen in the upper atmosphere by analyzing the P, Q, and R branches of the 1.27 μm O<sub>2</sub> band
- This false positive remains very challenging to discriminate, but this method is a promising start

## Future Work

- Use a coupled climate-photochem model to better understand the atmospheric distribution of photochemically produced O<sub>2</sub>
- Investigate the detectability and retrieval of these types of atmospheres
- Investigate the use of atmospheric pressure indicators as another method to distinguish this false positive

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