

Microlensing Planets: Modeling and Unique Solutions

Jennifer Yee

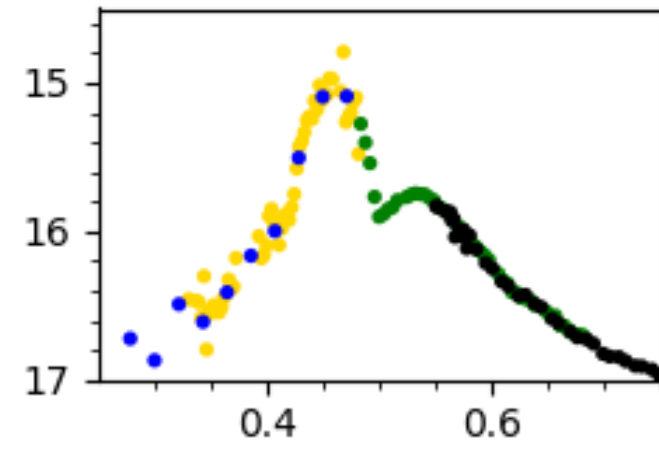
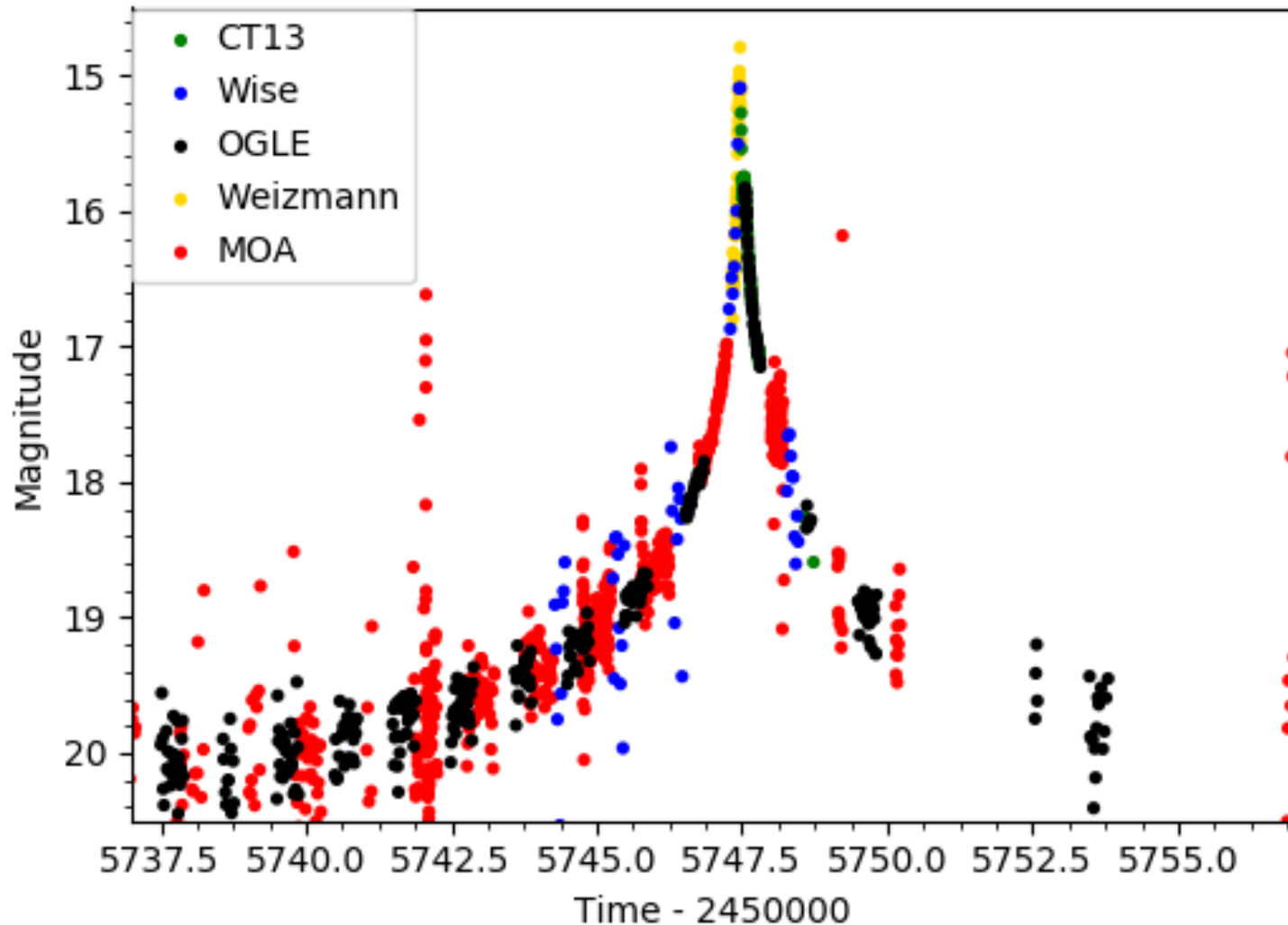
Center for Astrophysics

How Do I Figure Out If It's Really a Planet?

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Center for Astrophysics

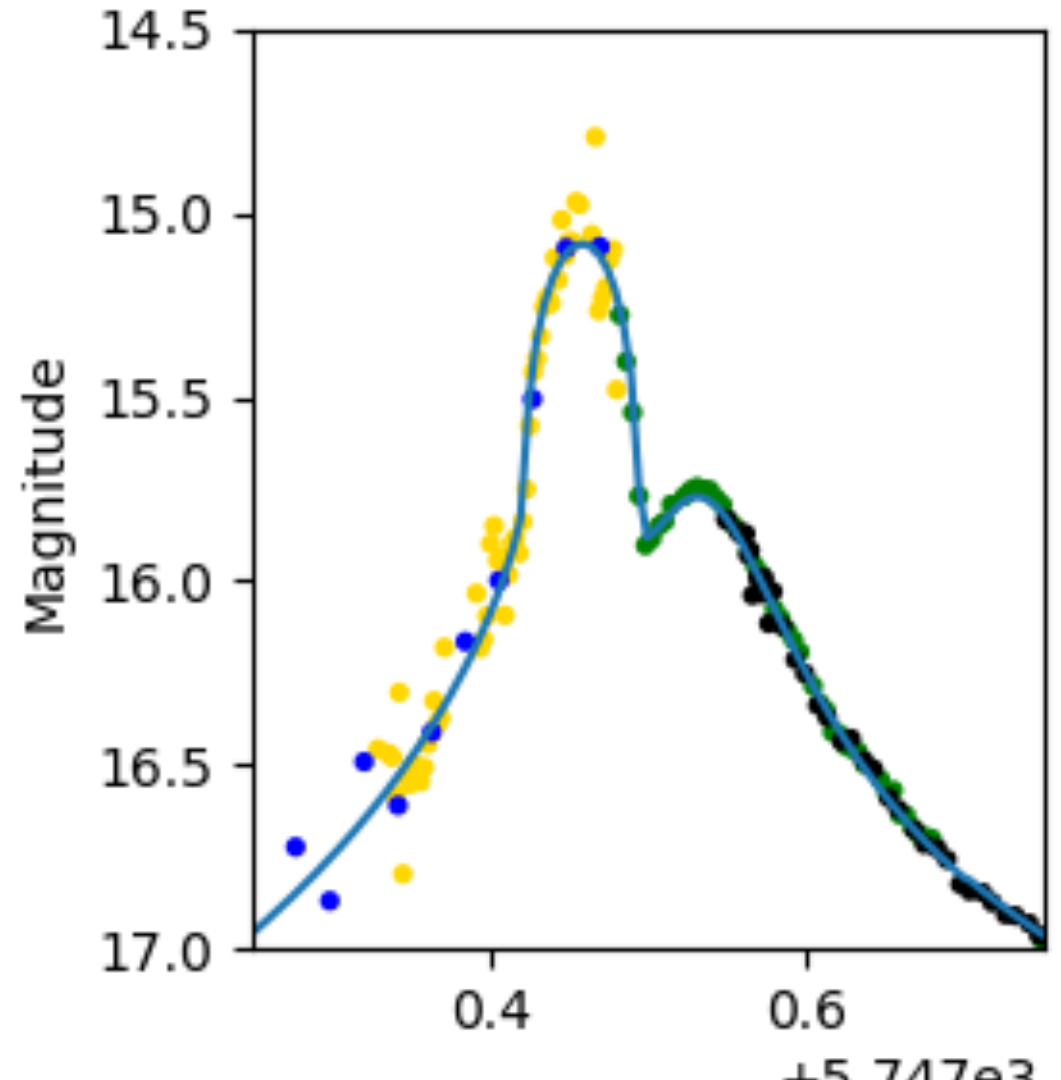
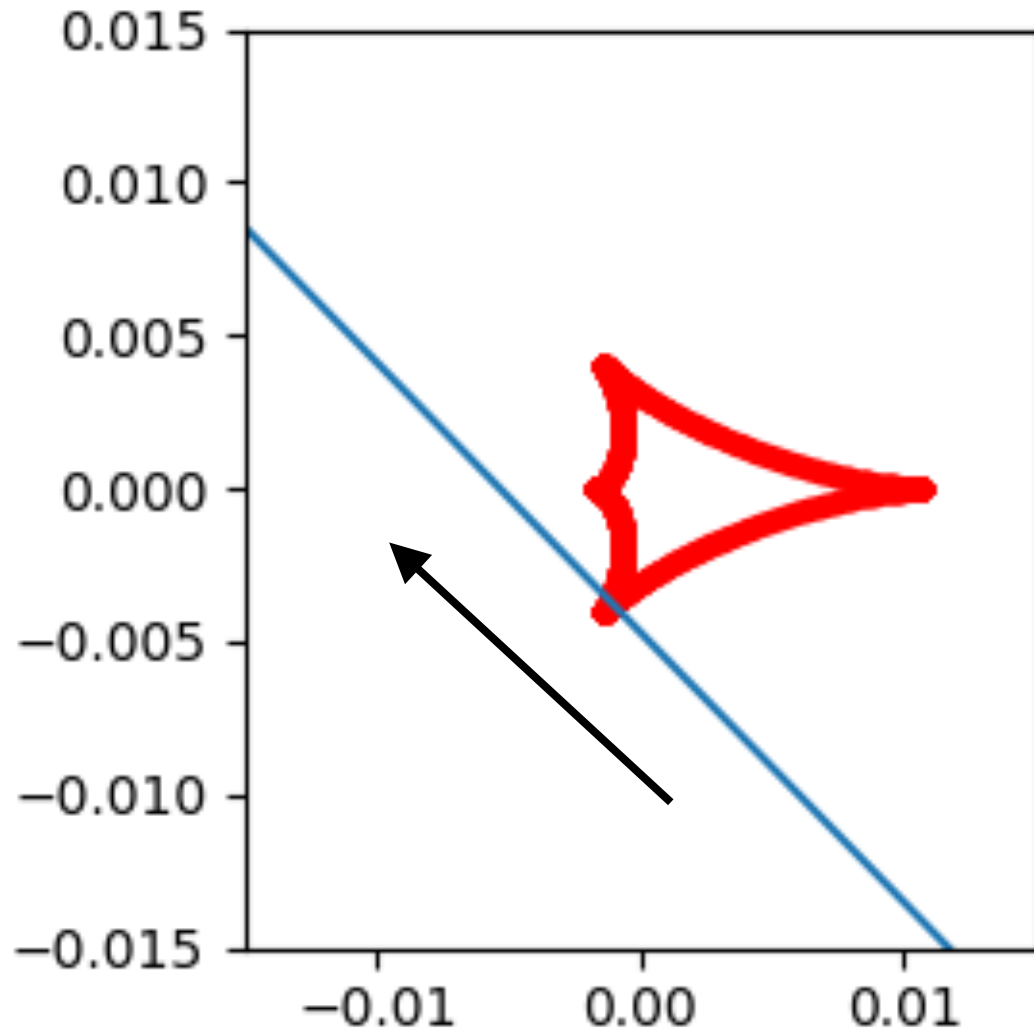
Why is this a Planet Candidate?



1. Mostly follows a point lens \rightarrow dominated by a single star.
2. The perturbation occurs at high magnification \rightarrow caustic is small.
3. The perturbation is small (< 1 mag) \rightarrow Could be a planet!

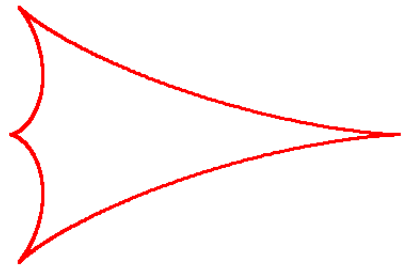
Part I: Relating
Caustics/Magnification Patterns
to Light Curves

A model with mass ratio (q) = 0.005 = 5 $M_{\text{Jup}}/M_{\text{Sun}}$

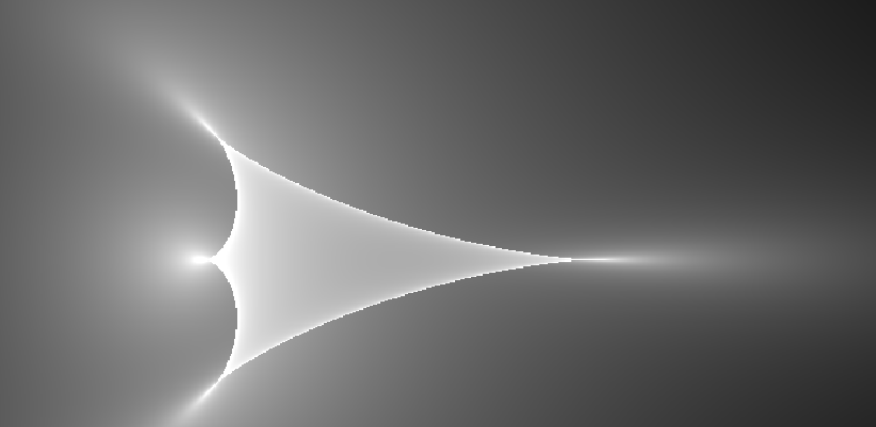


Caustic Reflects the Full Magnification Pattern

Caustic

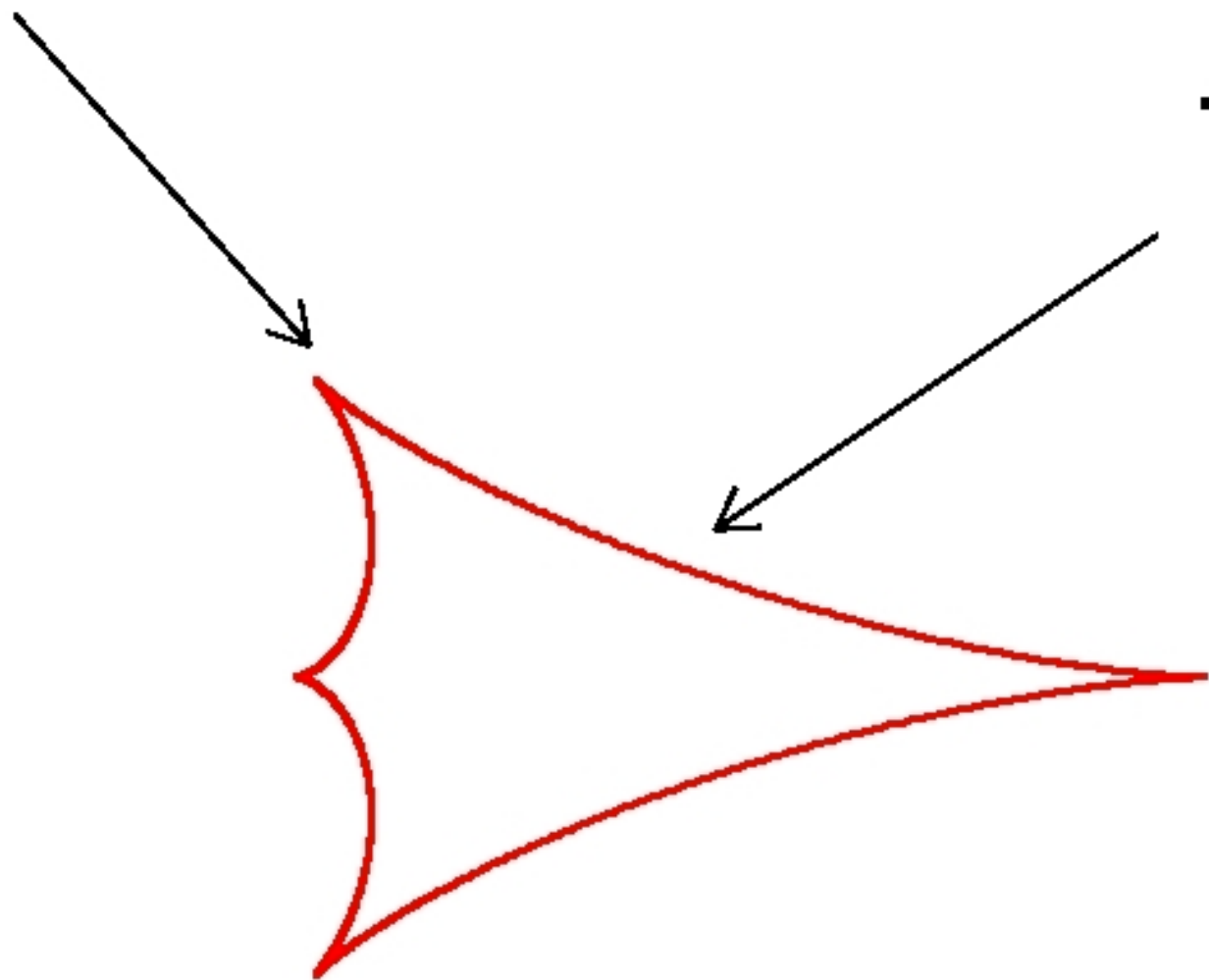


Magnification Pattern

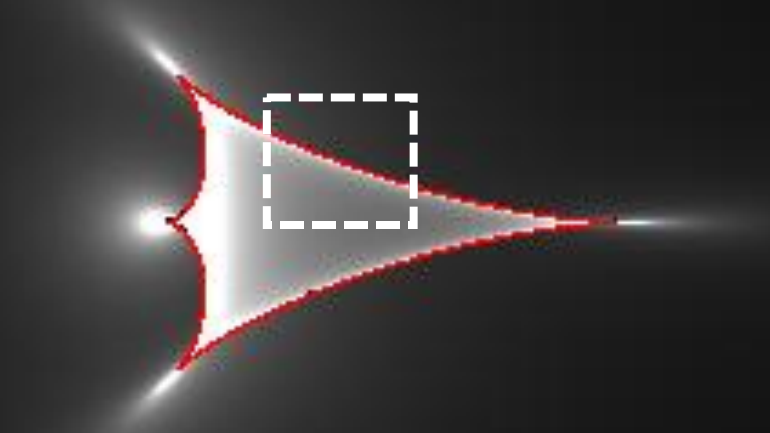


cusp

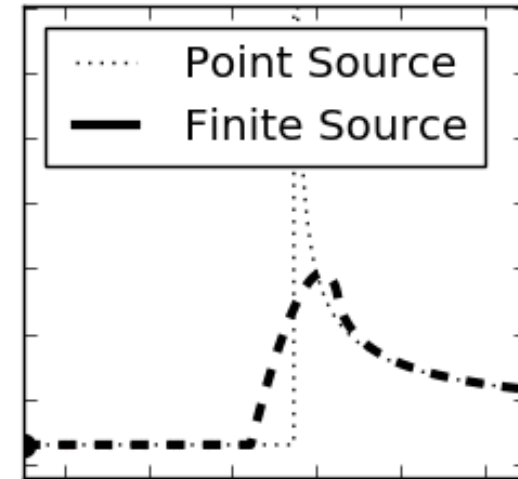
fold



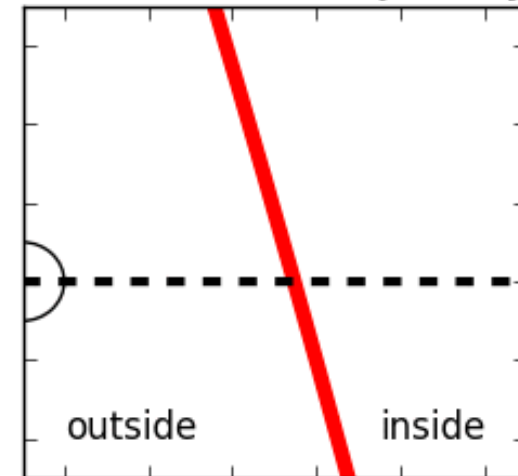
There is a strong magnification change at a **fold**.



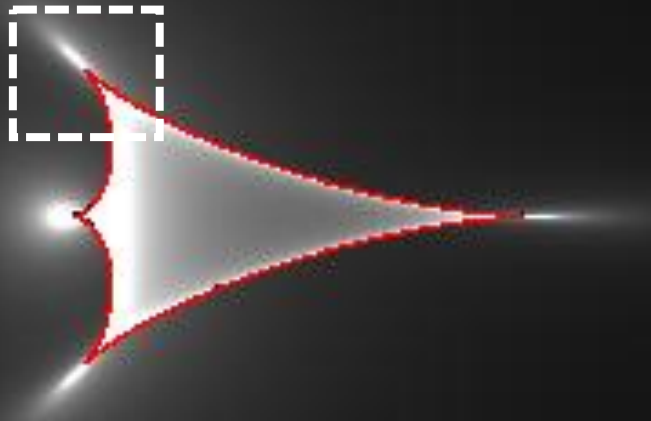
Magnification Curve



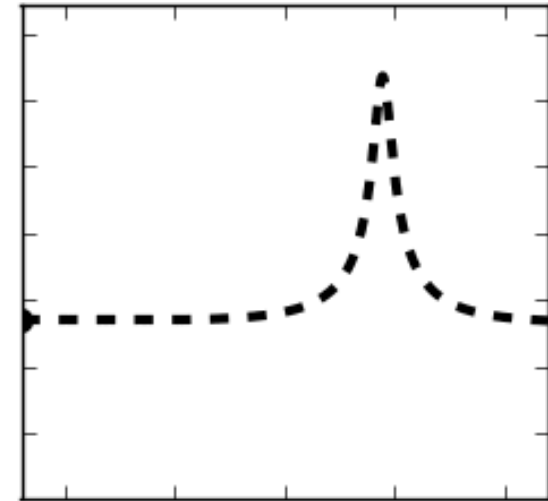
Caustics and Trajectory



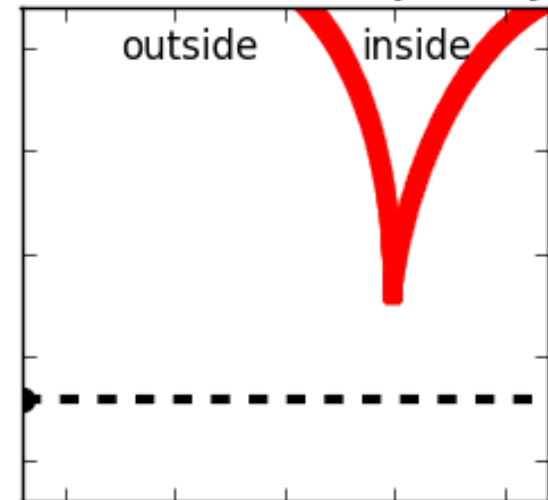
A **cusp** creates a spike in magnification outside the caustic.



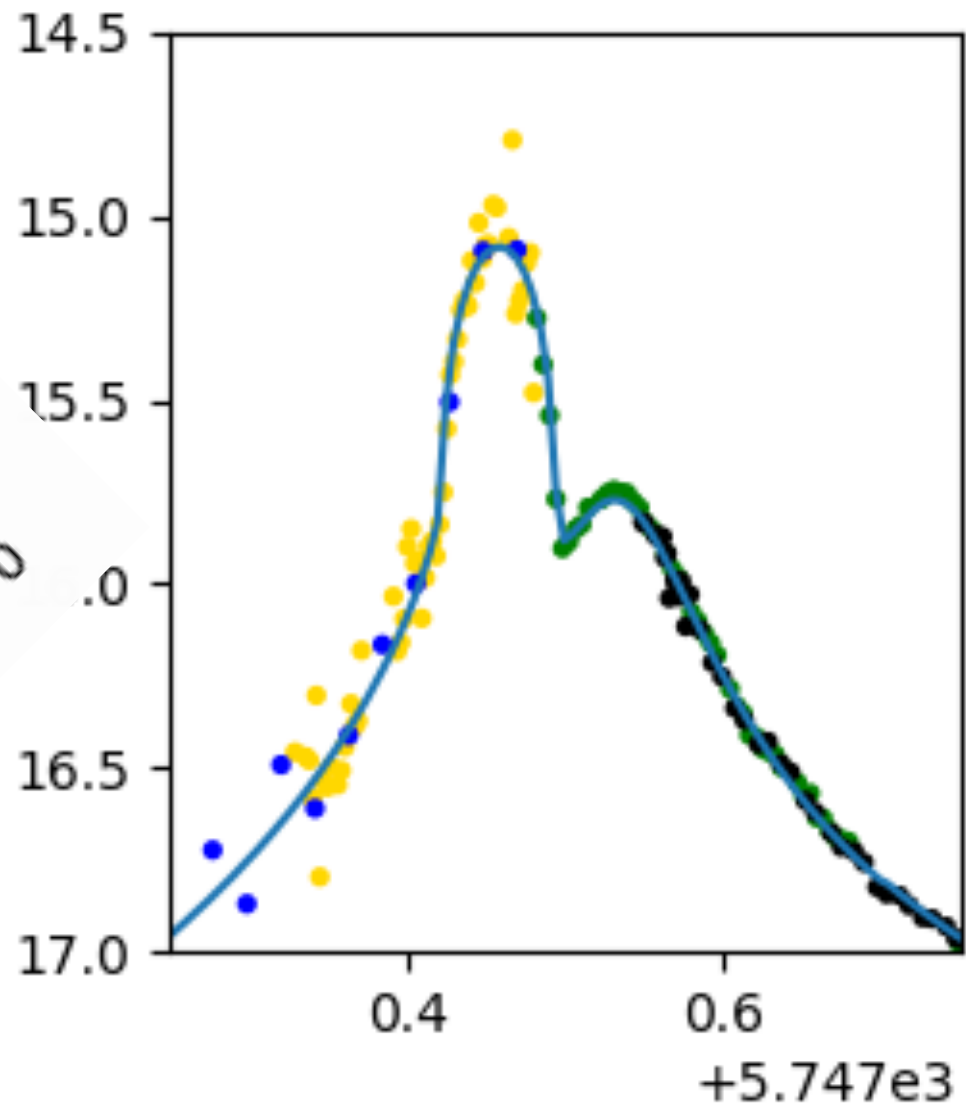
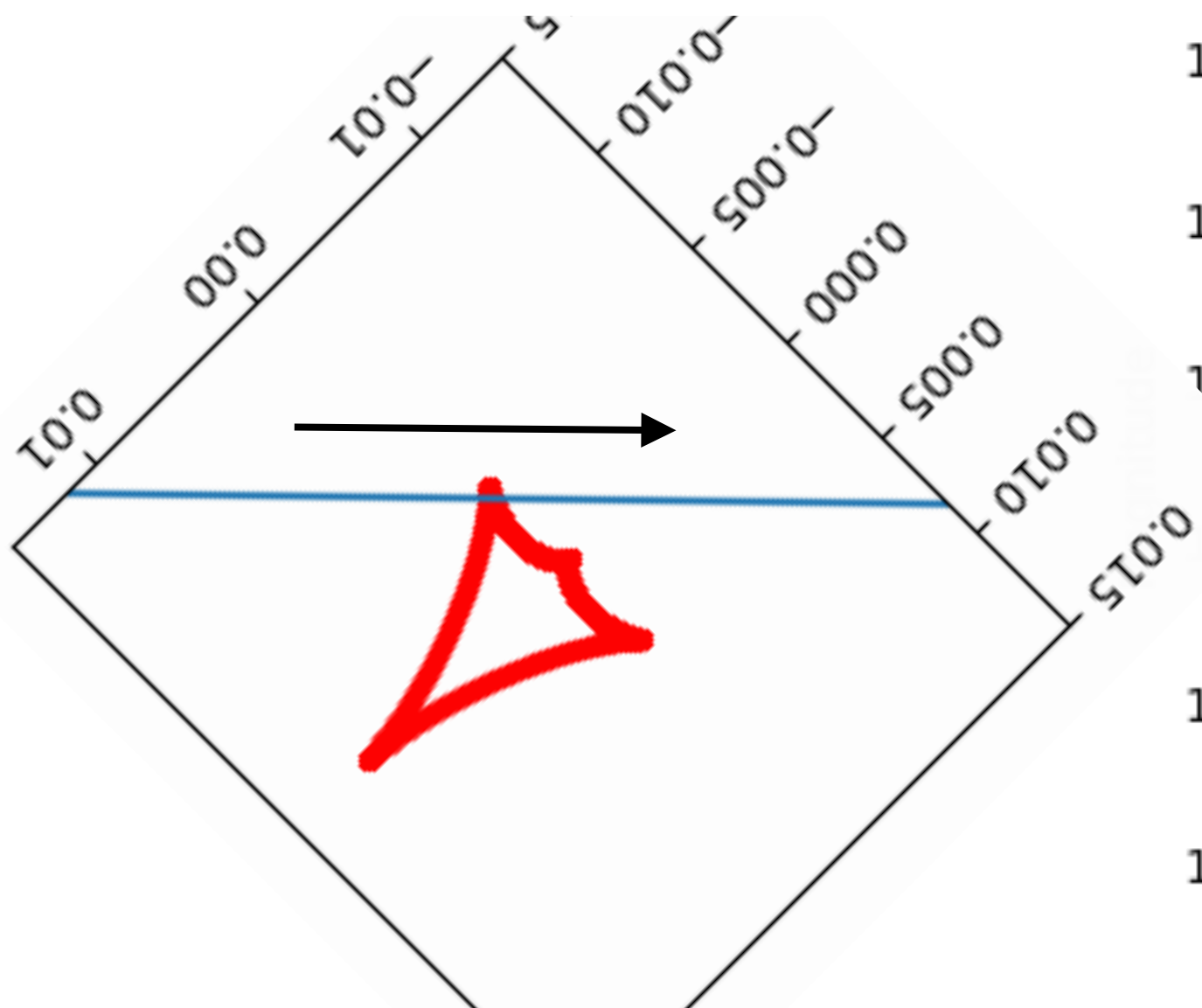
Magnification Curve



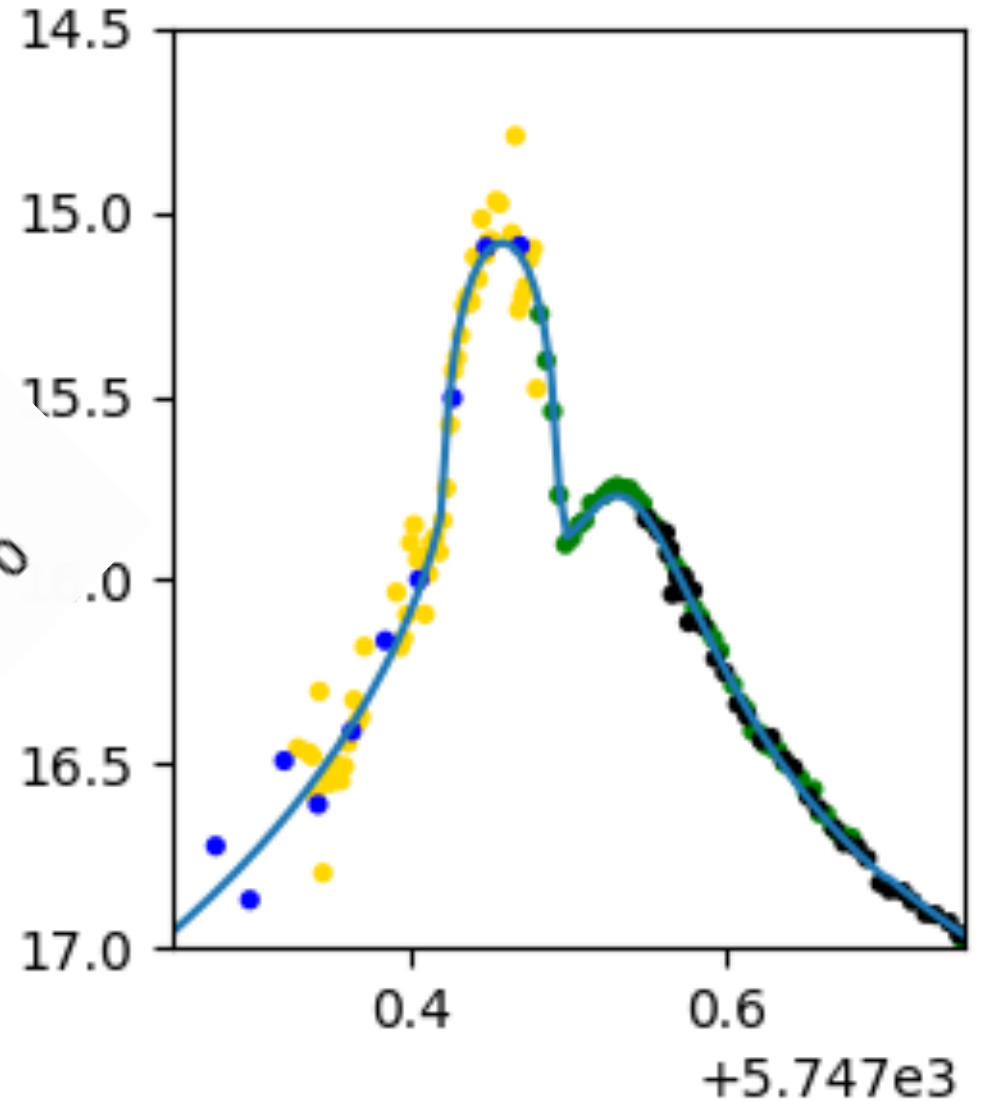
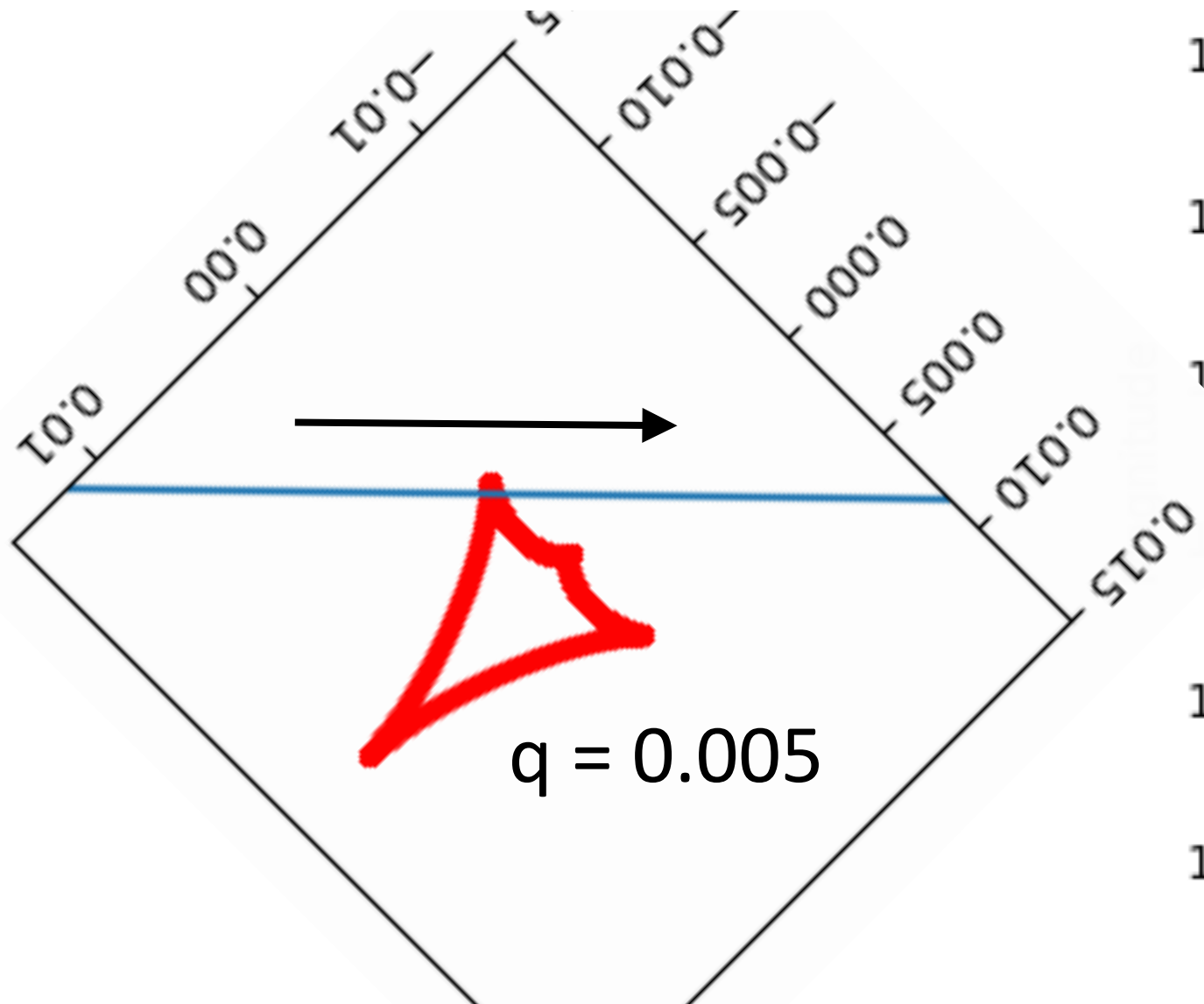
Caustics and Trajectory



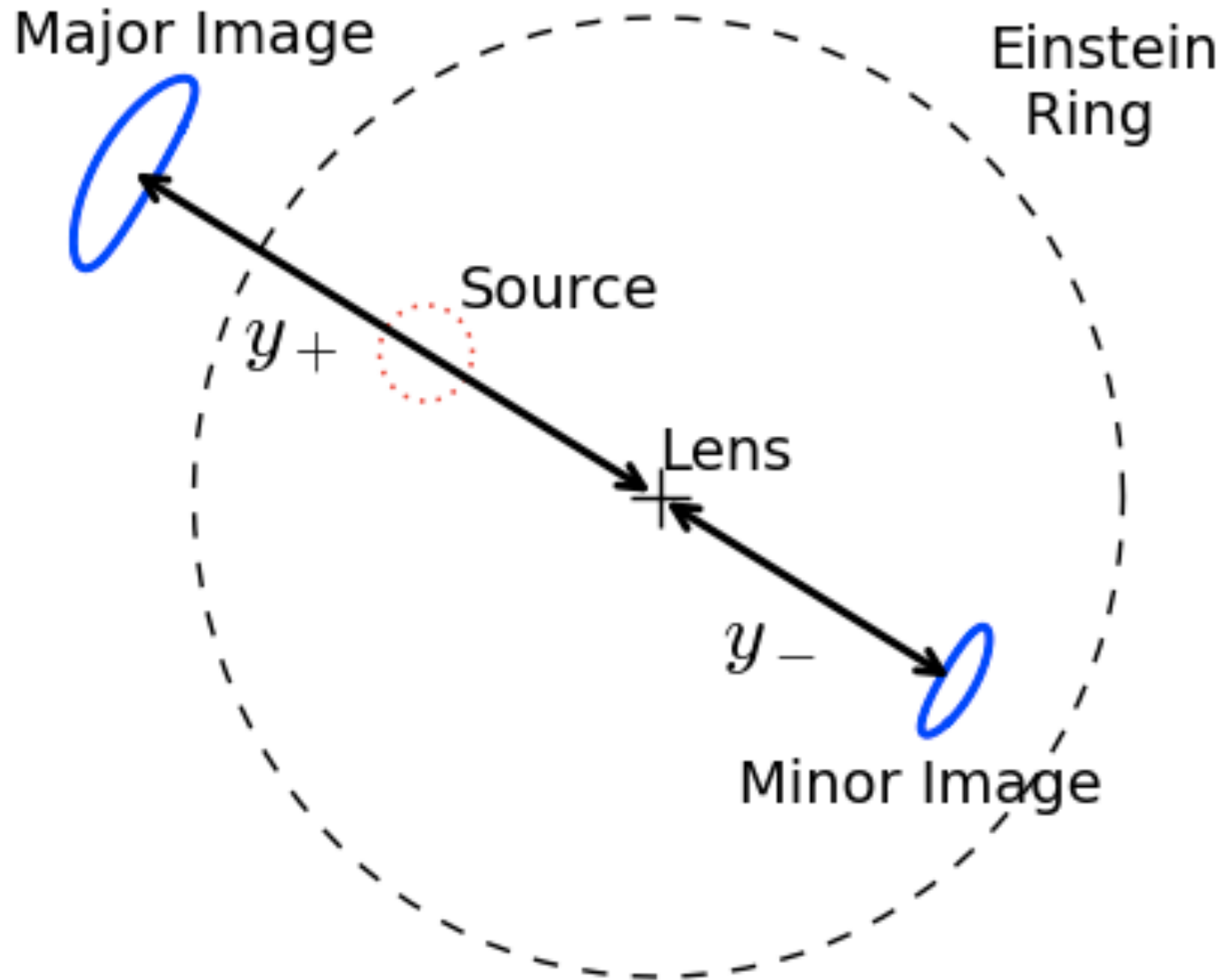
Two cusps \rightarrow two bumps



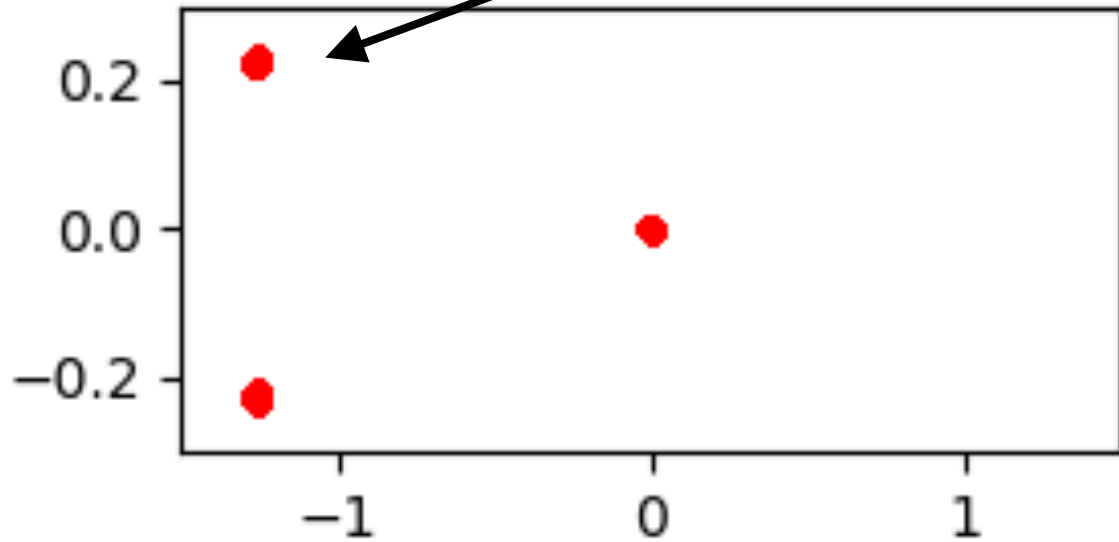
Light curve can be explained by a planet! But...



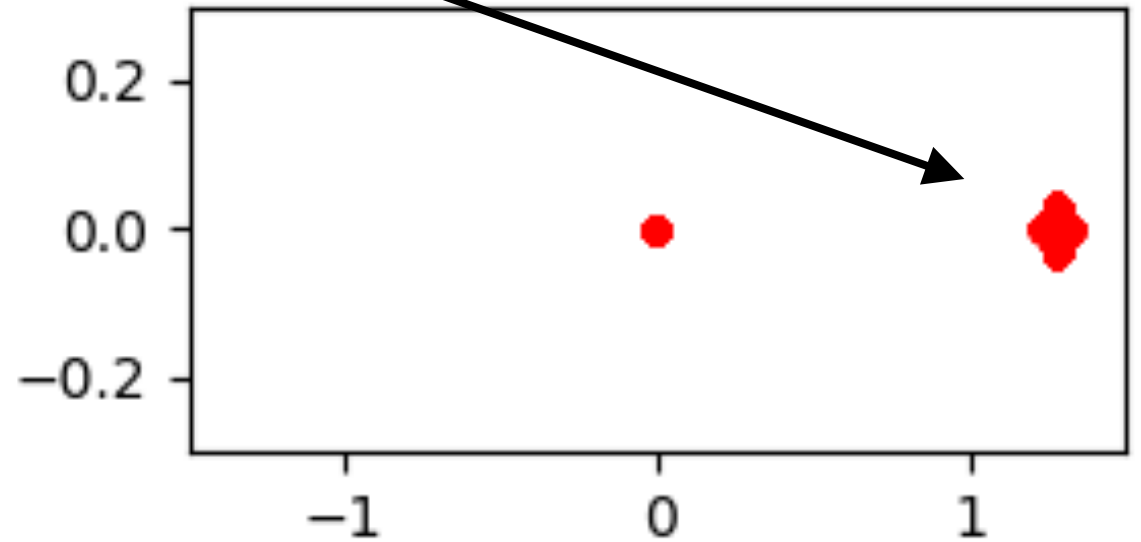
There are TWO images \rightarrow TWO solutions



The planetary caustics are different...

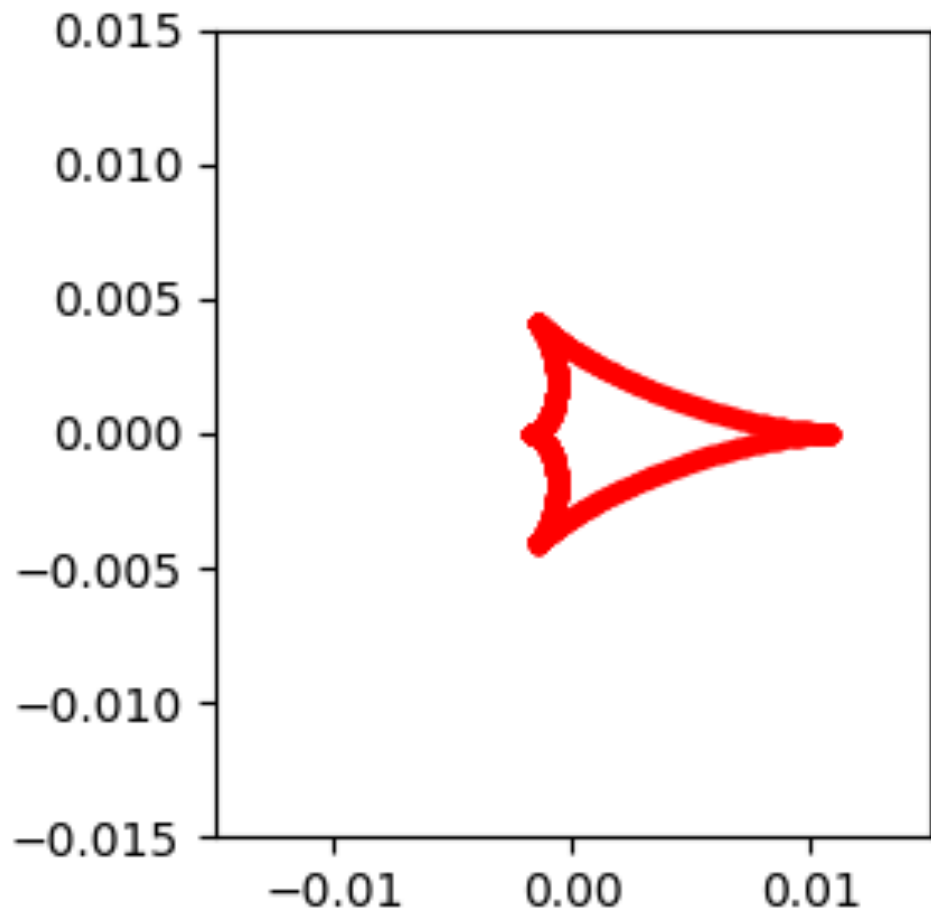


separation = $s = 0.55$
“close”



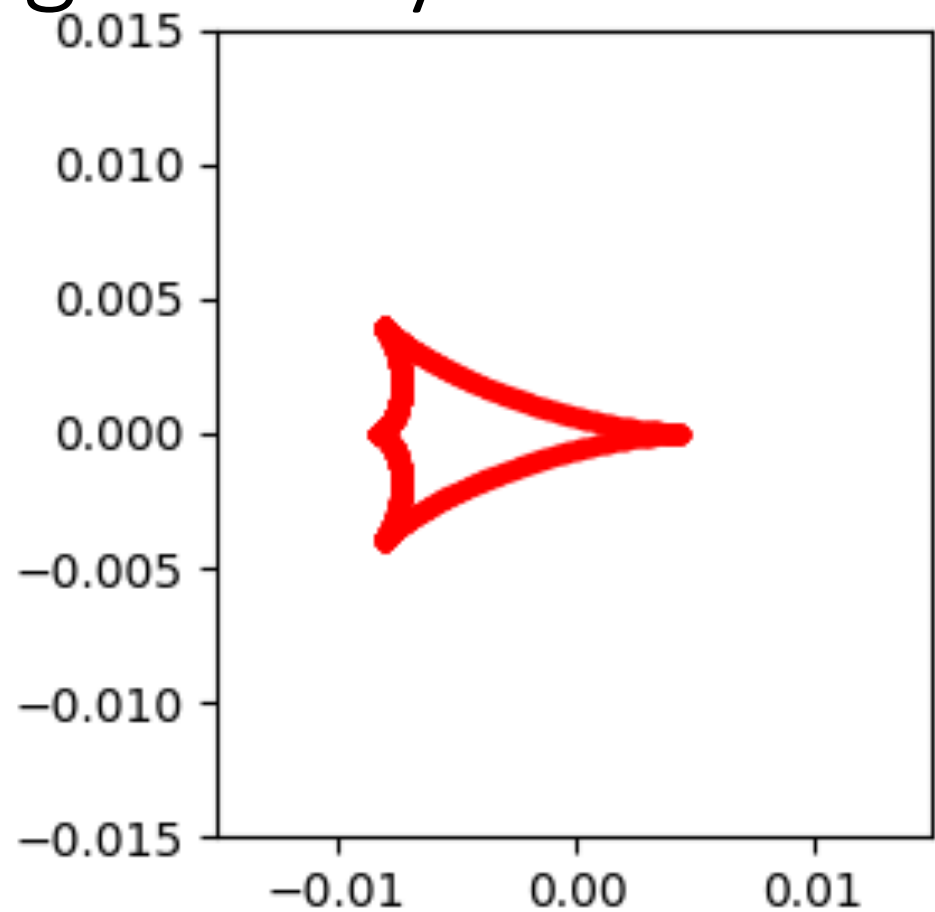
$s = 1.83$
“wide”

The central caustics are the same
“ $s \leftrightarrow 1/s$ degeneracy”



separation = $s = 0.55$

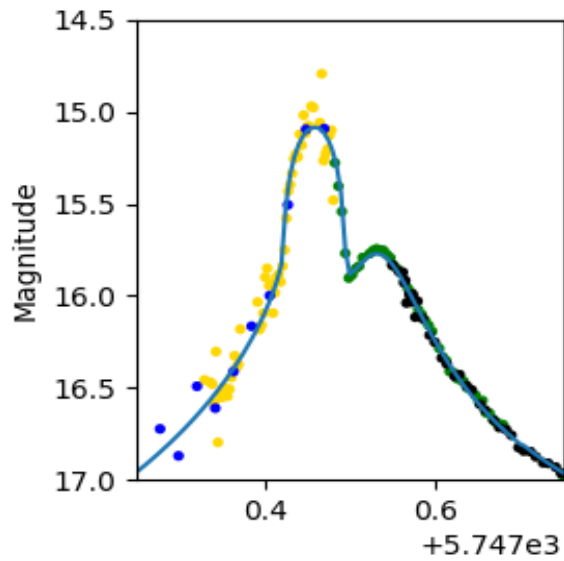
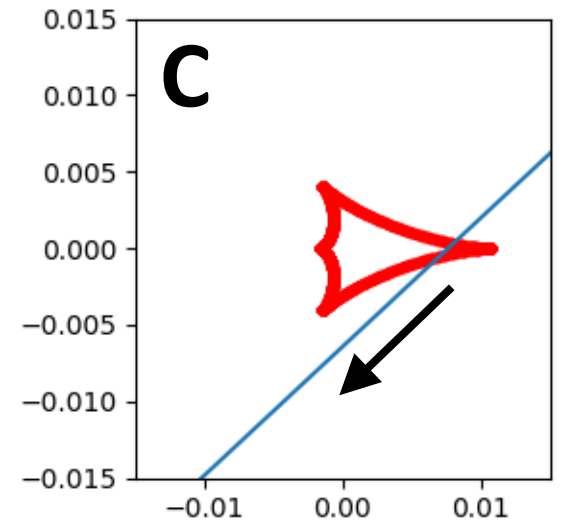
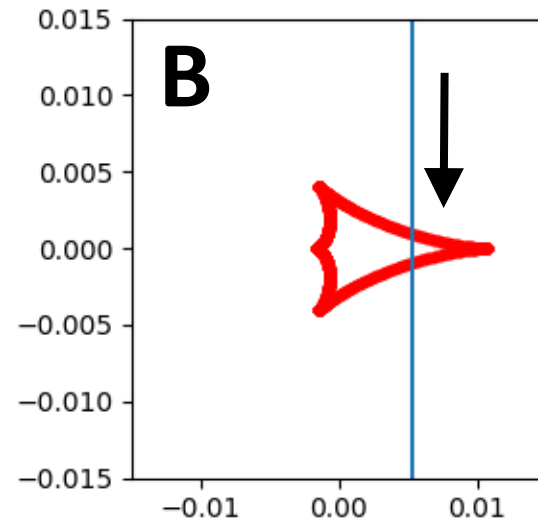
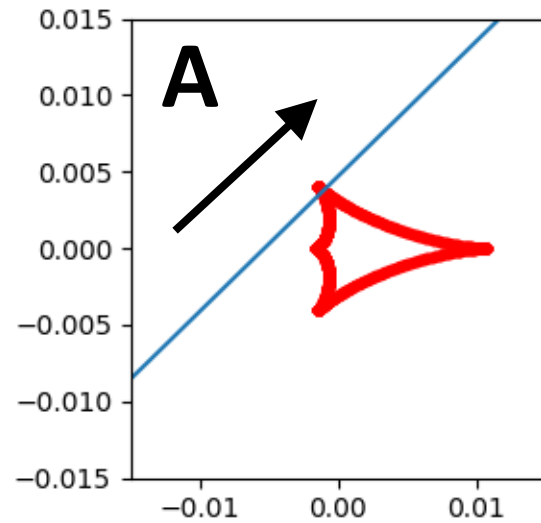
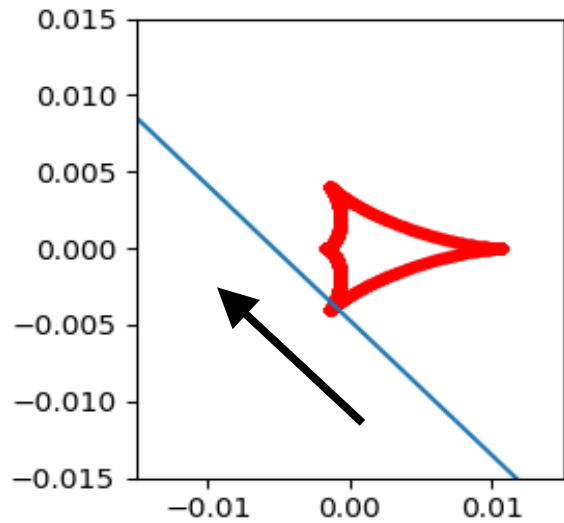
“close”



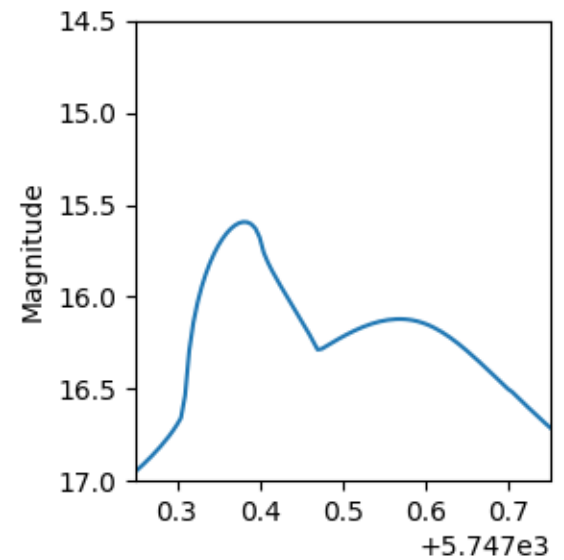
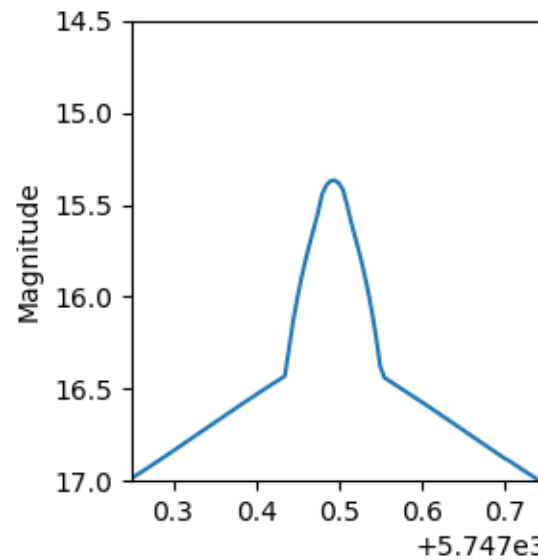
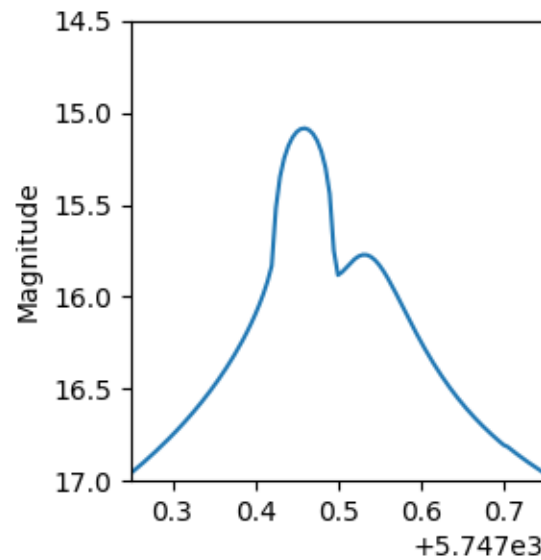
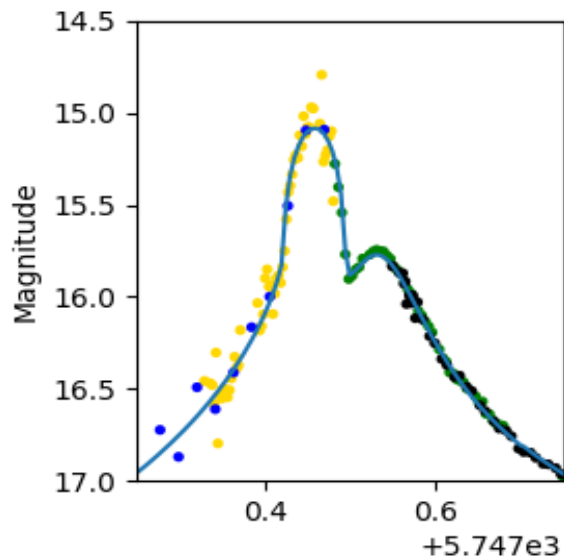
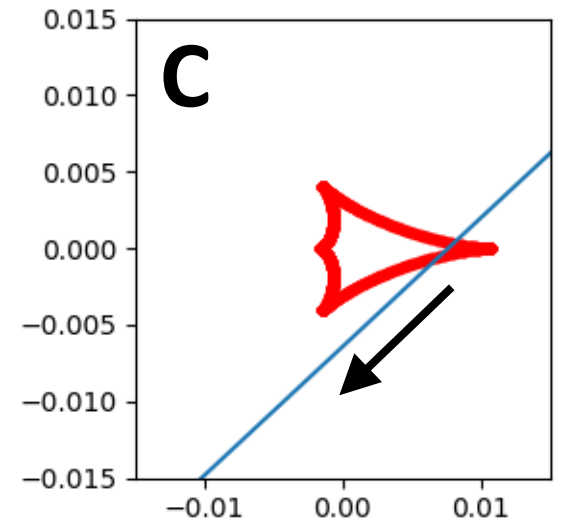
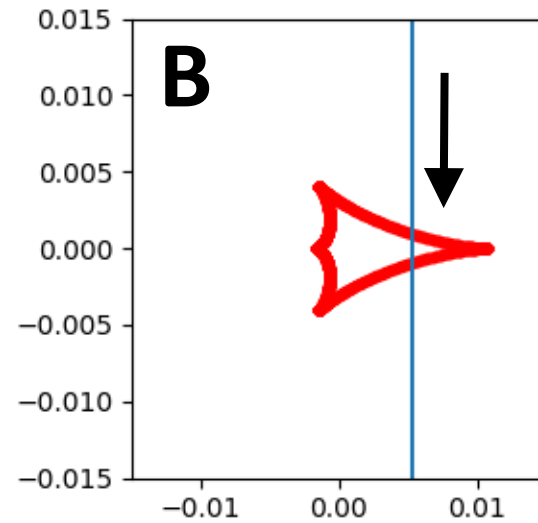
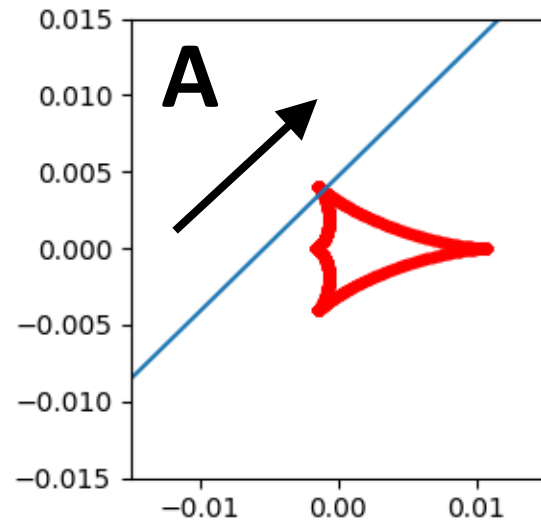
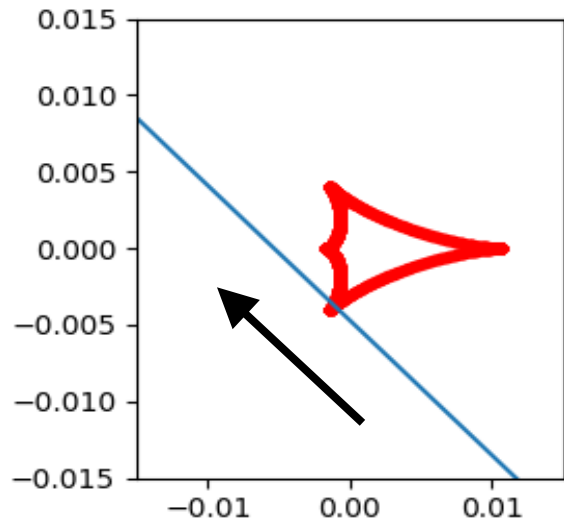
$s = 1.83$

“wide”

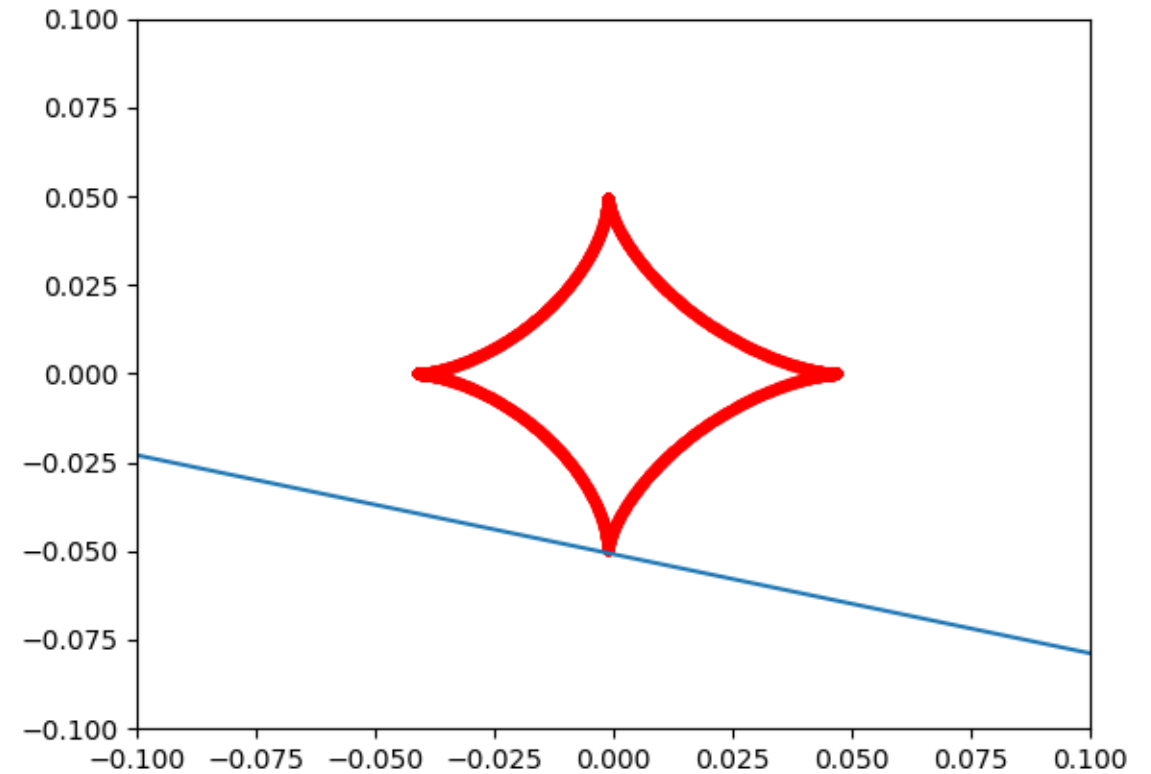
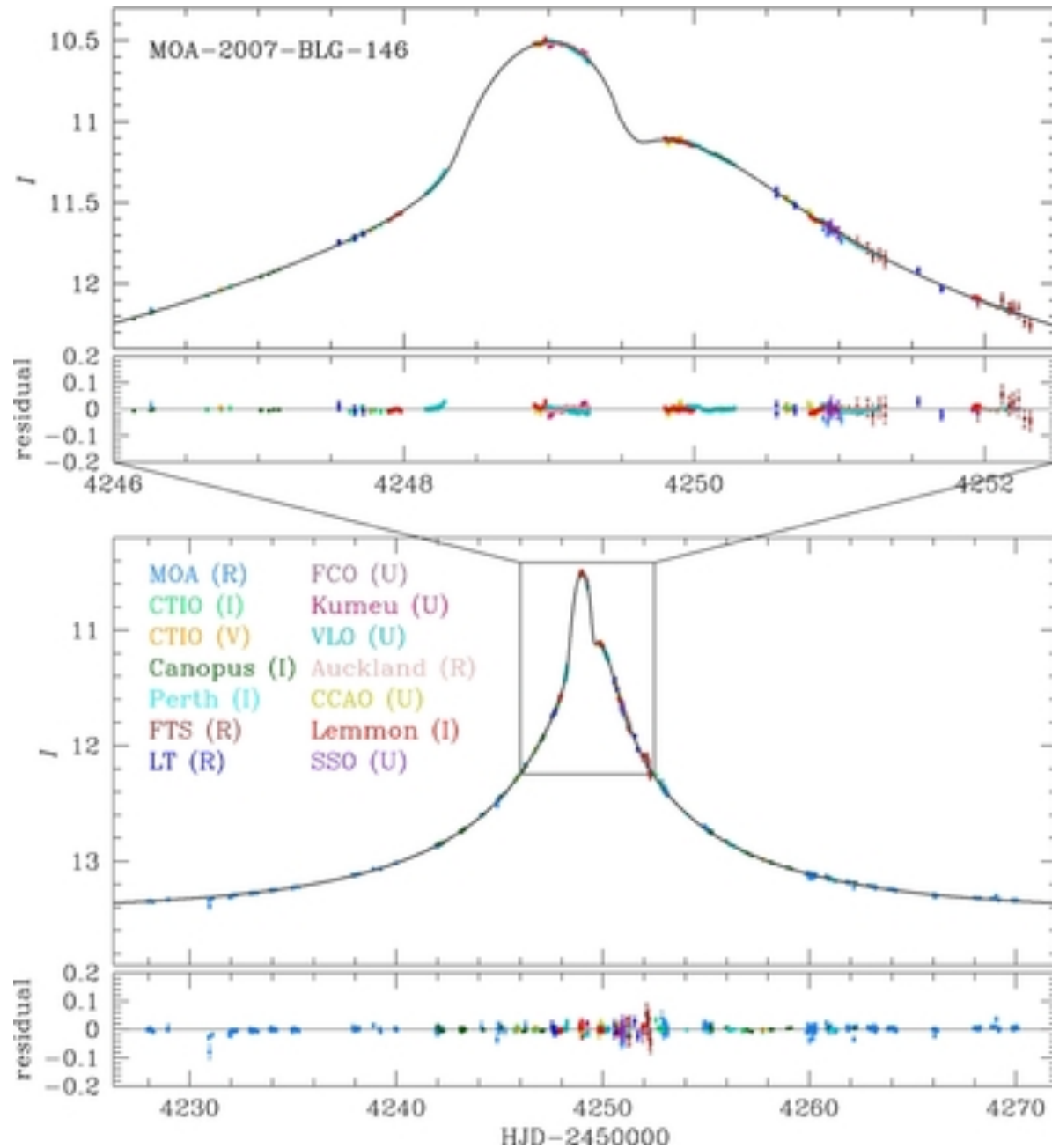
Which Trajectories Produce a Similar Light curve?



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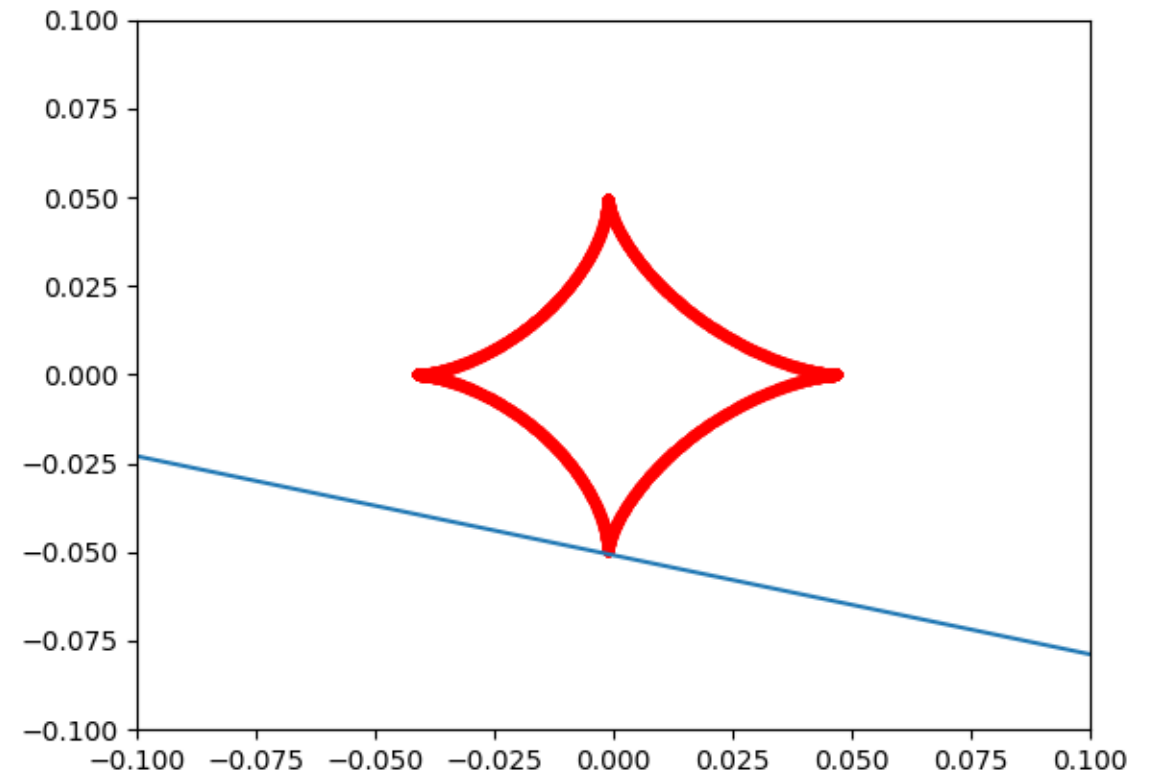
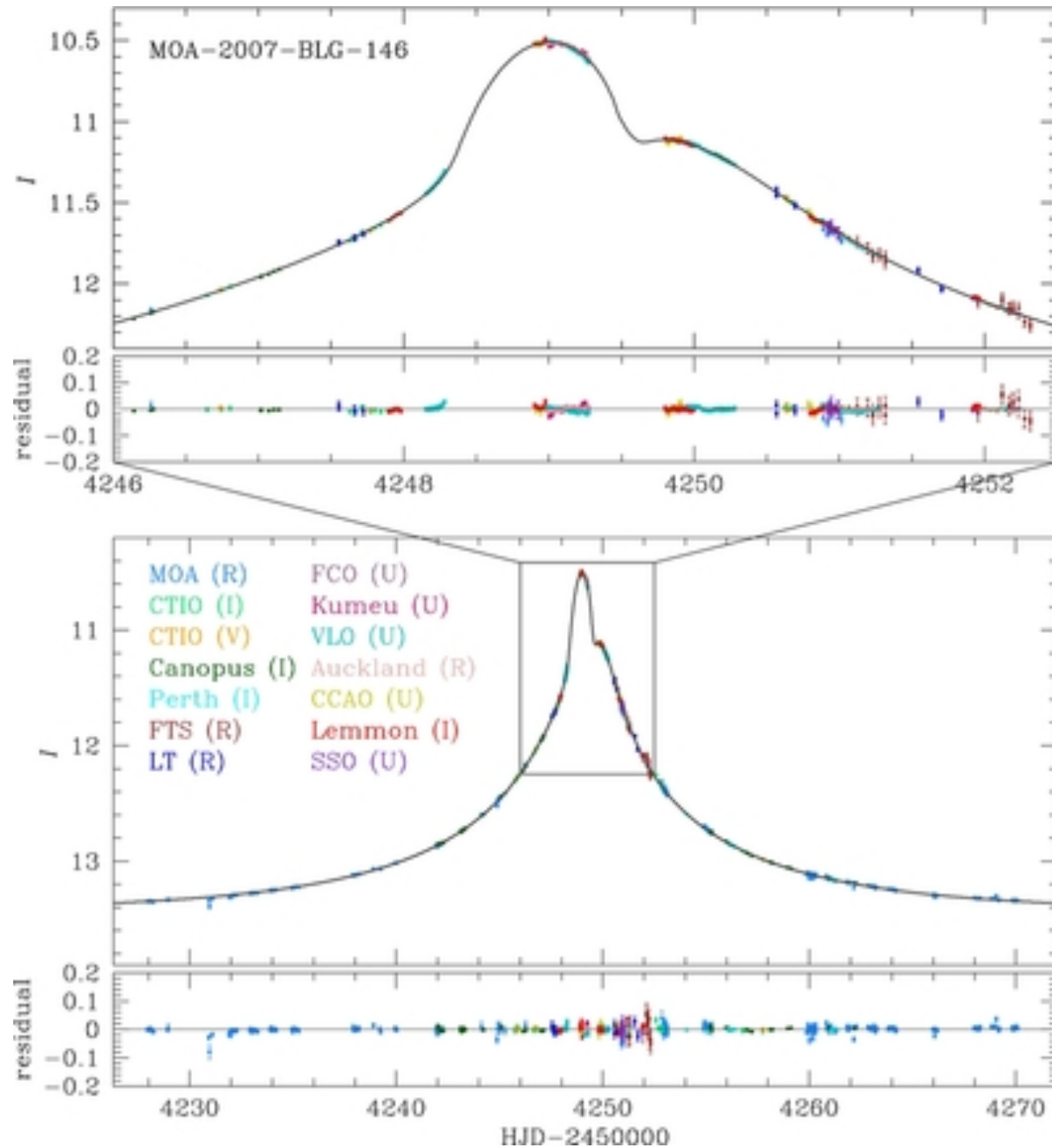


Is this also a planet?

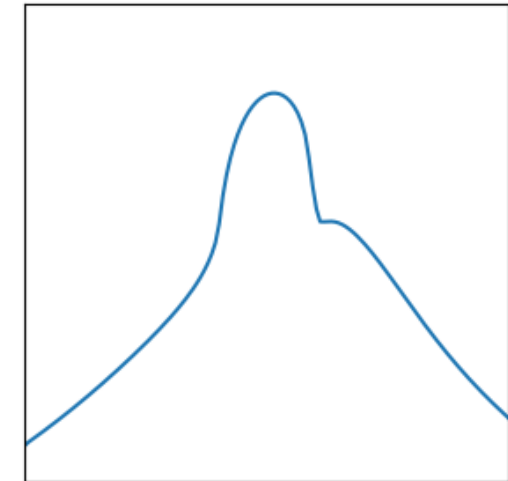
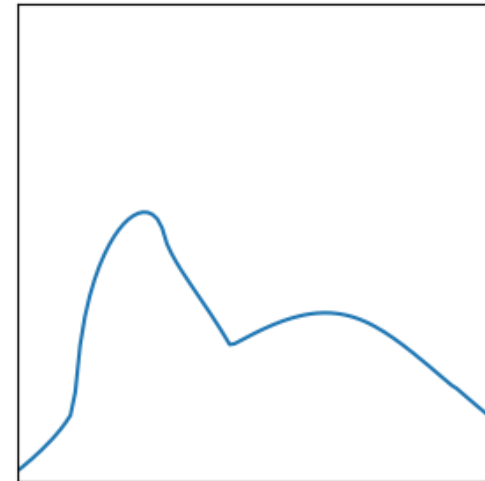
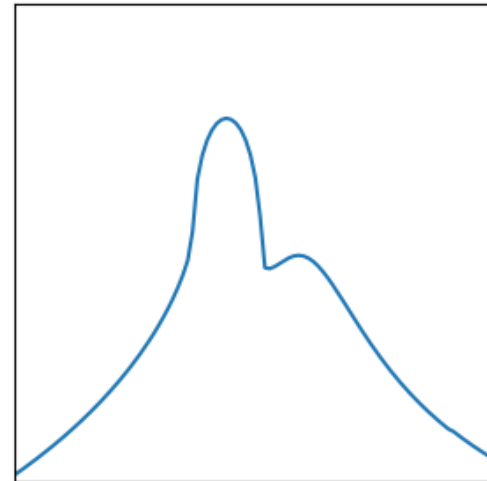
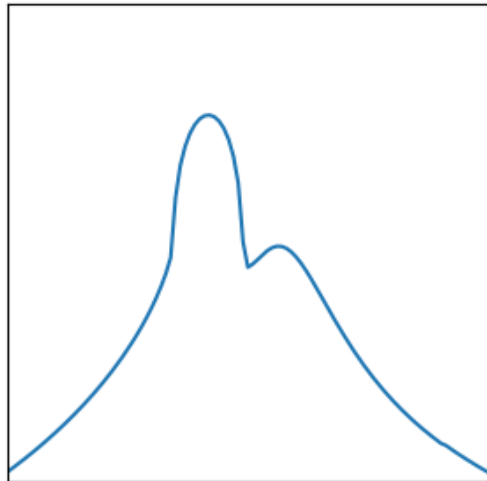
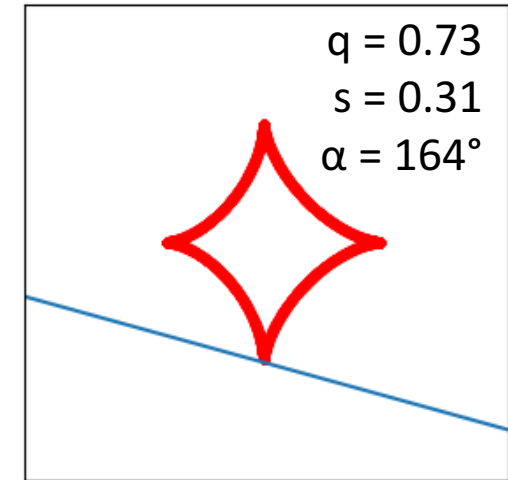
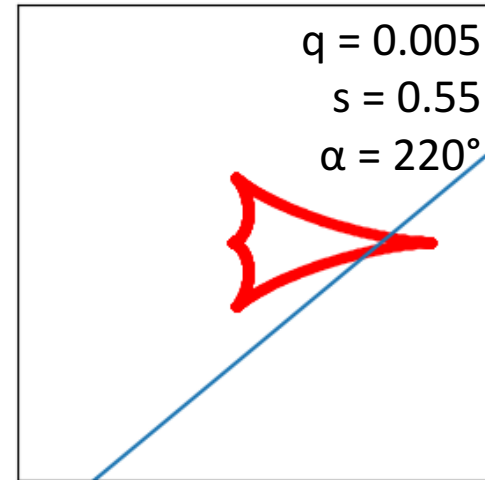
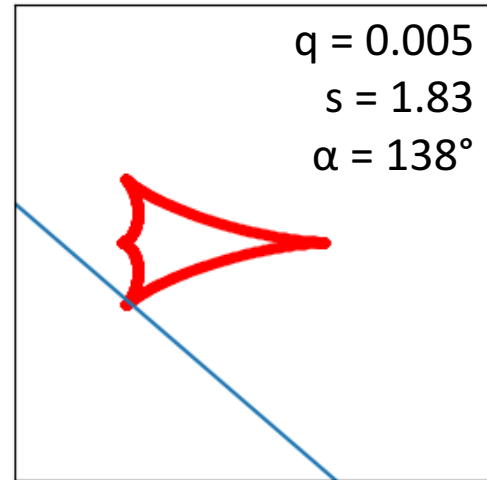
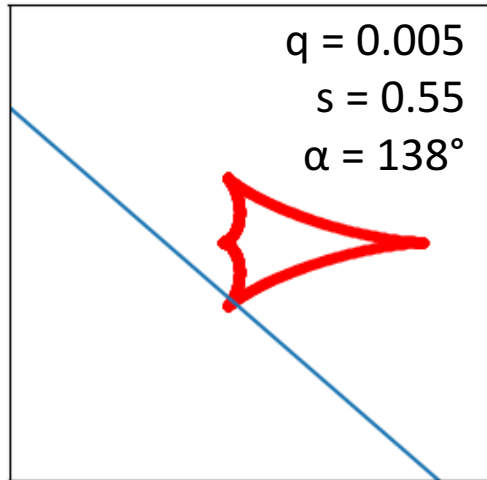


Is this also a planet?

NO. $q = 0.729$

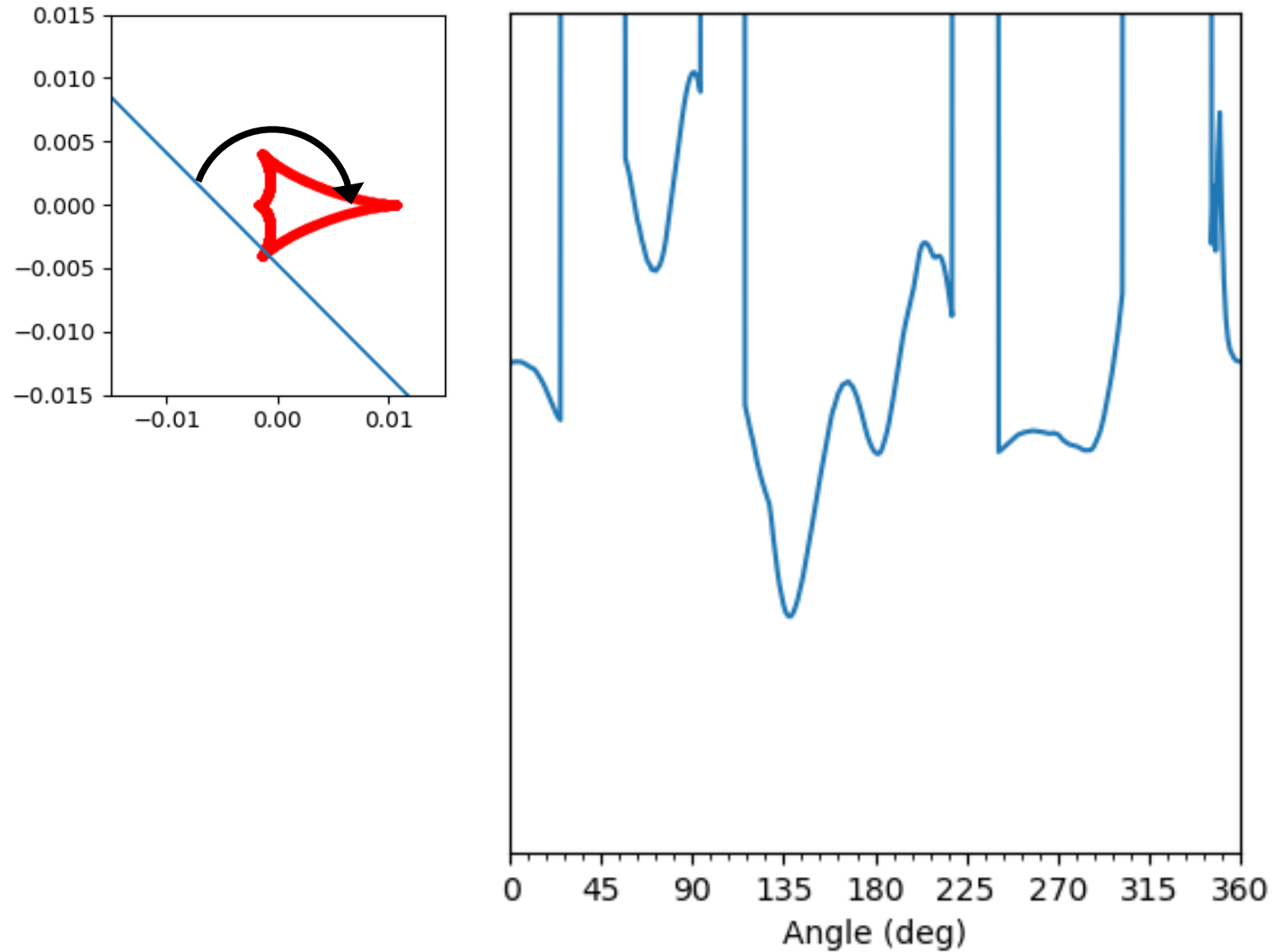


Symmetries/Similarities in the Magnification Pattern Make It Non-Trivial to Solve a Light curve

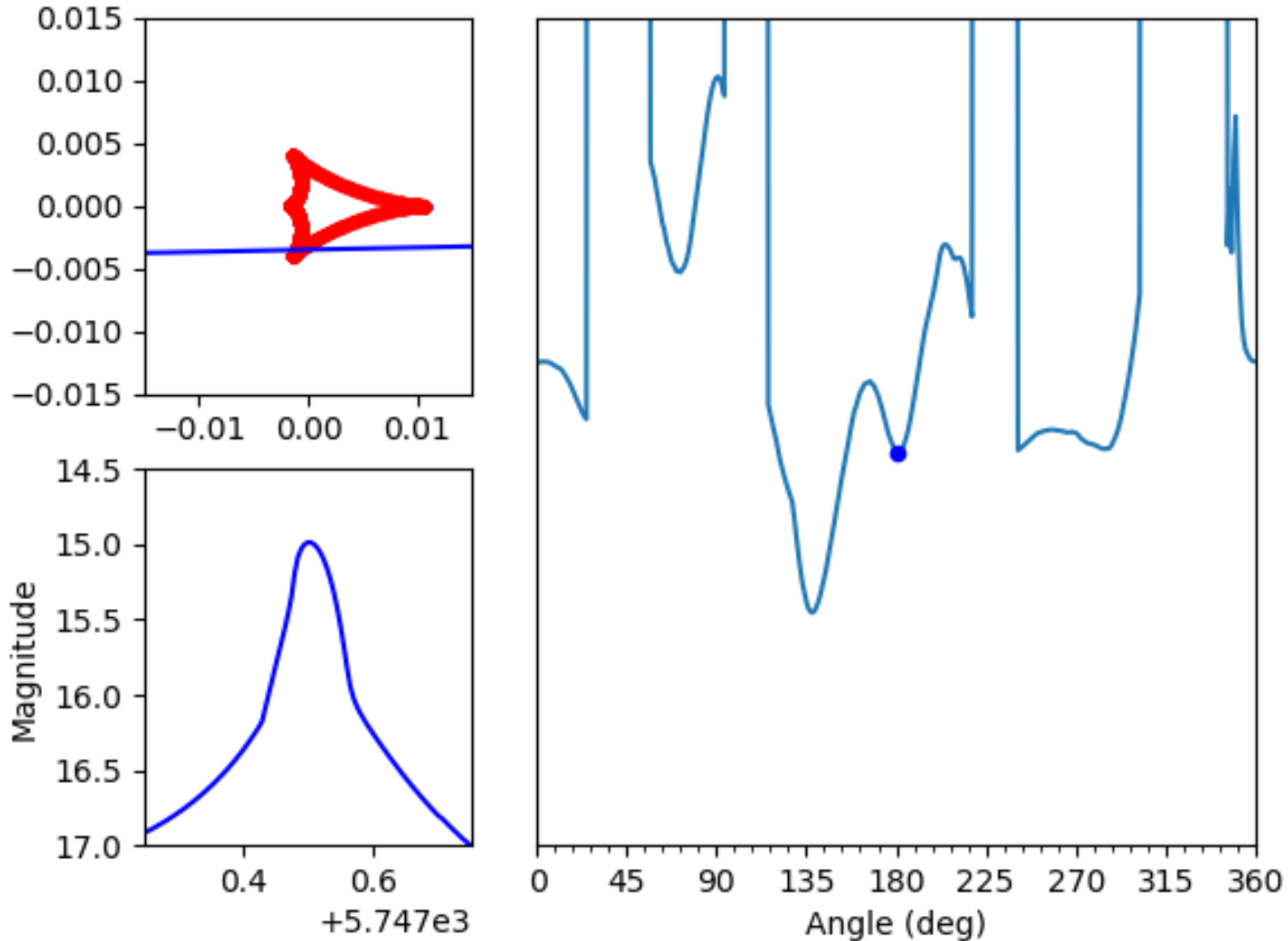


Part II: Techniques for Finding the Right Model

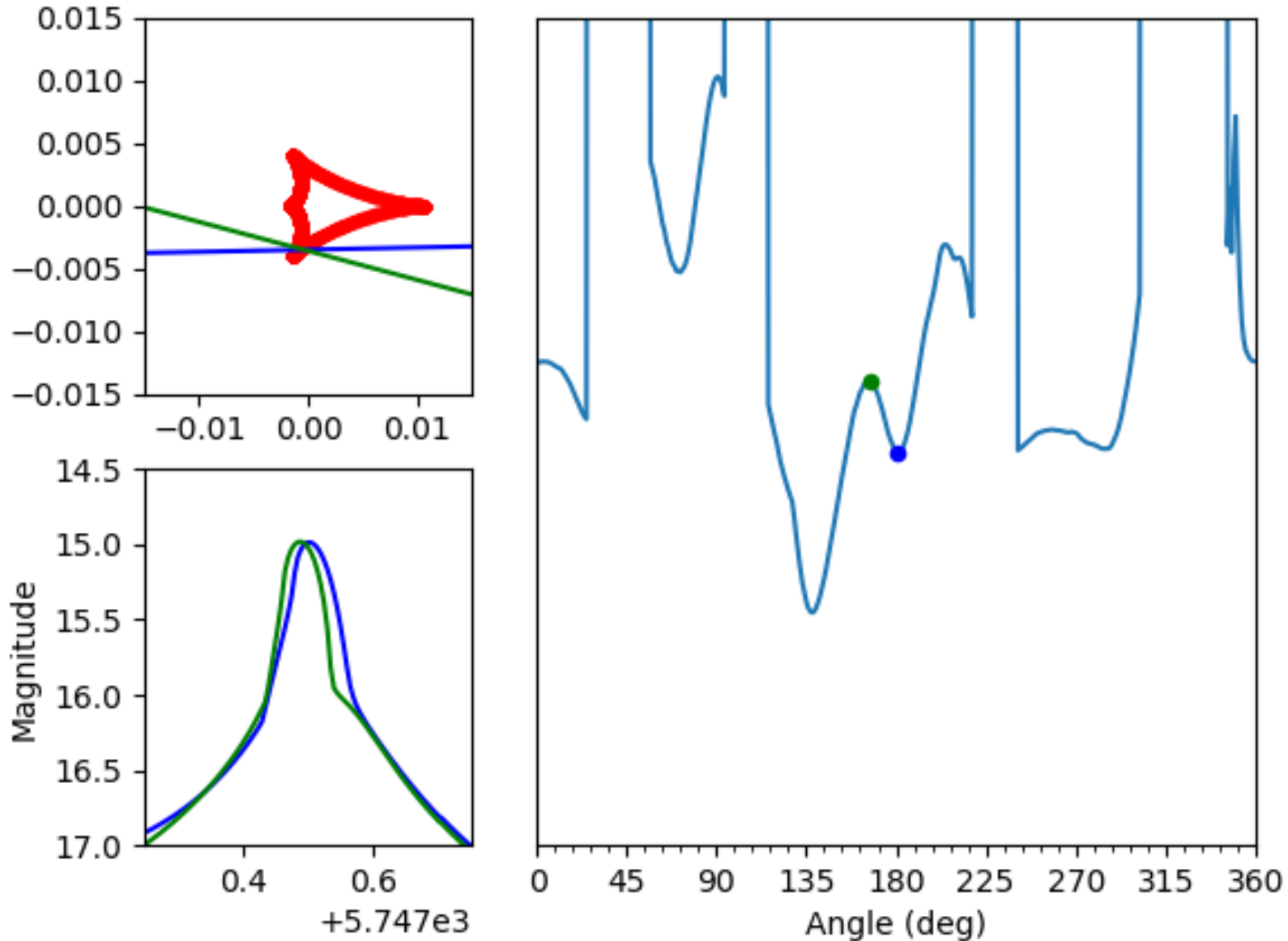
The χ^2 surface is complex.



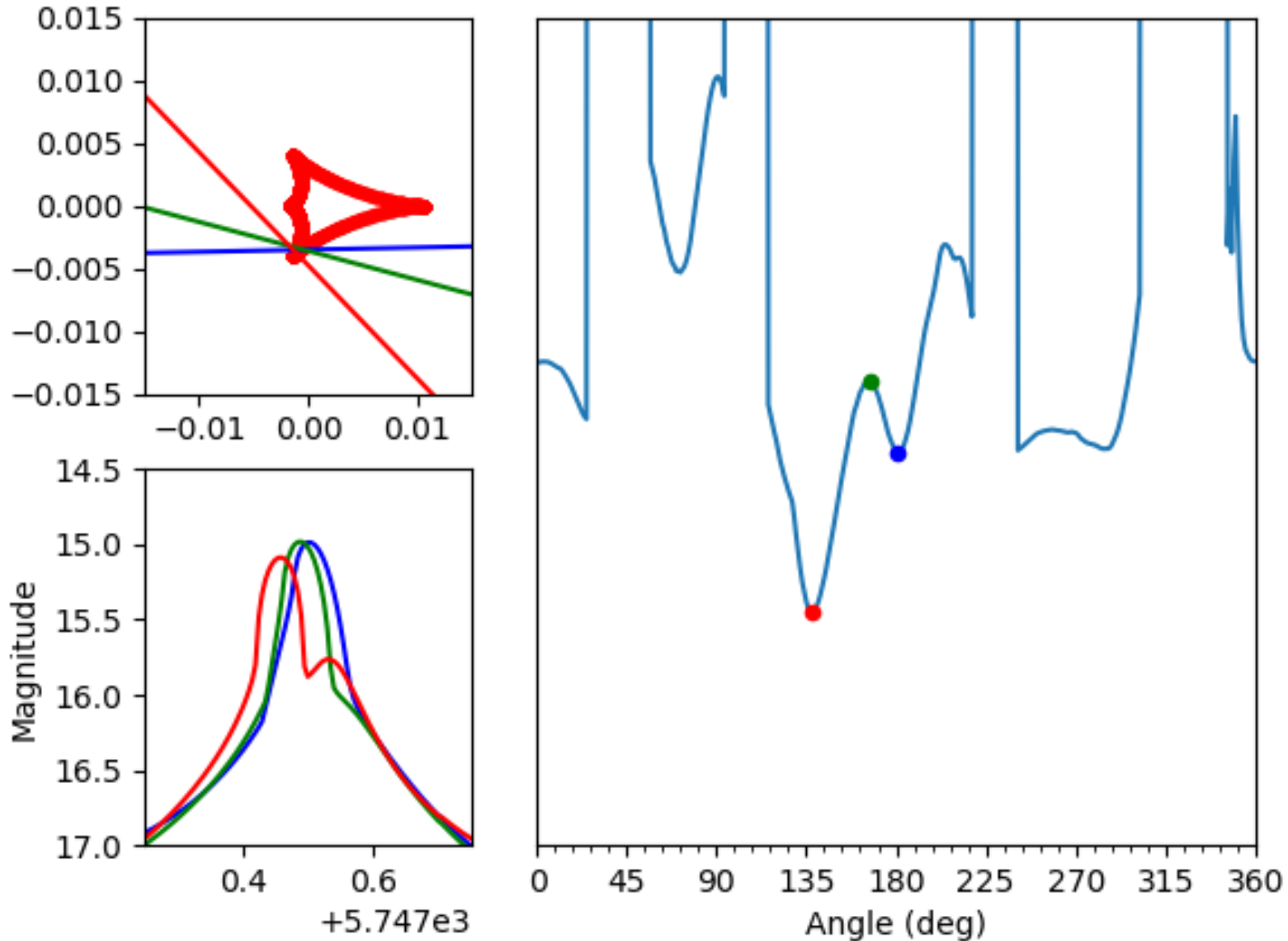
A simple downhill algorithm won't work.



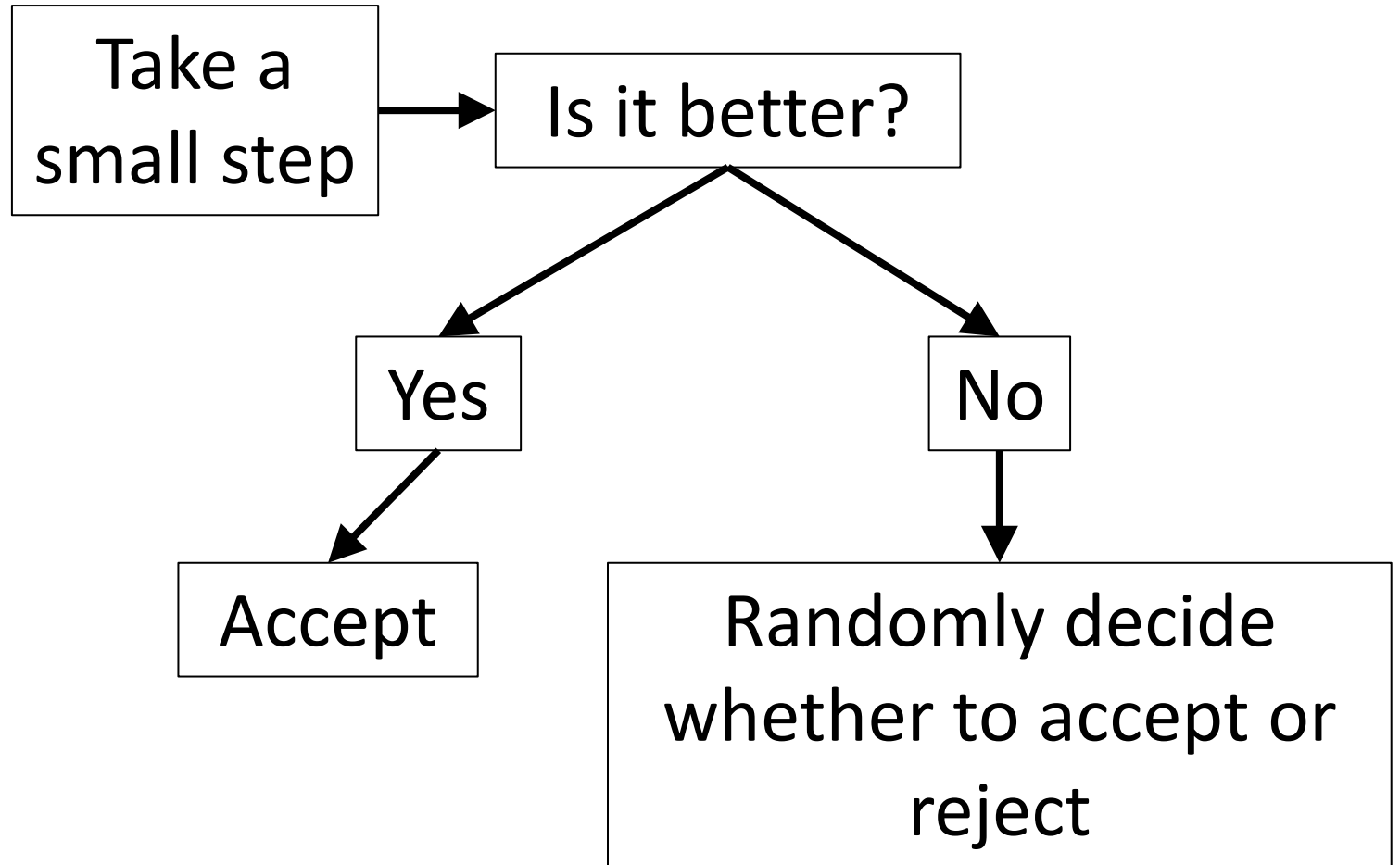
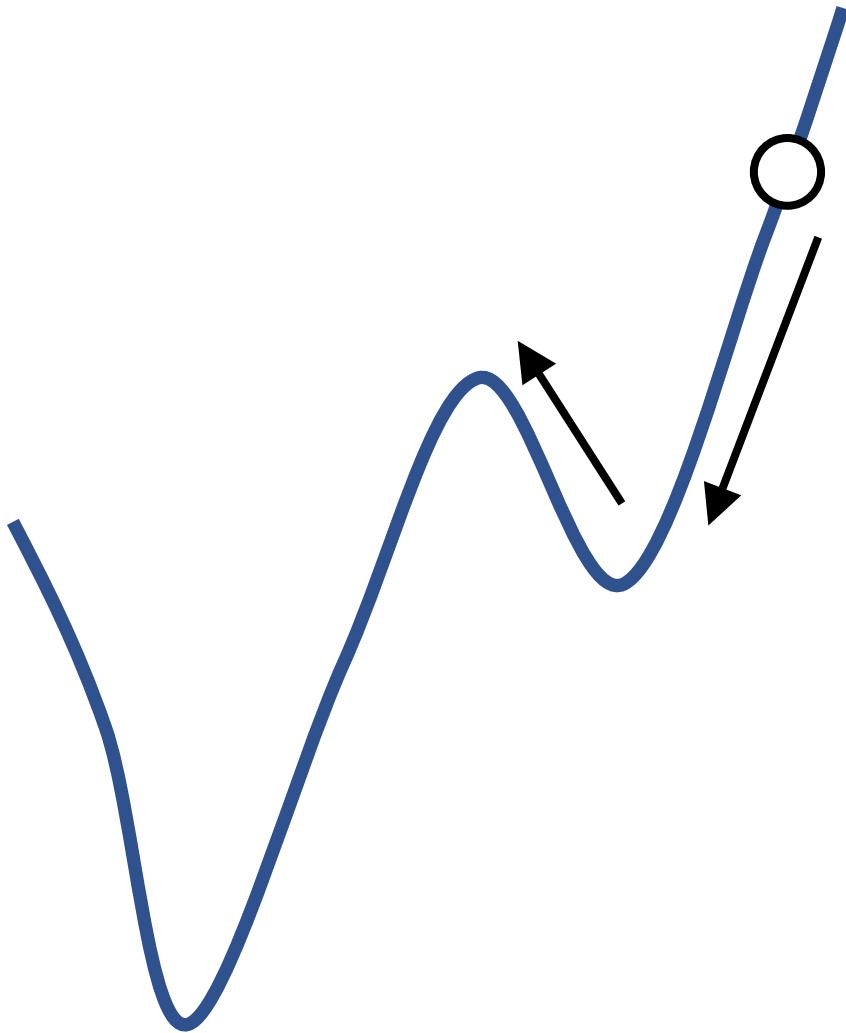
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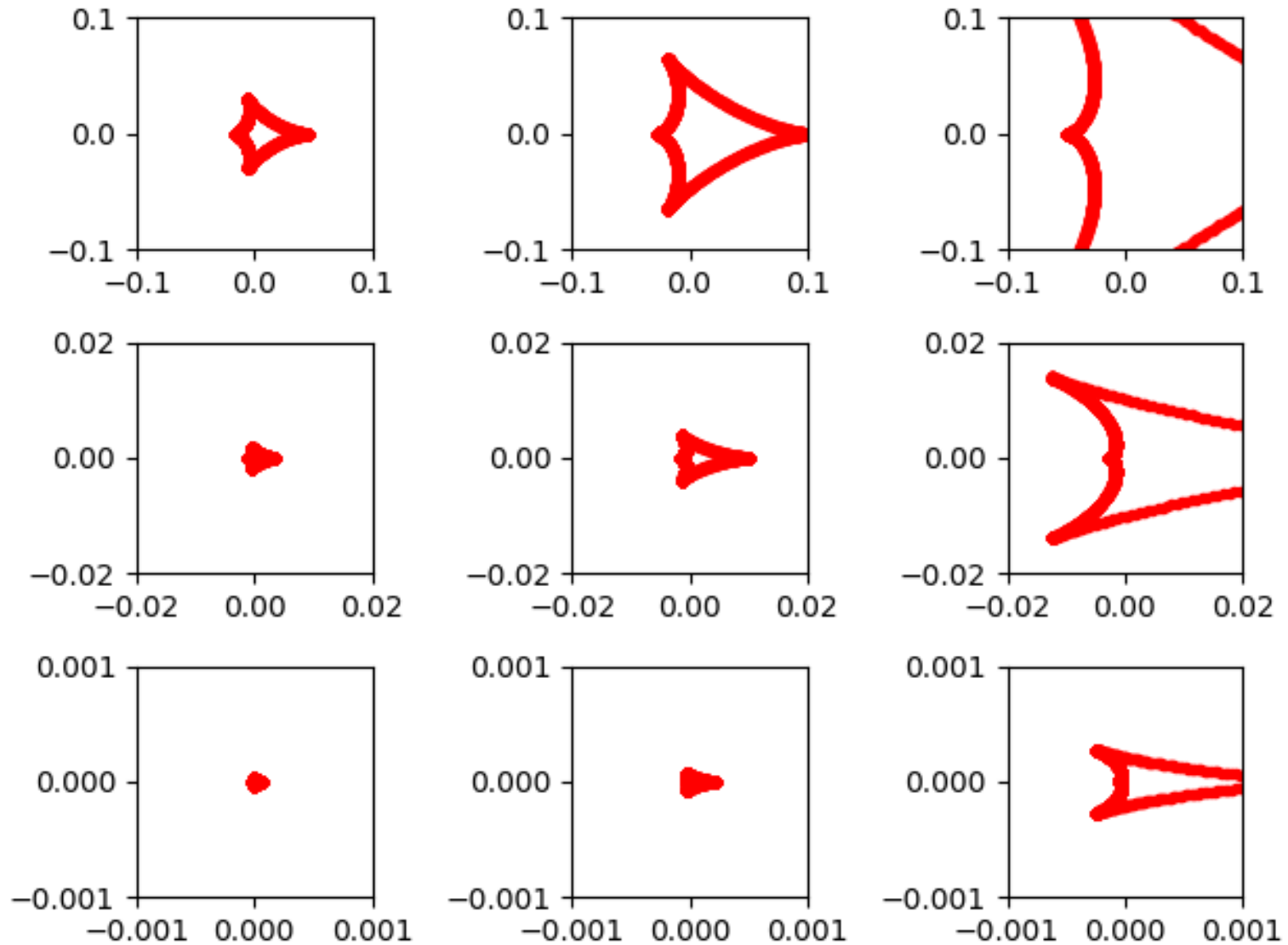
One Approach: Markov Chain Monte Carlo



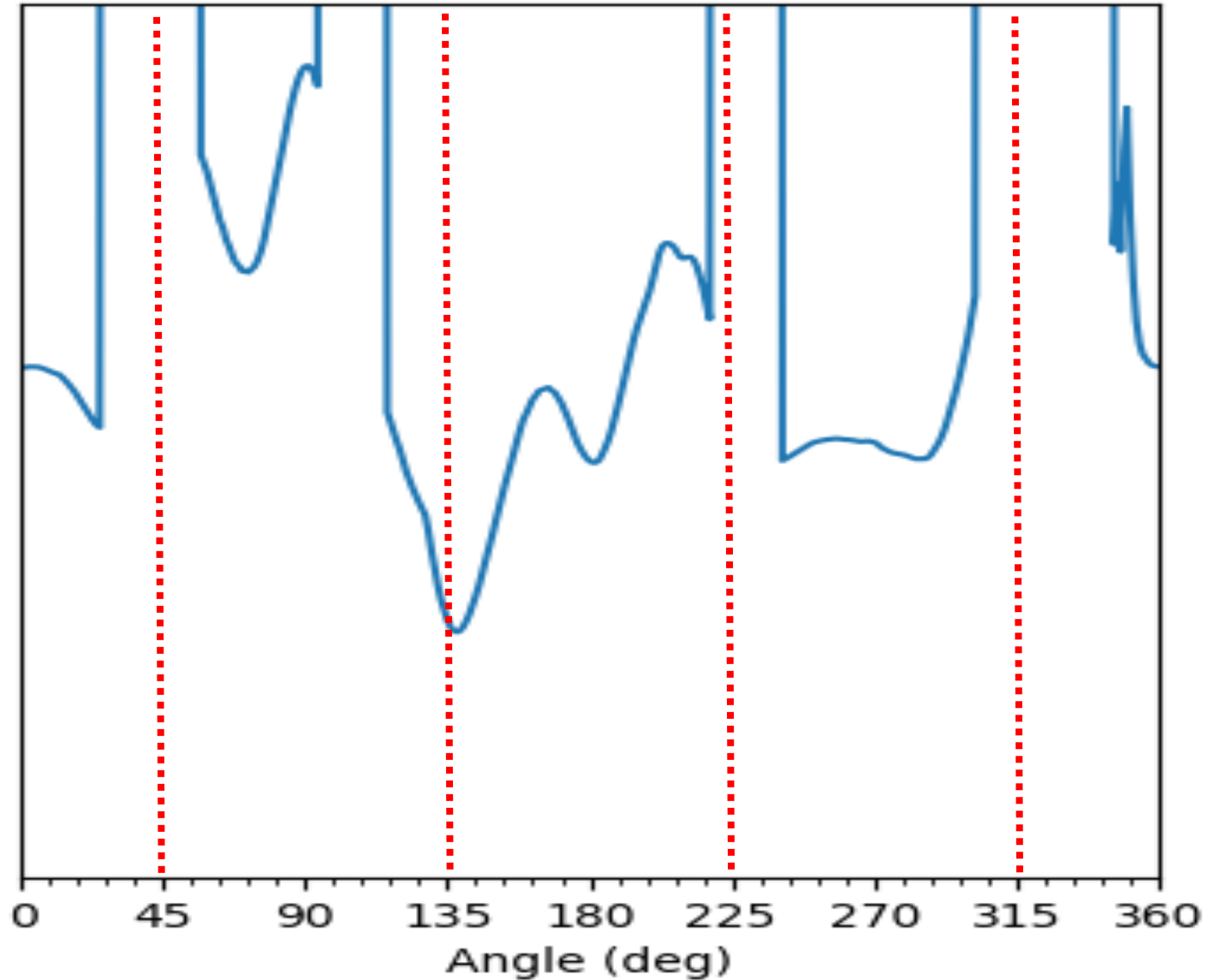
Markov Chain Monte Carlo (Random Walk)

- Advantage: It will eventually explore all of parameter space including multiple minima.
- Disadvantage: It may take the age of the universe to do it.

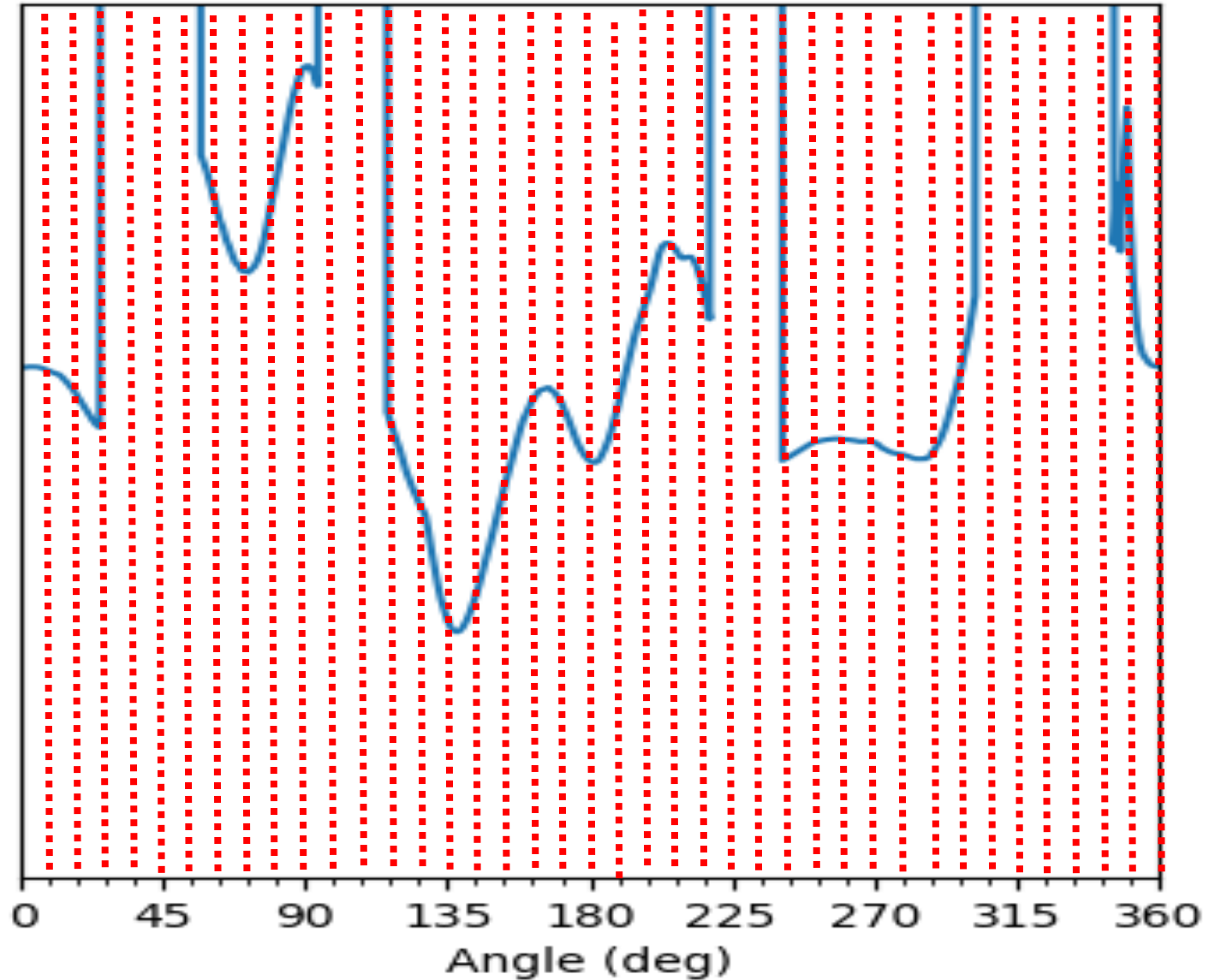
Hybrid Approach: Grid Search in s , q , angle



If the grid is too sparse, solutions are missed.



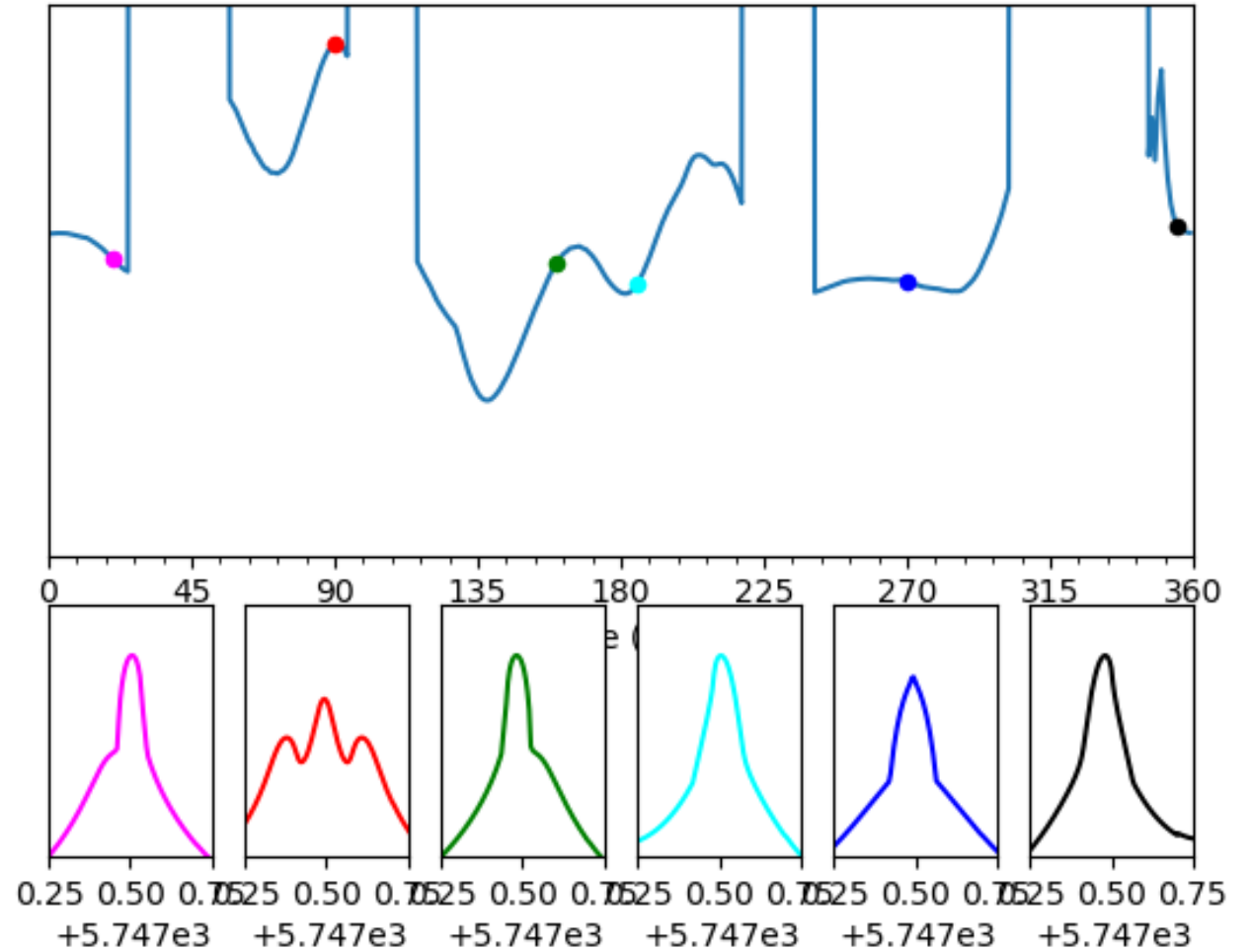
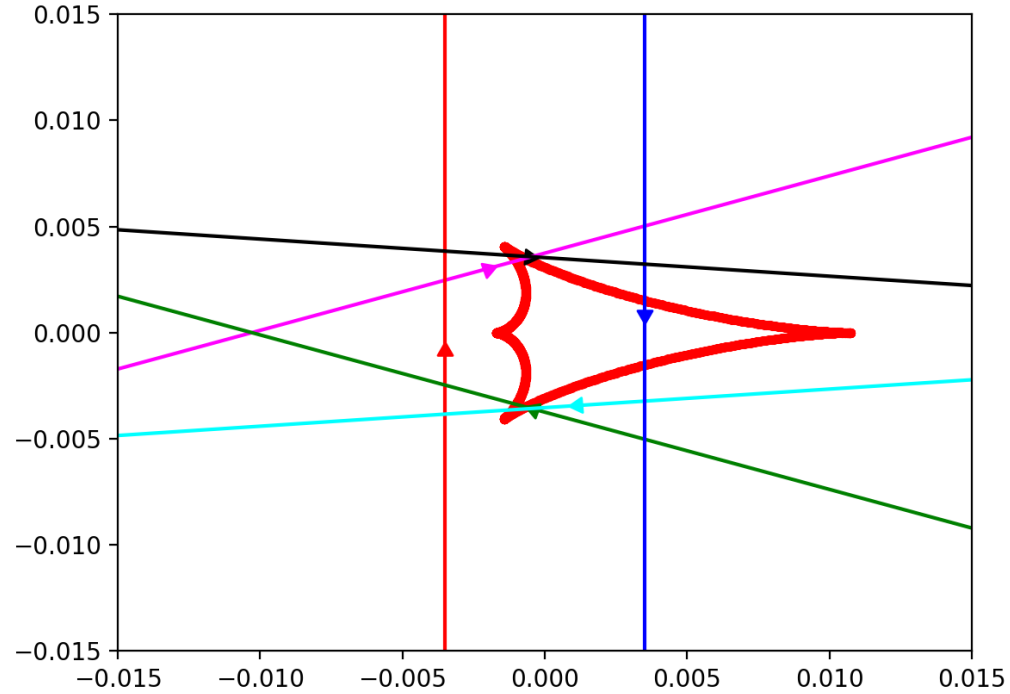
If the grid is too dense, it takes too long (inefficient).



MCMC + Grid Search

- Advantage: It can improve efficiency of MCMC alone (fewer parameters)
- Advantage: Good for a blind, systematic search
- Disadvantage: There is an art to selecting the grid size

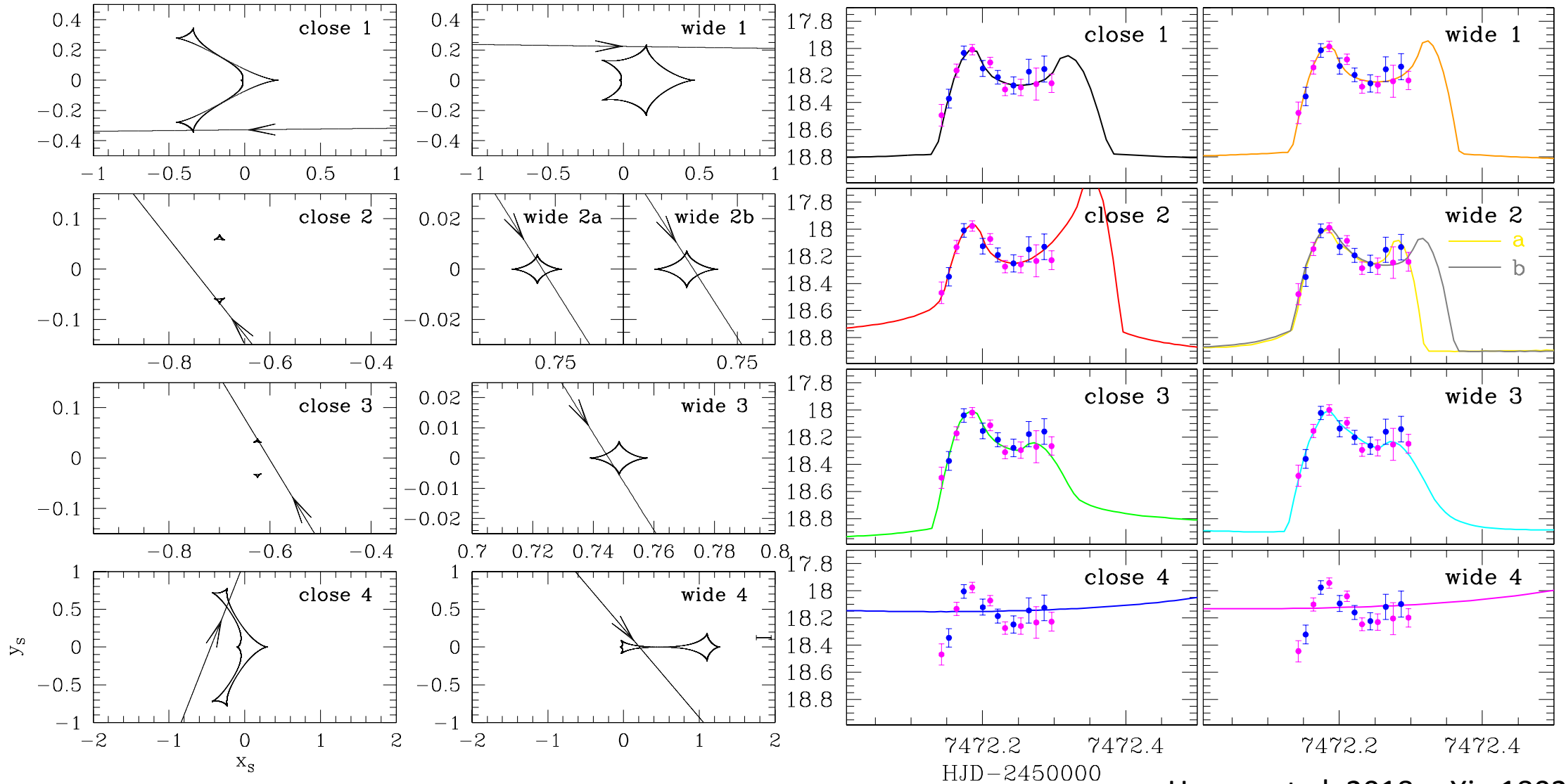
Alternative Approach: Light Curve Library



Light Curve Library

- Advantage: Optimal sampling of parameter space
 - Less likely to miss solutions
 - Can be more efficient
- Disadvantage: Significant effort to set up. Requires a deep knowledge of caustic structures.

Partial data = More degeneracies



Summary

- The magnification pattern + source trajectory determines the light curve
- Symmetries/similarities in the magnification make it hard to find the true solution
- Methods for searching for the correct model:
 - MCMC + grid search
 - Light curve library

Part III: Resources

Resources: Modeling

- MulensModel: Python, generates models, includes examples
<https://github.com/rpoleski/MulensModel/releases/tag/v1.4.0>
- pyLIMA: Python, fits models to data, includes examples
<https://github.com/ebachelet/pyLIMA/tree/master/pyLIMA>
- VBBL (Valerio Bozza Binary Lensing): C++, generates models
<http://www.fisica.unisa.it/GravitationAstrophysics/VBBinaryLensing.htm>

Resources: Public Data

- Exoplanet Archive: Data from Published Planets
<https://exoplanetarchive.ipac.caltech.edu/>
- Korea Microlensing Telescope Network: Data for Microlensing Events from 2015 and K2C9 data are fully public.
<http://kmtnet.kasi.re.kr/ulens/>
- UKIRT Microlensing Survey: Data for all stars for 2015—2017 campaigns
<https://exoplanetarchive.ipac.caltech.edu/docs/UKIRTMission.html>

Microlensing Data Challenge: “Solve” 293 WFIRST-like lightcurves

<http://microlensing-source.org/data-challenge-guidelines/>

