

Pale blue  
dot

# The promise of transits



Josh Winn  
Massachusetts Institute of Technology

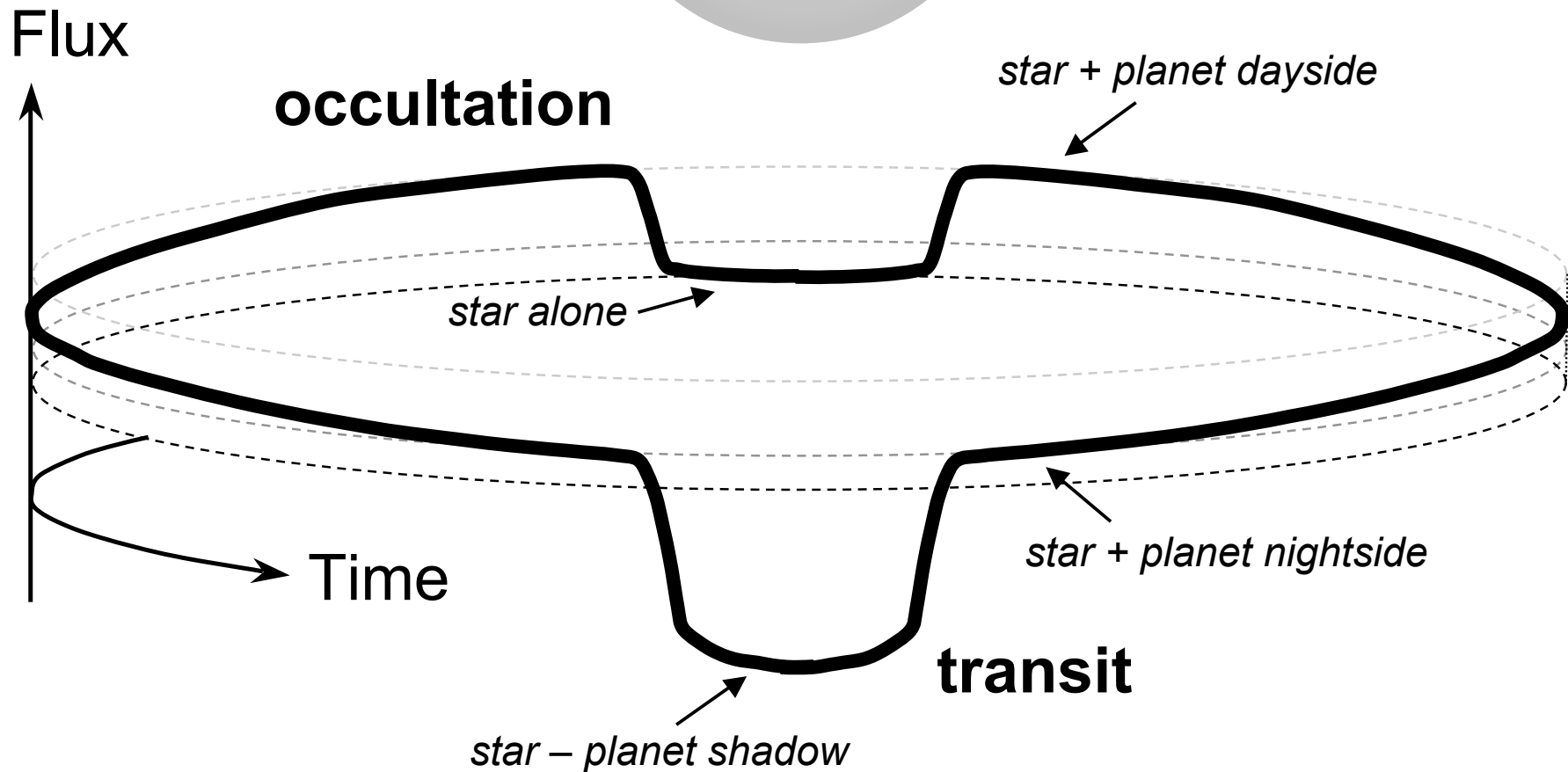
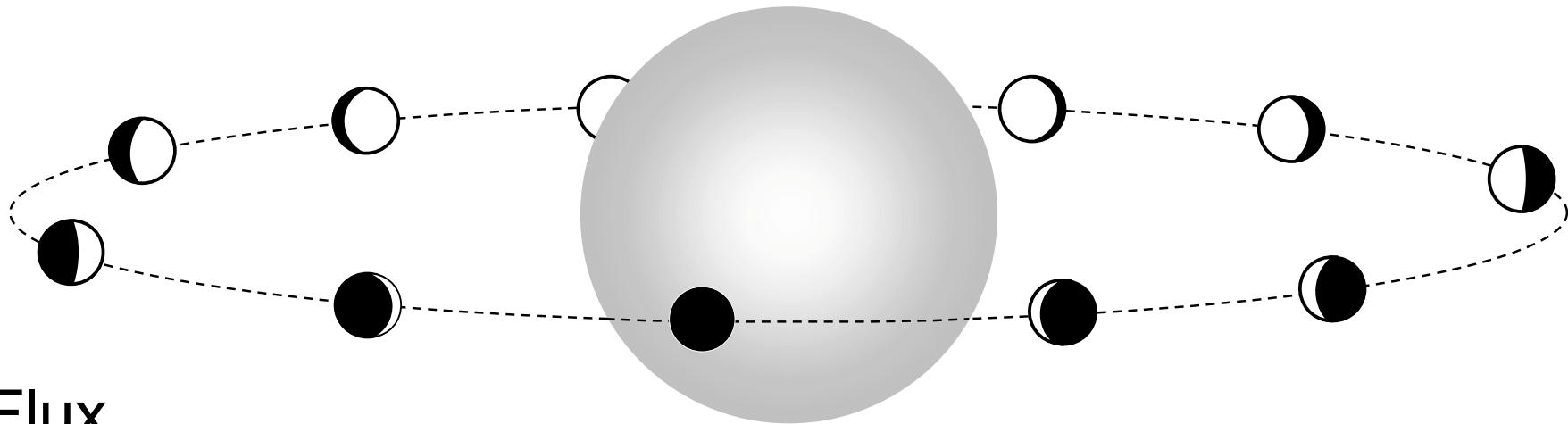


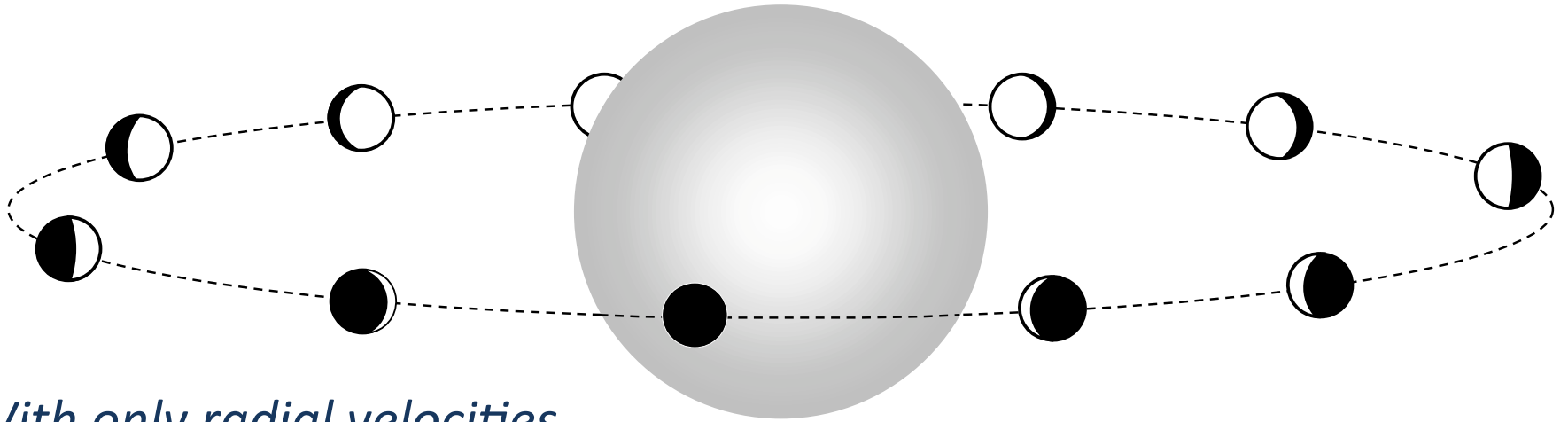
Pitch black  
dot

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*With only radial velocities...*

***Orbital properties***

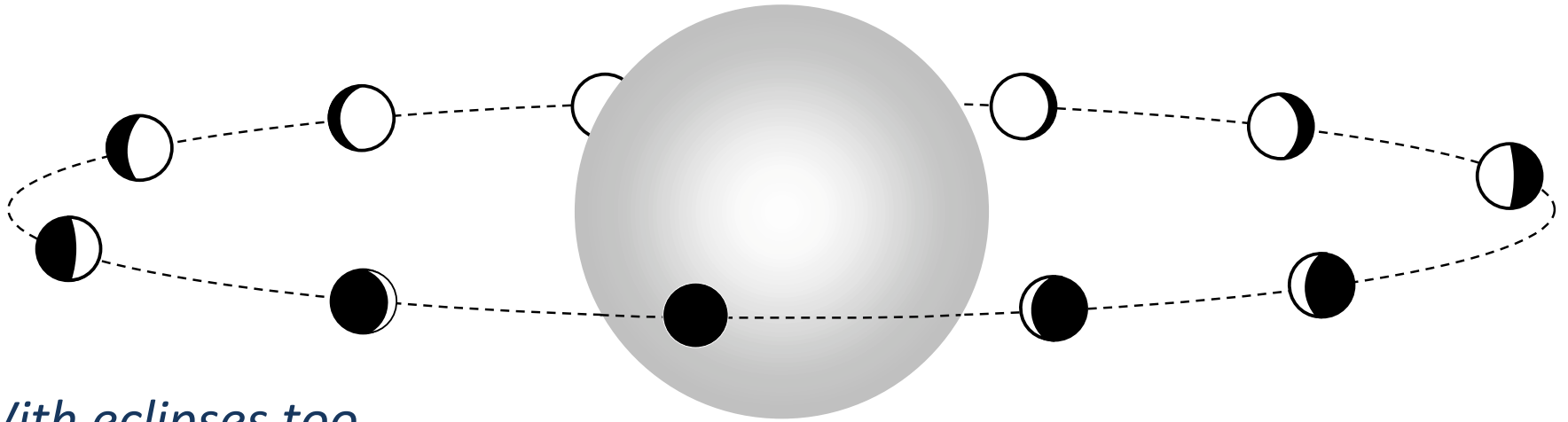
Period

Eccentricity

***Planet bulk properties***

Minimum mass





*With eclipses too...*

***Orbital properties***

Period  
Eccentricity

***Planet bulk properties***

Mass  
Radius

***Planet atmospheric properties***

Emission spectrum  
Transmission spectrum  
Reflectance spectrum  
Phase function

***Stellar properties***

Limb darkening function  
Gravity darkening function  
Spots and plages  
Obliquity

***Multiple-planet systems***

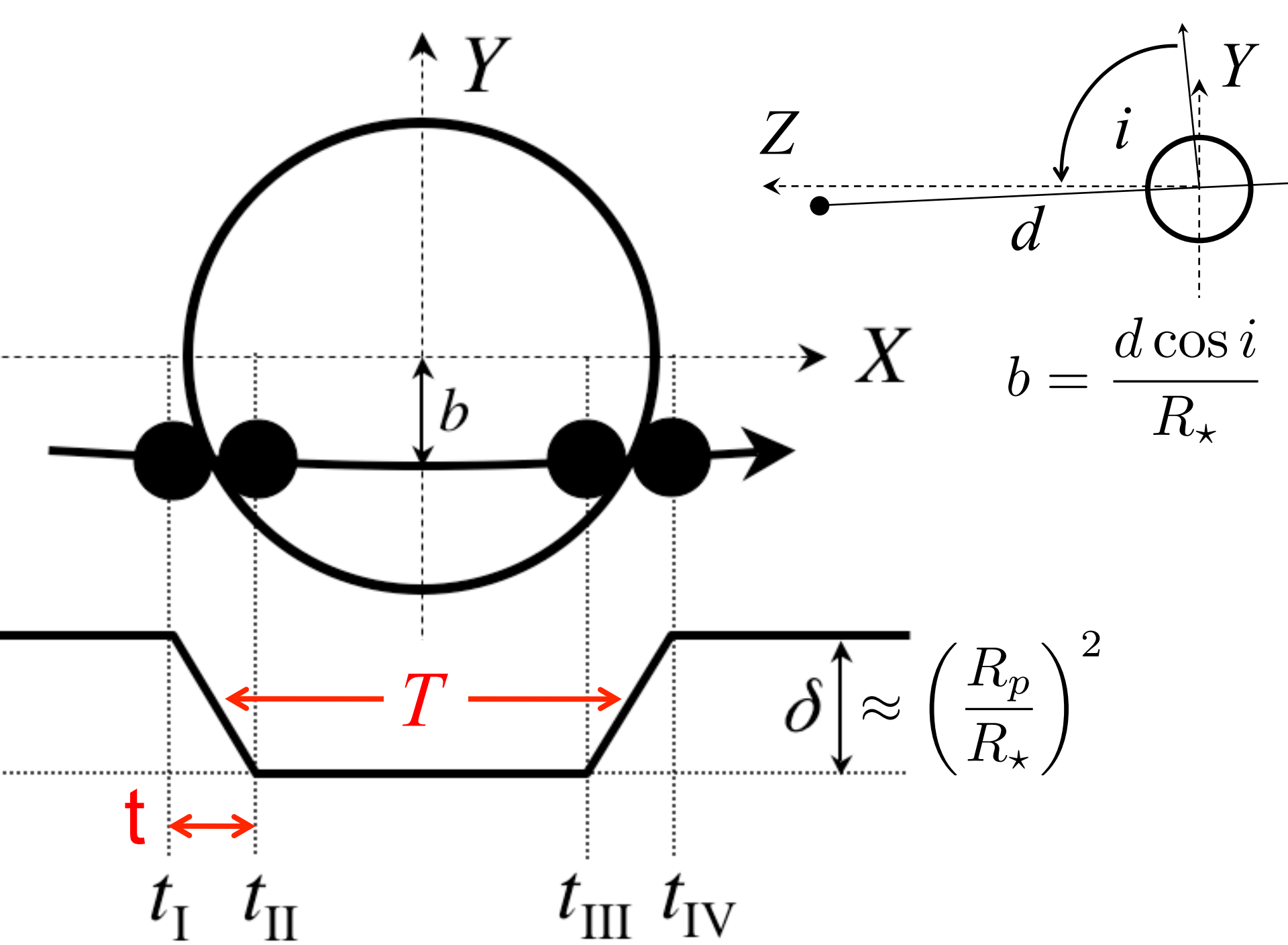
Dynamical masses  
Mutual inclinations  
Resonances and chaos

***Multiple-star systems***

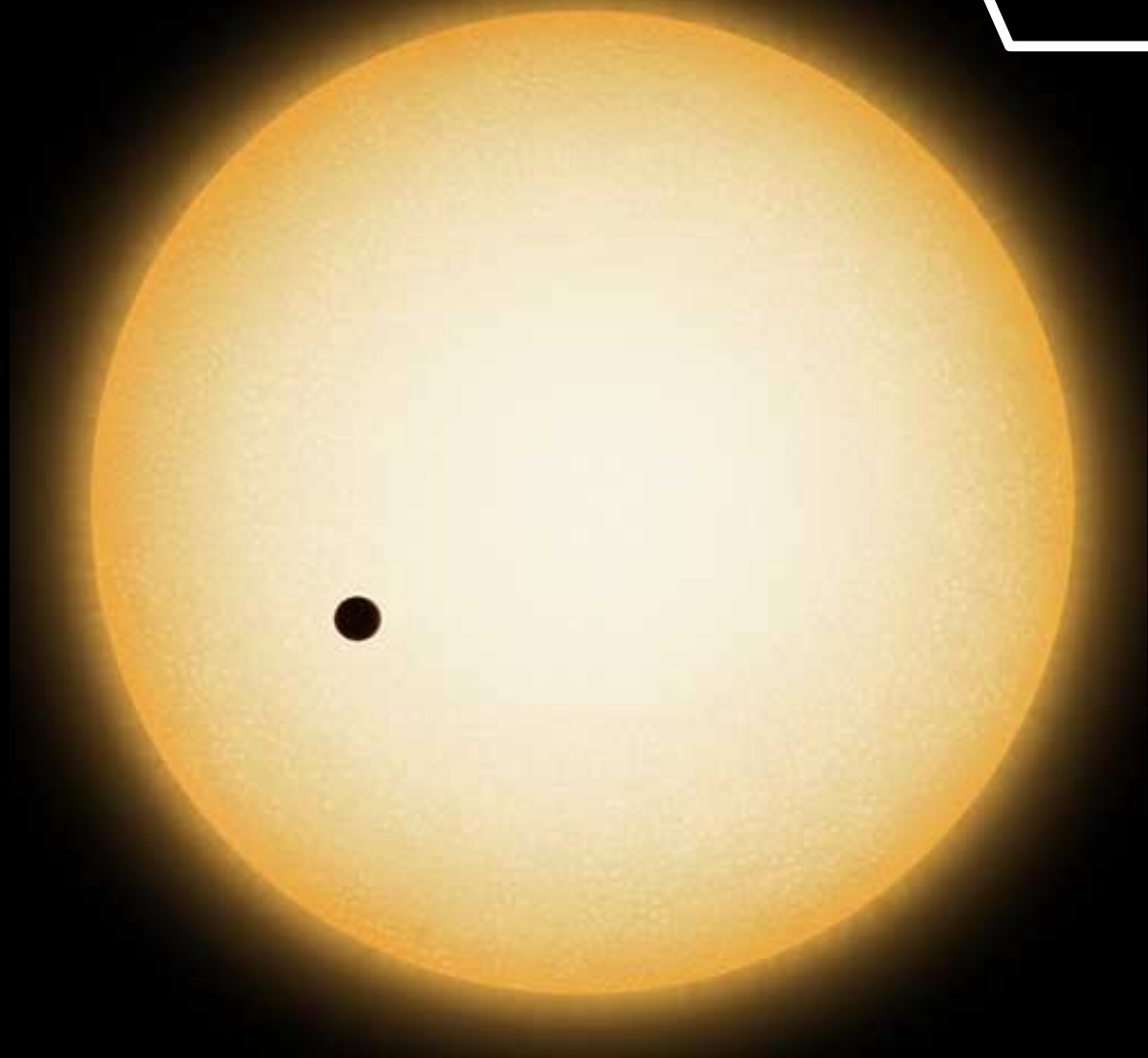
Circumbinary planets

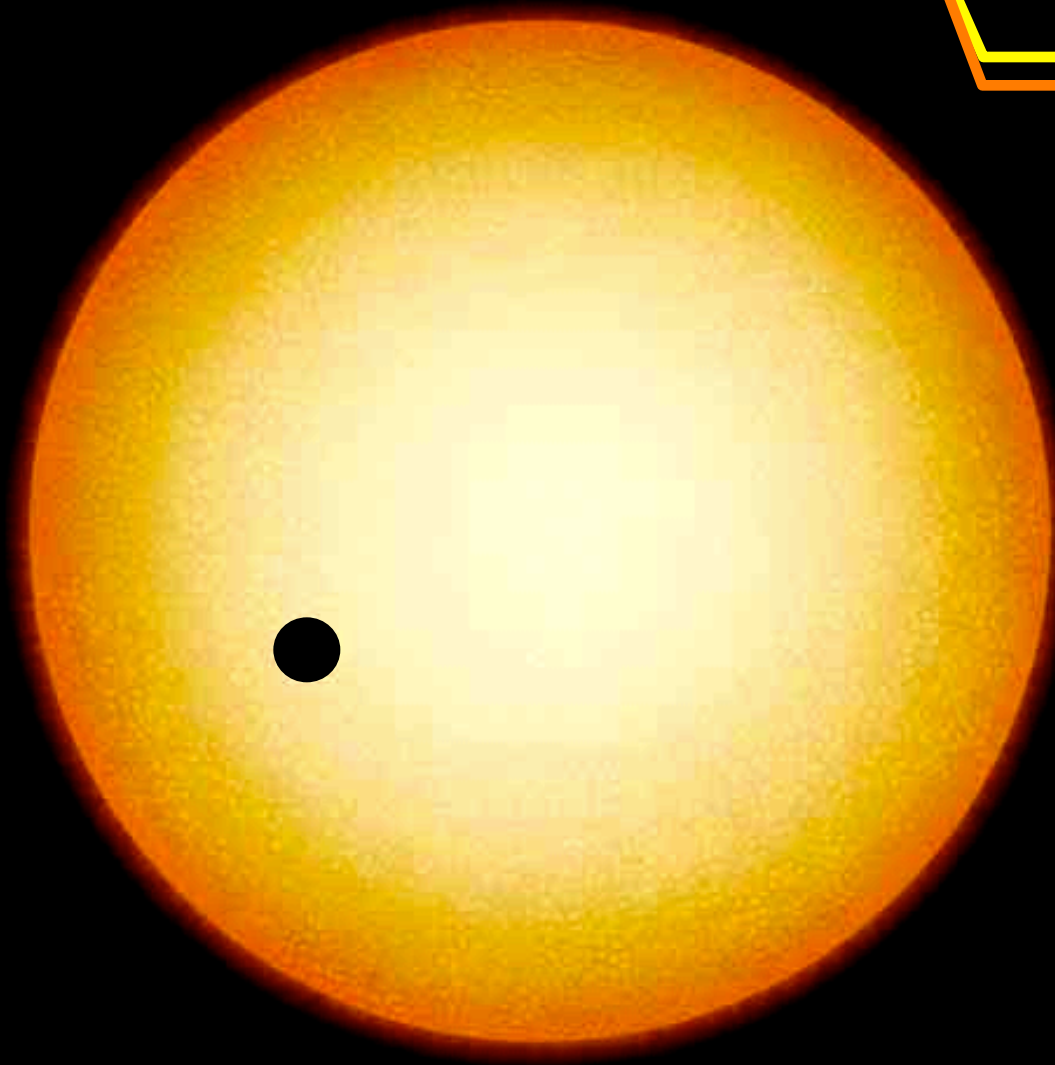
***For the future***

Moons and Trojans  
Oblateness and rings  
Apsidal motion constant  
Relativistic precession  
Orbital decay  
Applegate effect  
Planetary wind speed  
Yarkovsky effect  
Planetary magnetic field  
Planetary aurorae  
Artificial planet-sized objects



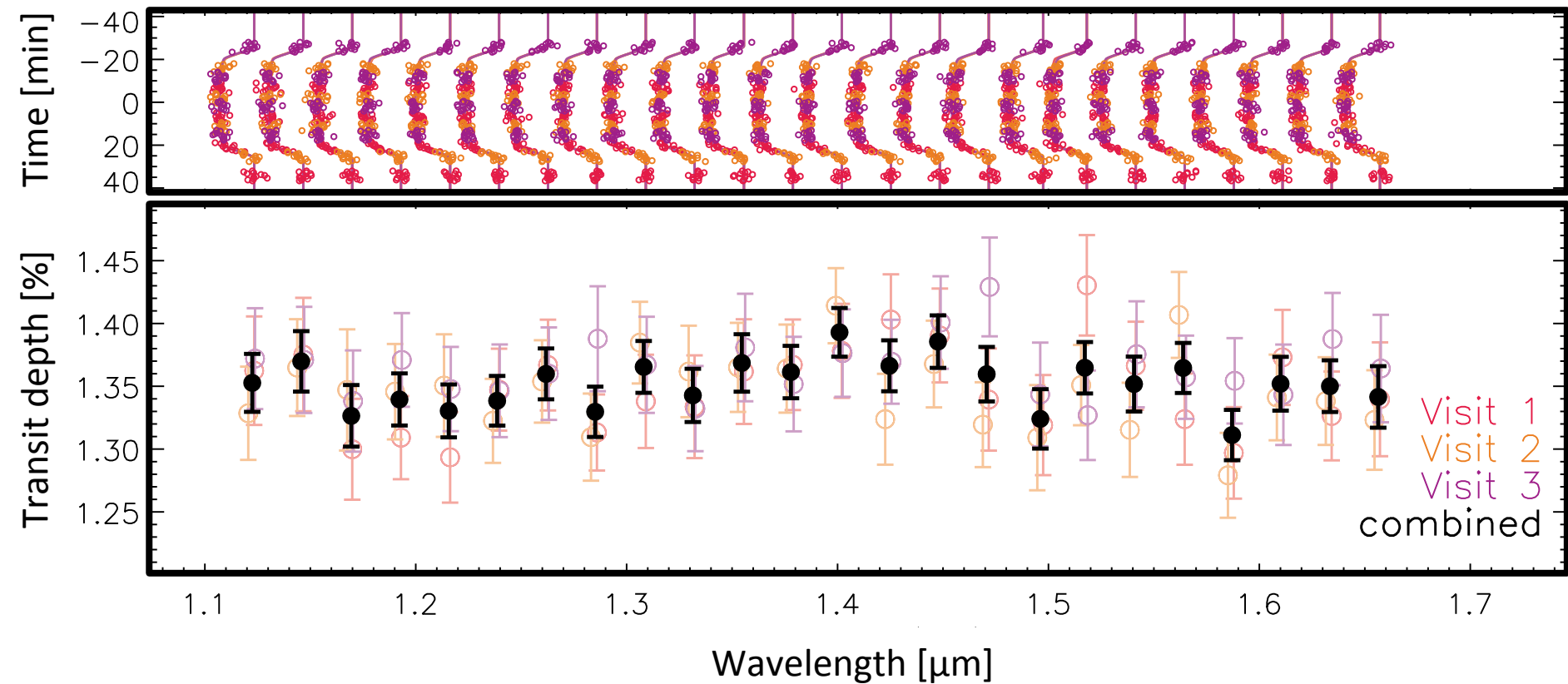




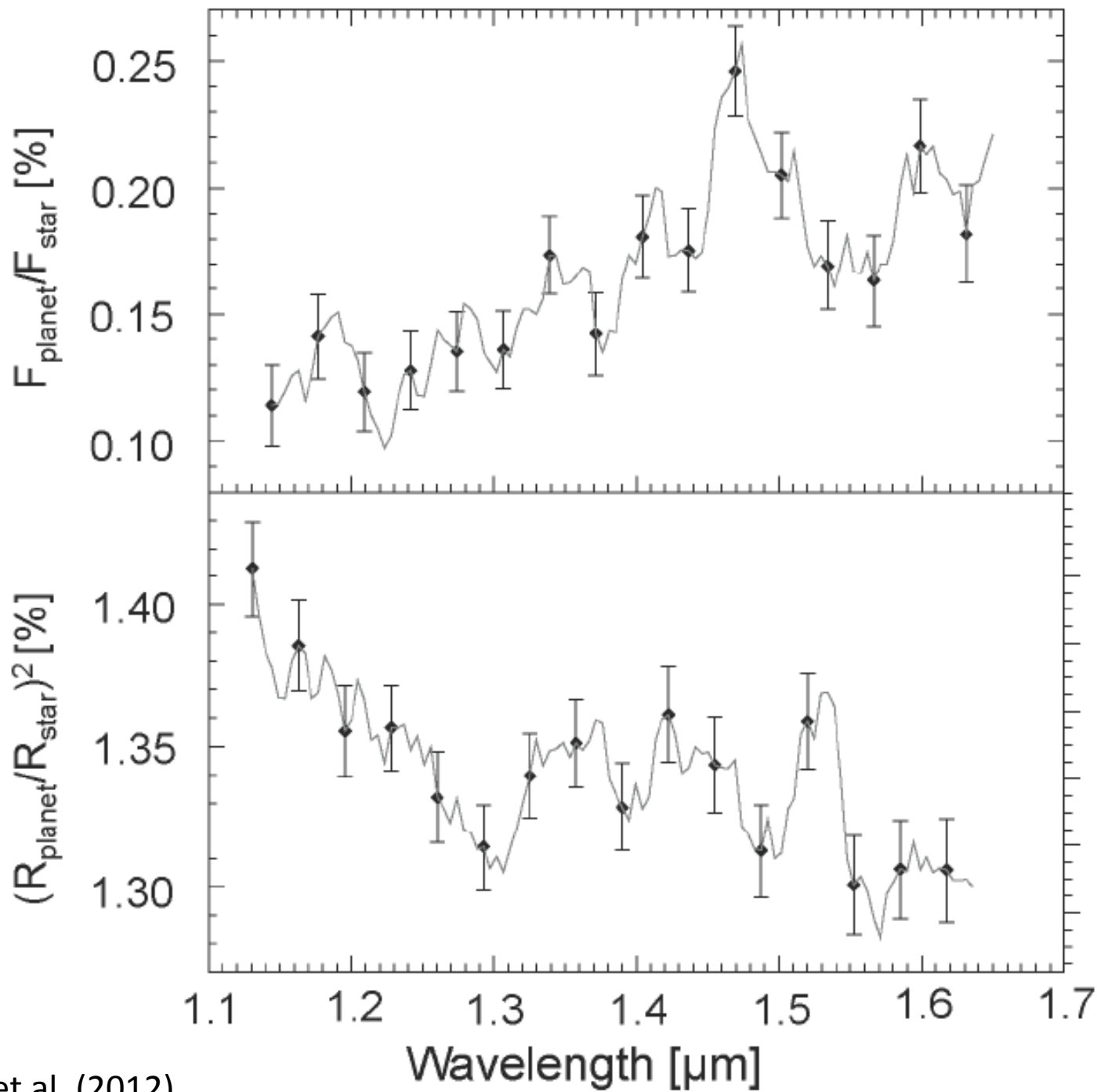




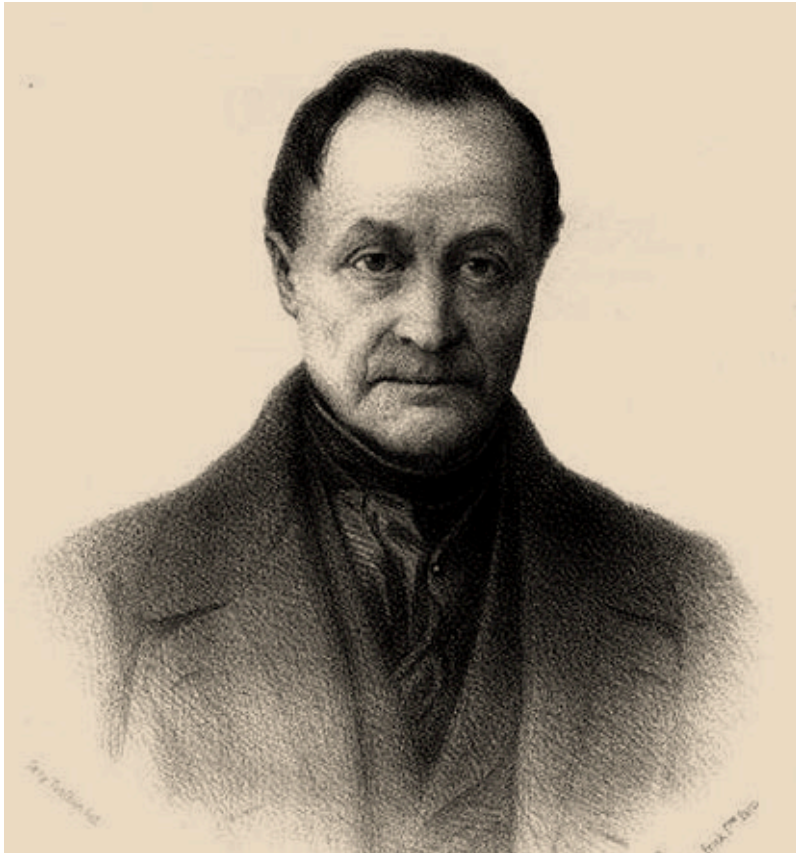
# Transmission spectroscopy



# Transit Occultation







Auguste Comte,  
*Positive Philosophy* (1853)

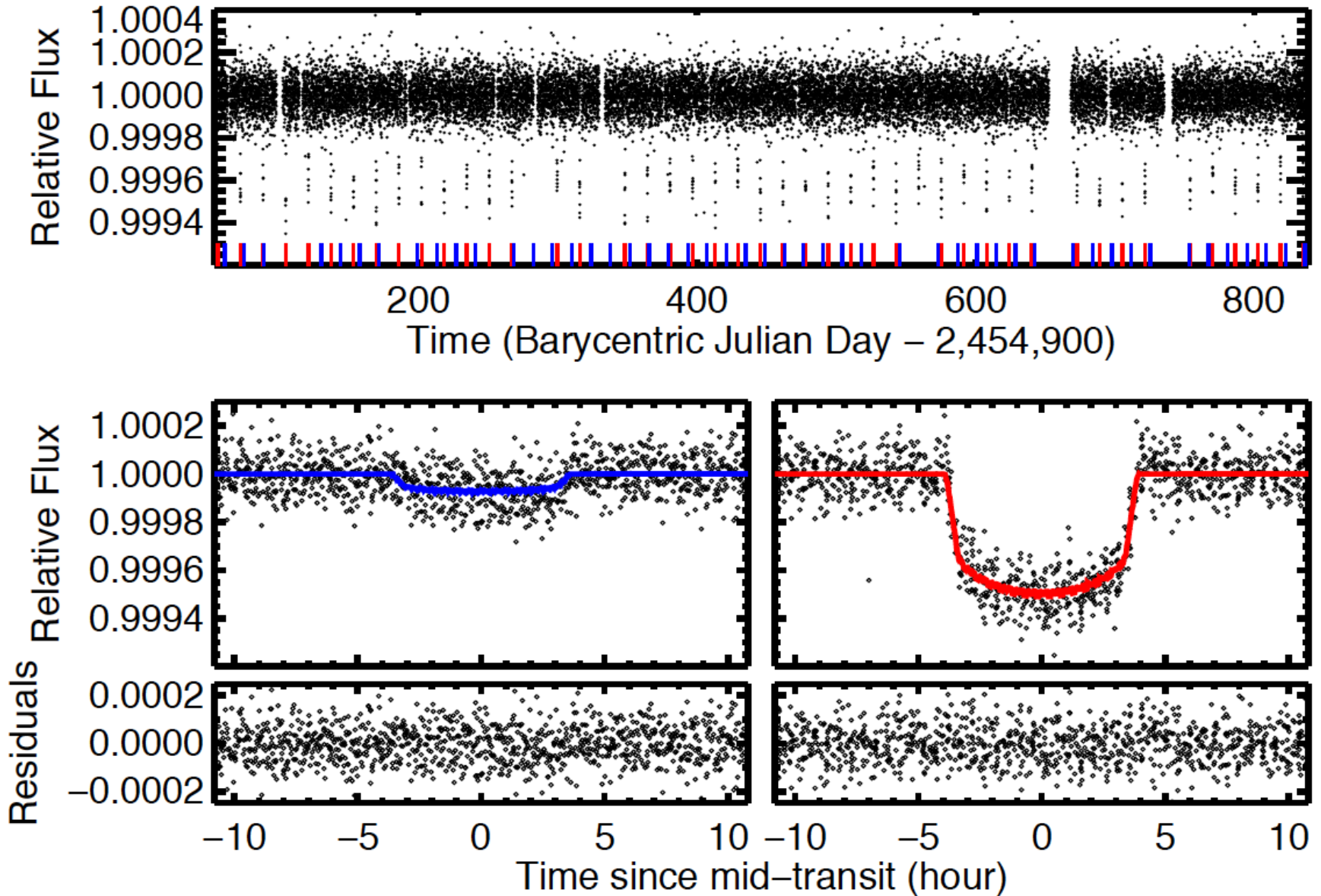
We see how we may determine their forms, their distances, their bulk, and their motions, but **we can never know anything of their chemical or mineralogical structure**; and, much less, that of organized beings living on their surface. ... All physical, chemical, physiological, and social researches are out of the question in regard to the planets.



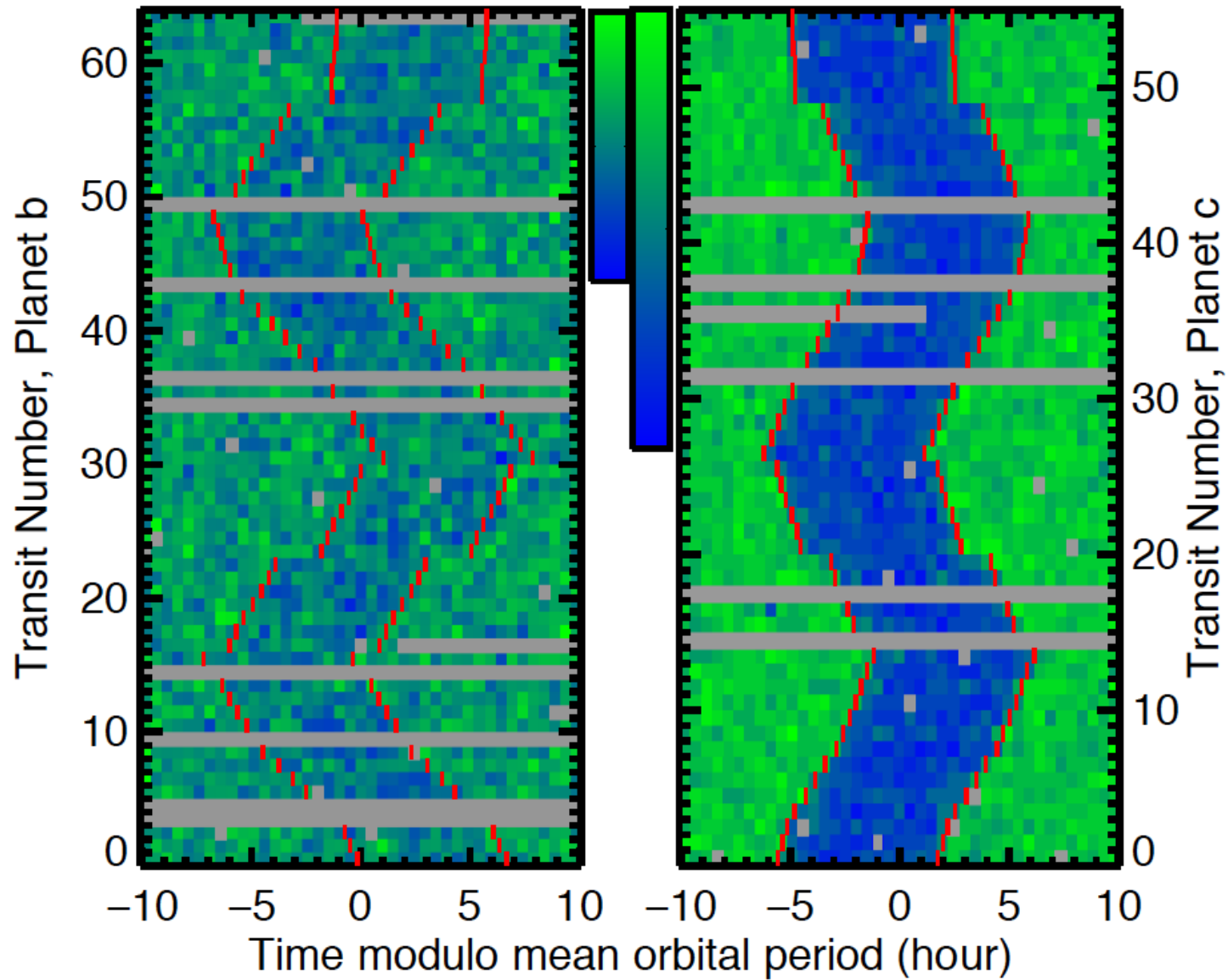
Josh Winn,  
*Exoplanets* (2010)

Ideally one would like to know the mass in kilograms, and the radius in kilometers, to allow for physical modeling and comparisons with solar system planets. **With only a transit light curve, this is impossible.** The light curve by itself reveals the planet-to-star radius ratio but not the planetary radius, and says nothing about the planetary mass.

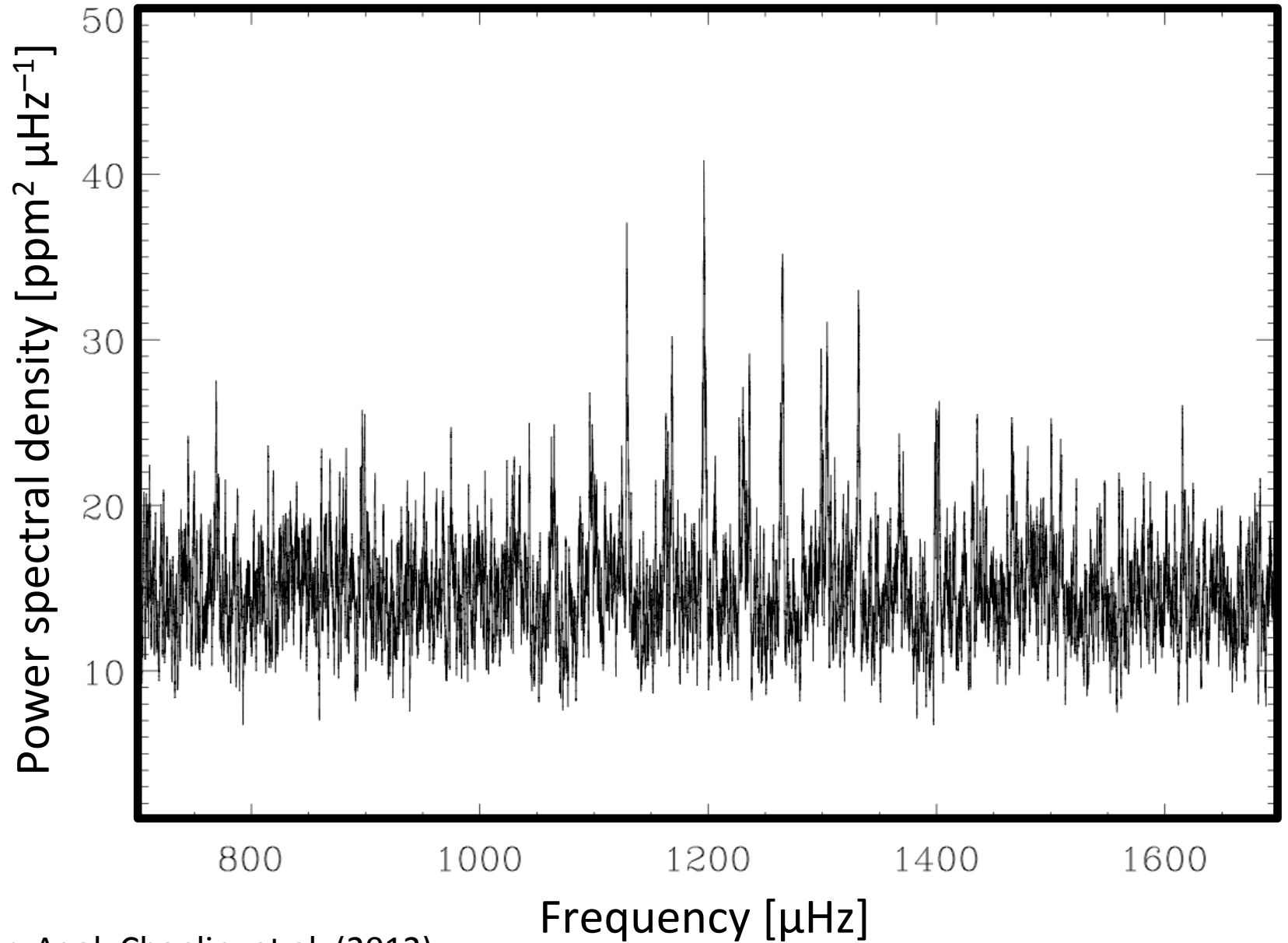
# Photodynamics



# Photodynamics

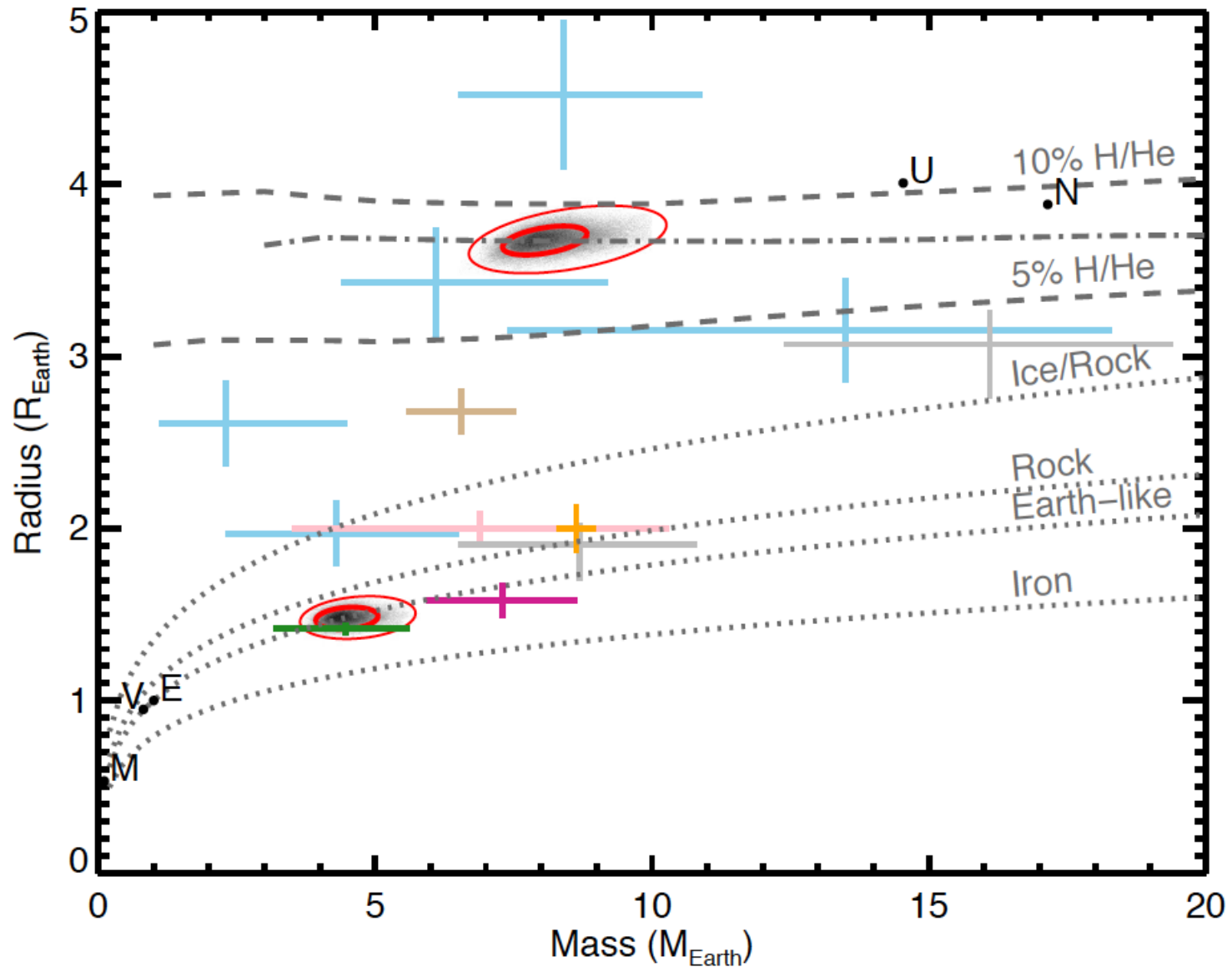


# Asteroseismology

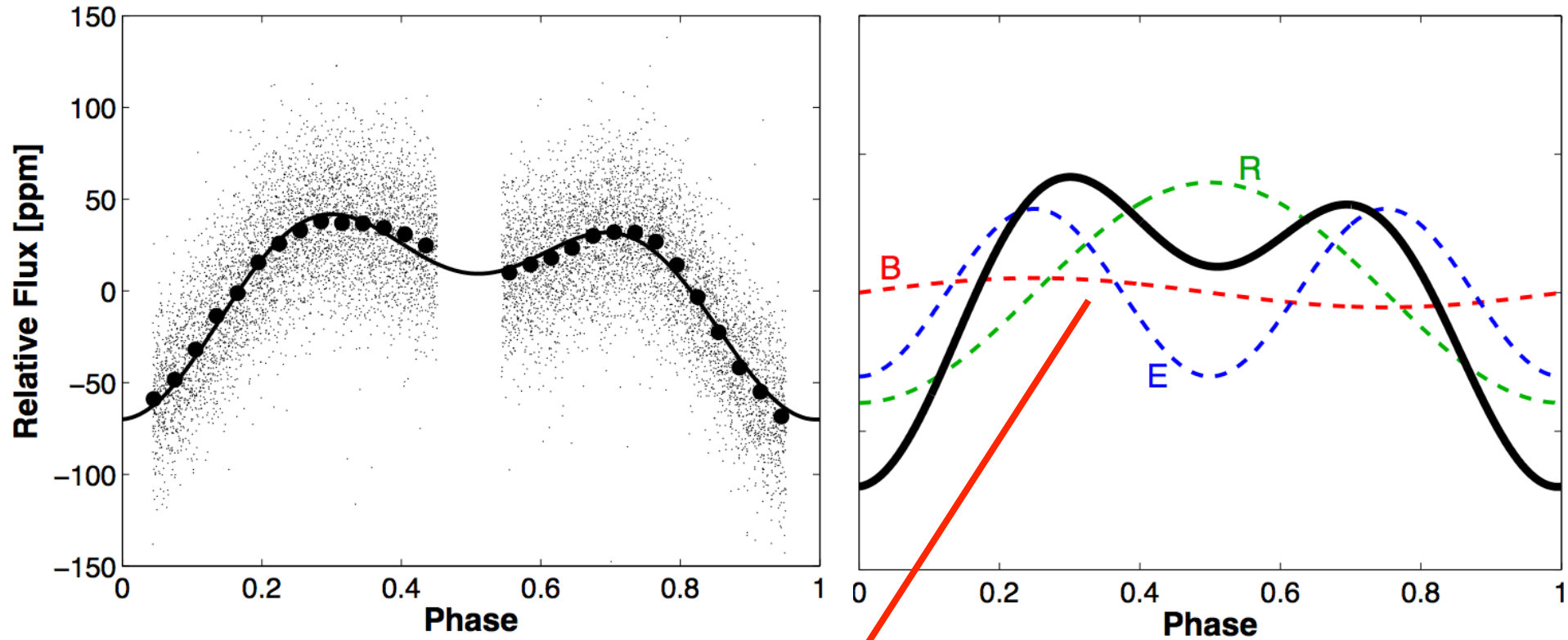




# Photodynamics + Asteroseismology

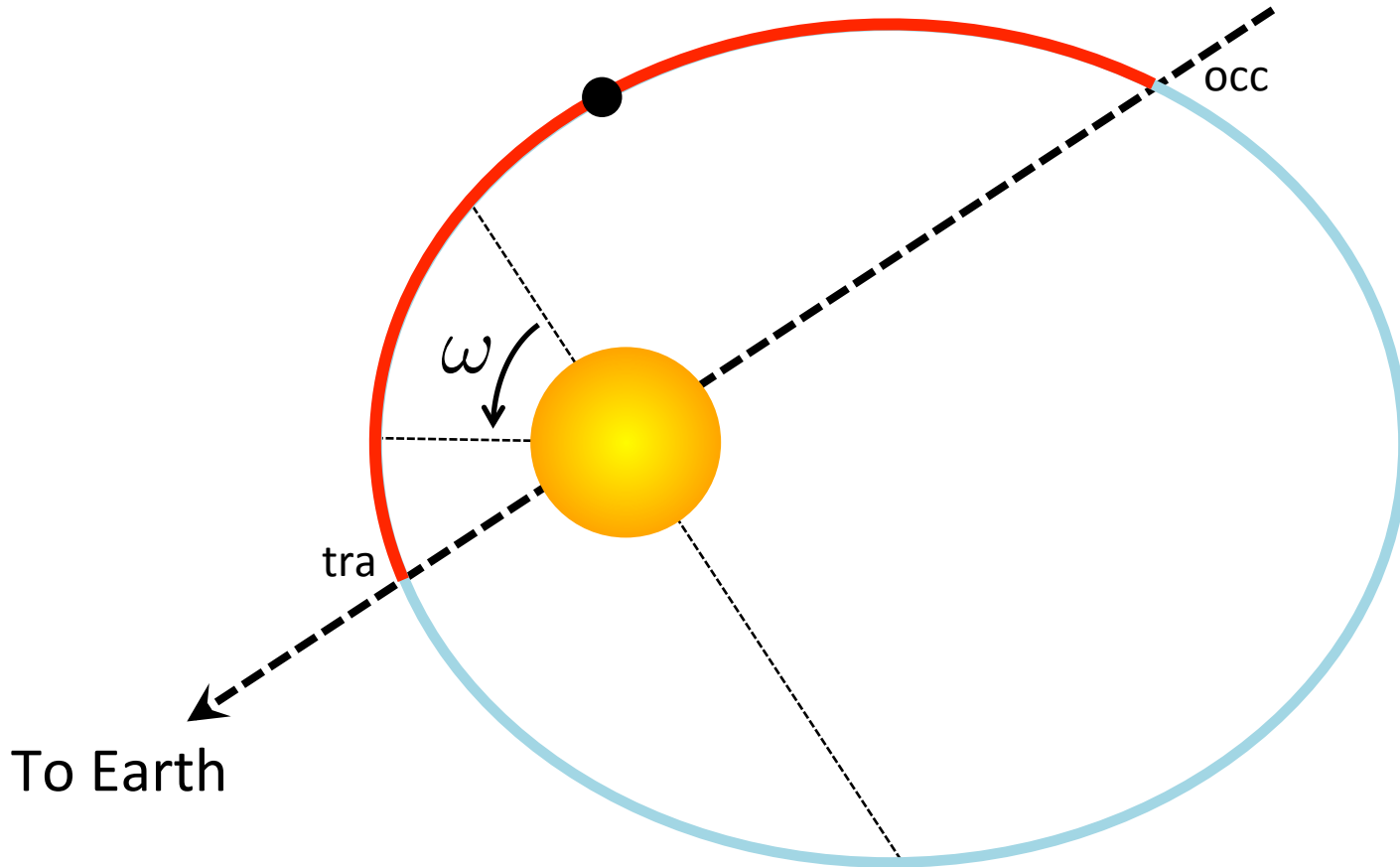


# Doppler boosting



$$A_{\text{boost}} \propto \frac{K_{\text{RV}}}{c}$$

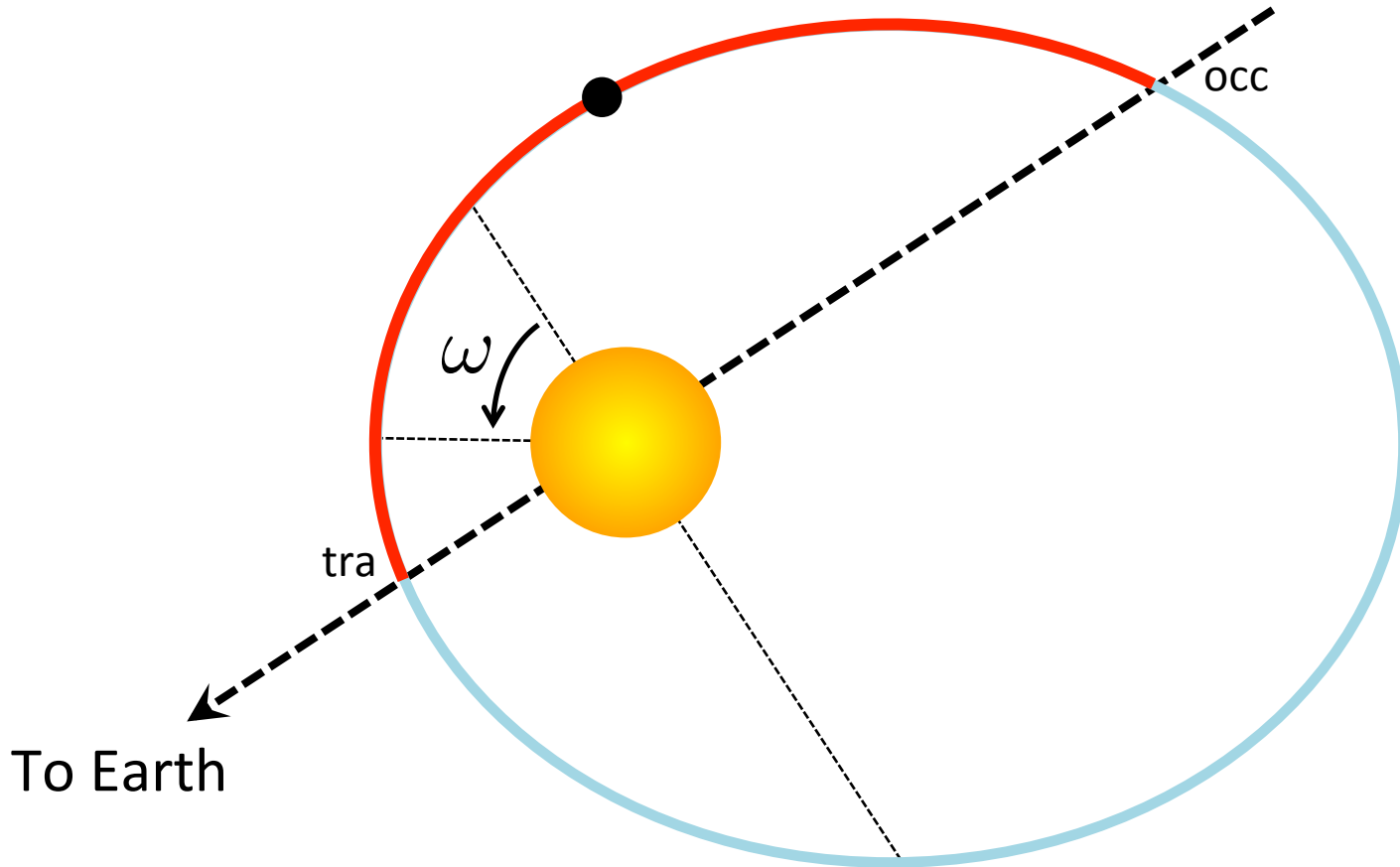
# Eccentricity



$$\Delta t_c \approx \frac{P}{2} \left[ 1 + \frac{4}{\pi} e \cos \omega \right]$$

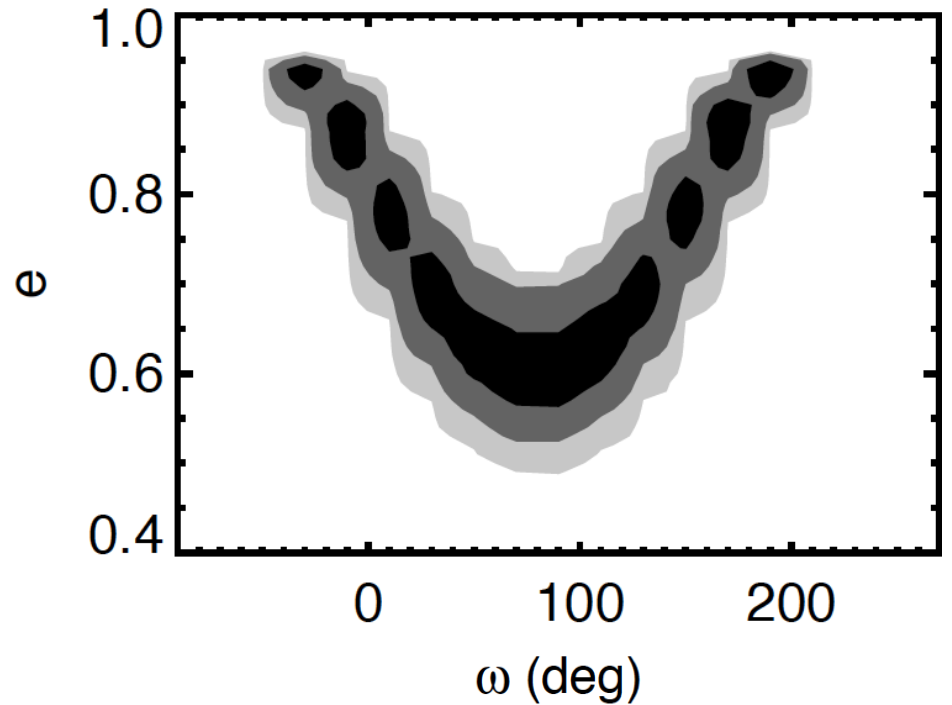
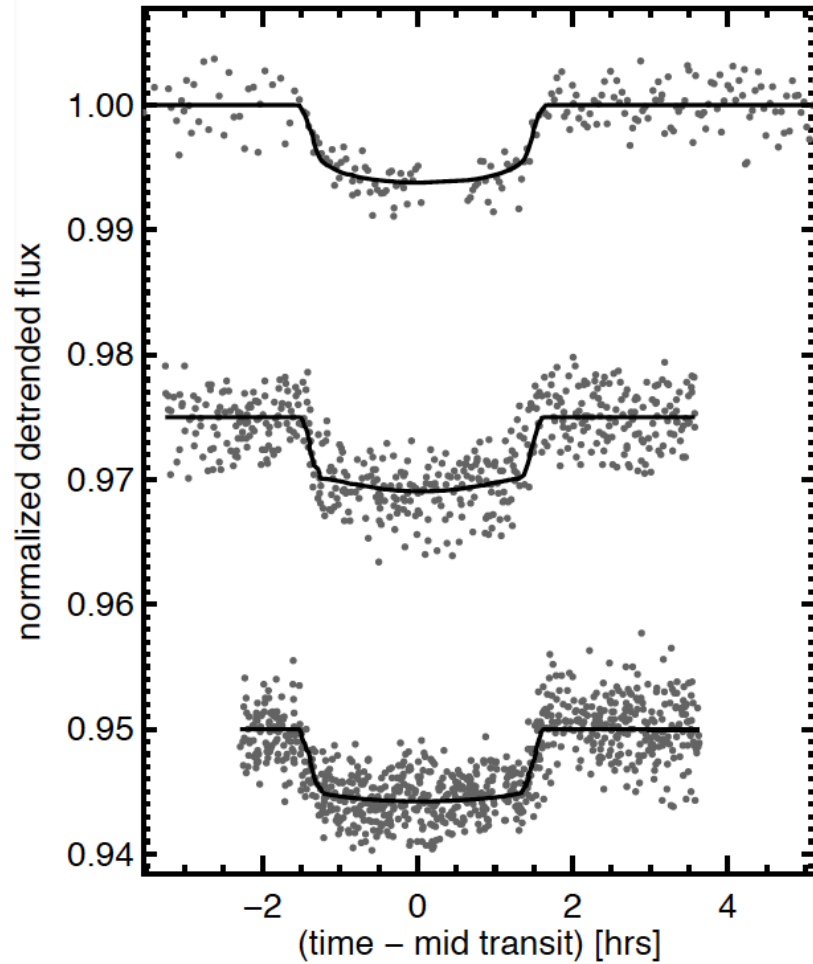
$$\frac{T_{\text{occ}}}{T_{\text{tra}}} \approx \frac{1 + e \sin \omega}{1 - e \sin \omega}$$

# Eccentricity

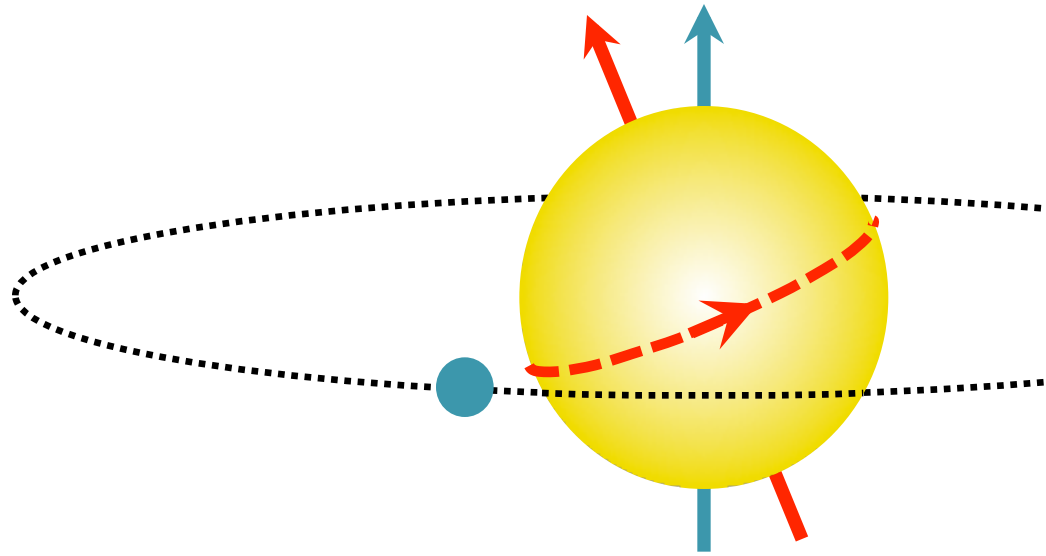


$$T_{\text{tra}} \propto \rho_{\star}^{-1/3} \frac{\sqrt{1 - e^2}}{1 + e \sin \omega}$$

# The photoeccentric effect



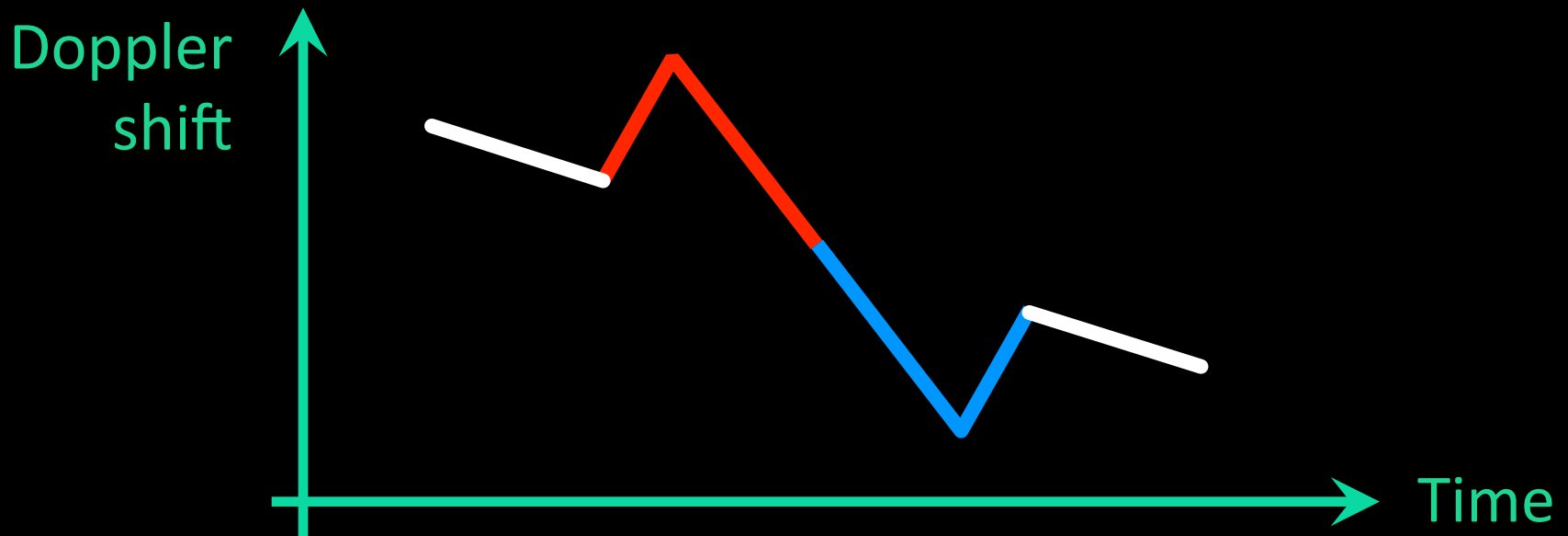
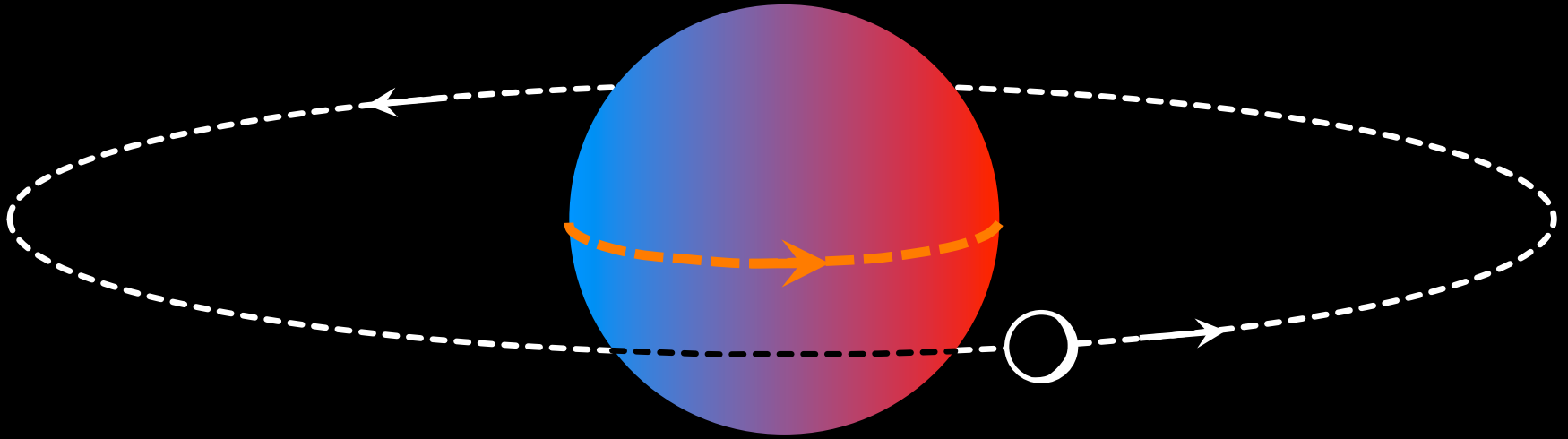




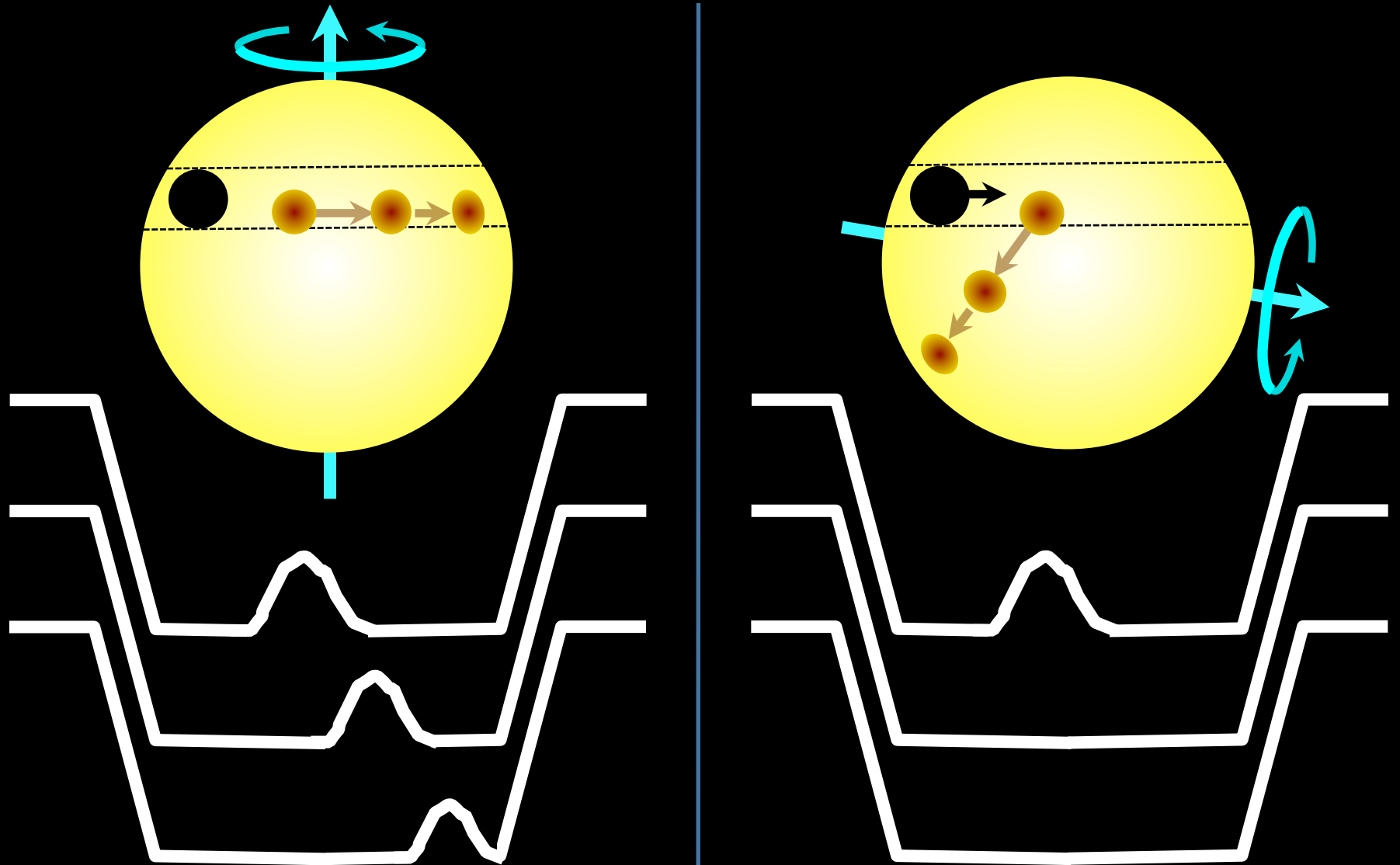
# Stellar obliquity

- Sun's obliquity is  $7^\circ$  — how typical is this?
- Whatever produces *hot Jupiters* may also perturb inclinations

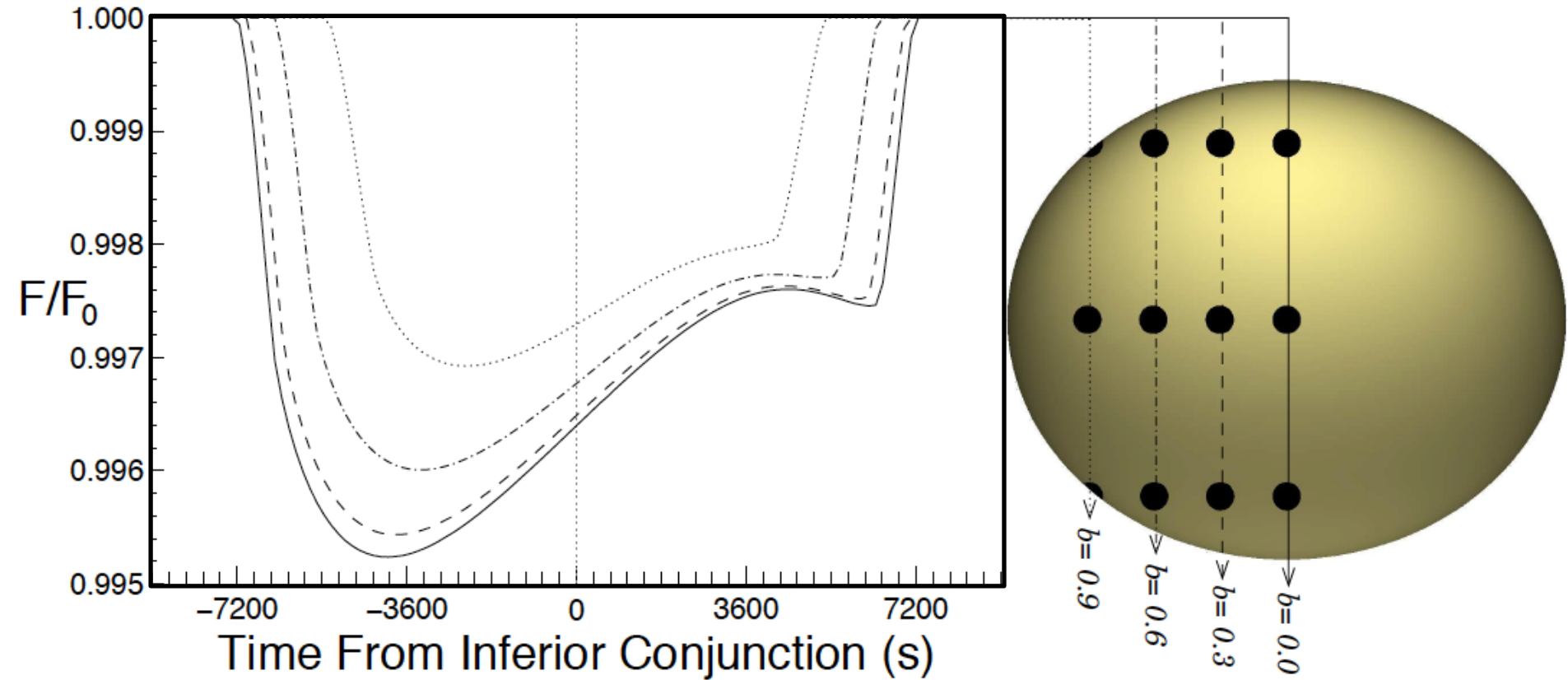
# Rossiter-McLaughlin (RM) effect



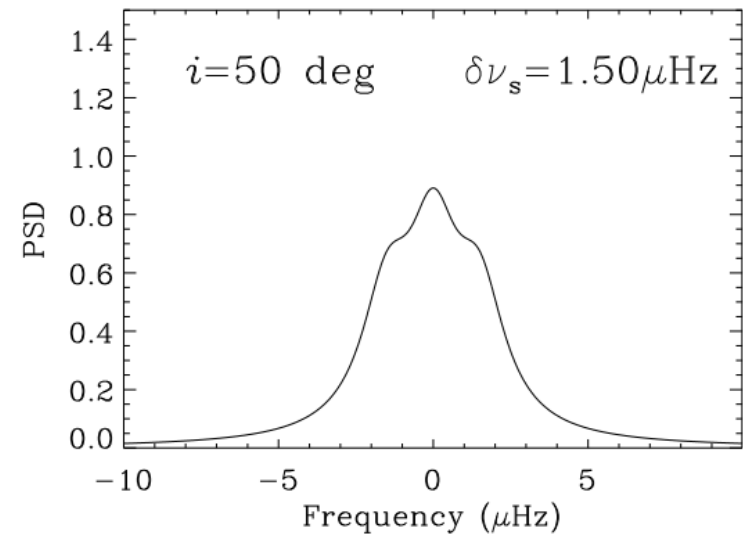
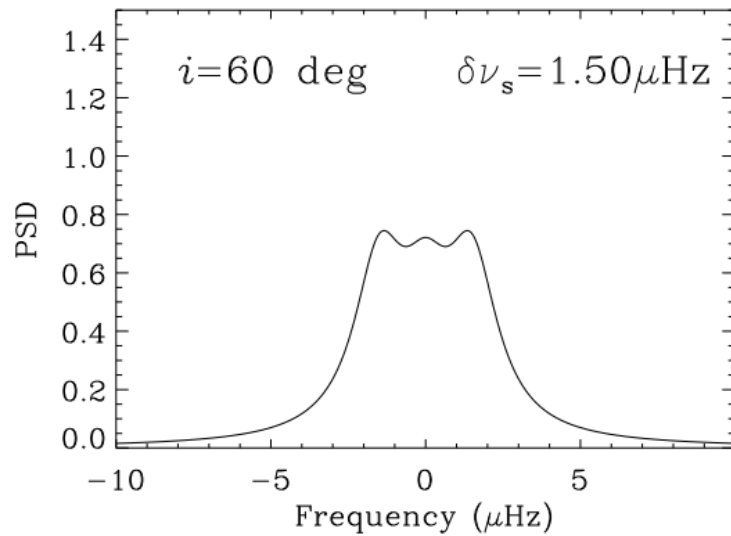
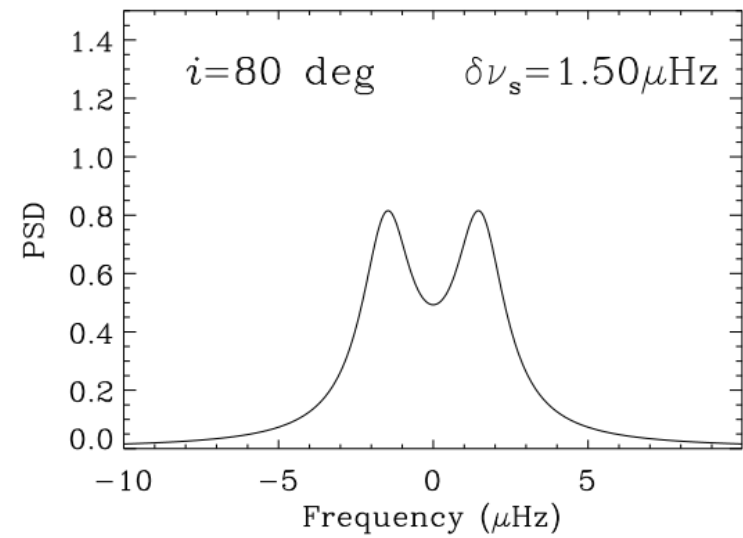
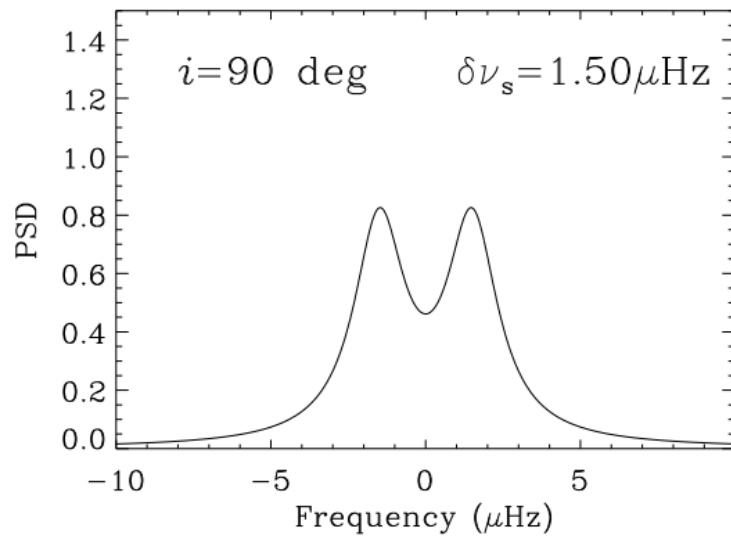
# Sanchis-Nutzman (SN) effect



# Barnes-Szabo (BSz) effect



# Gizon-Solanki (GS) effect

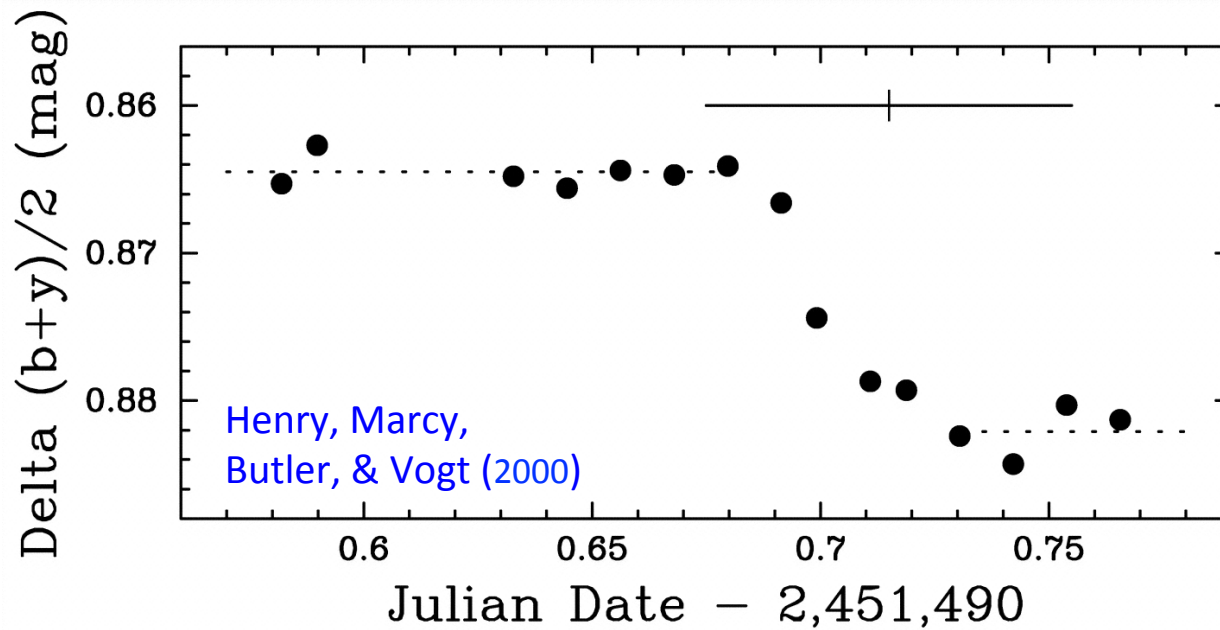
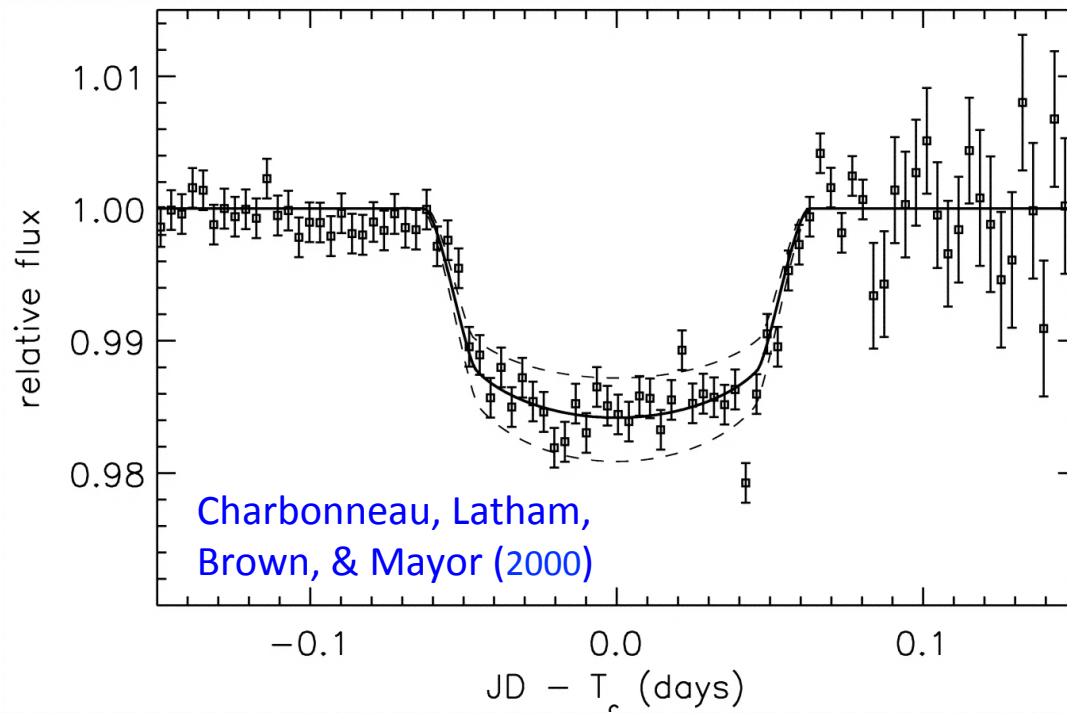






The promise of transits  
**FULFILLED**

When was the first  
announcement of the  
detection of a planetary  
transit?





Jeremiah Horrocks (1639)



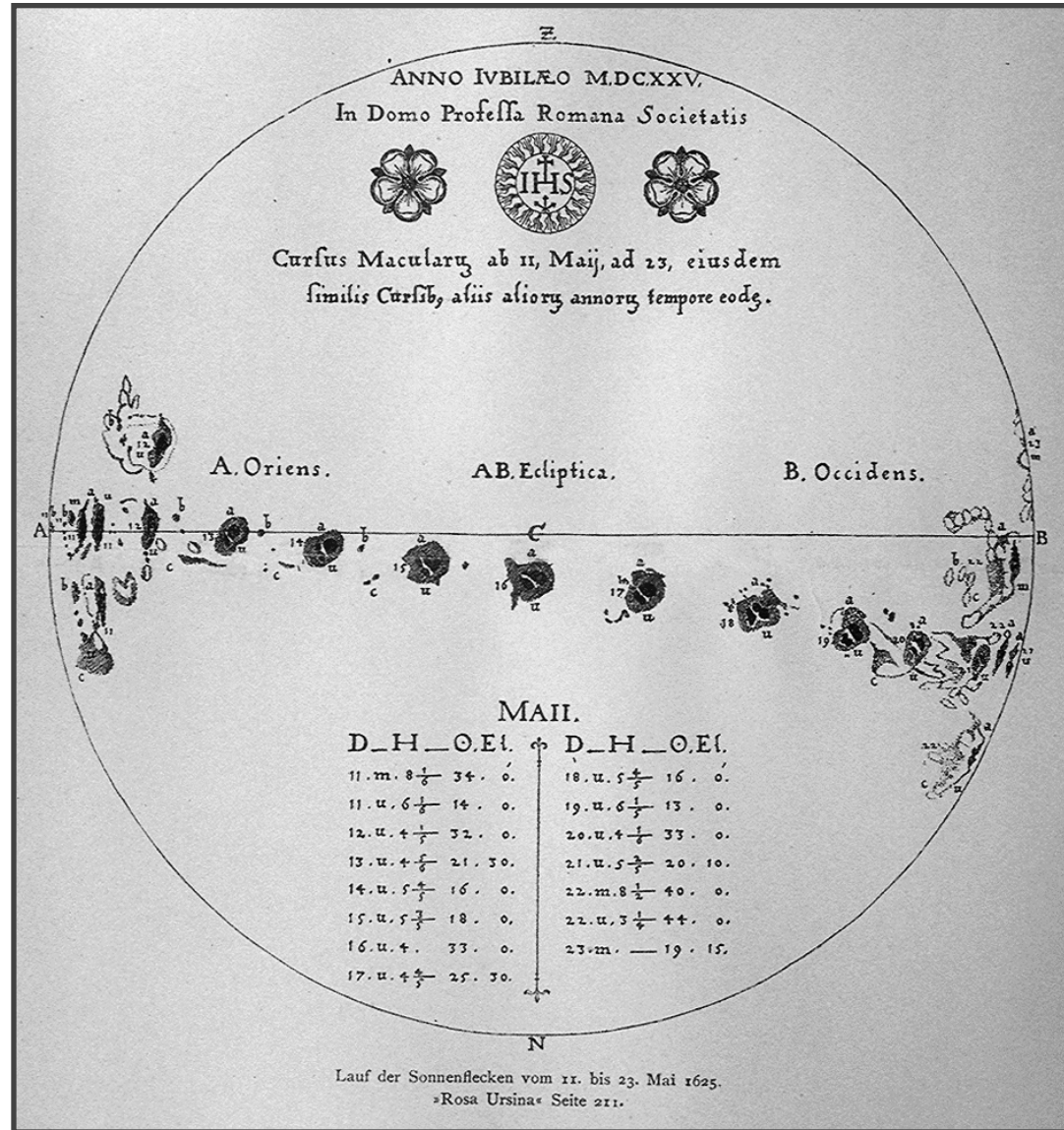
Christoph Scheiner (1611)



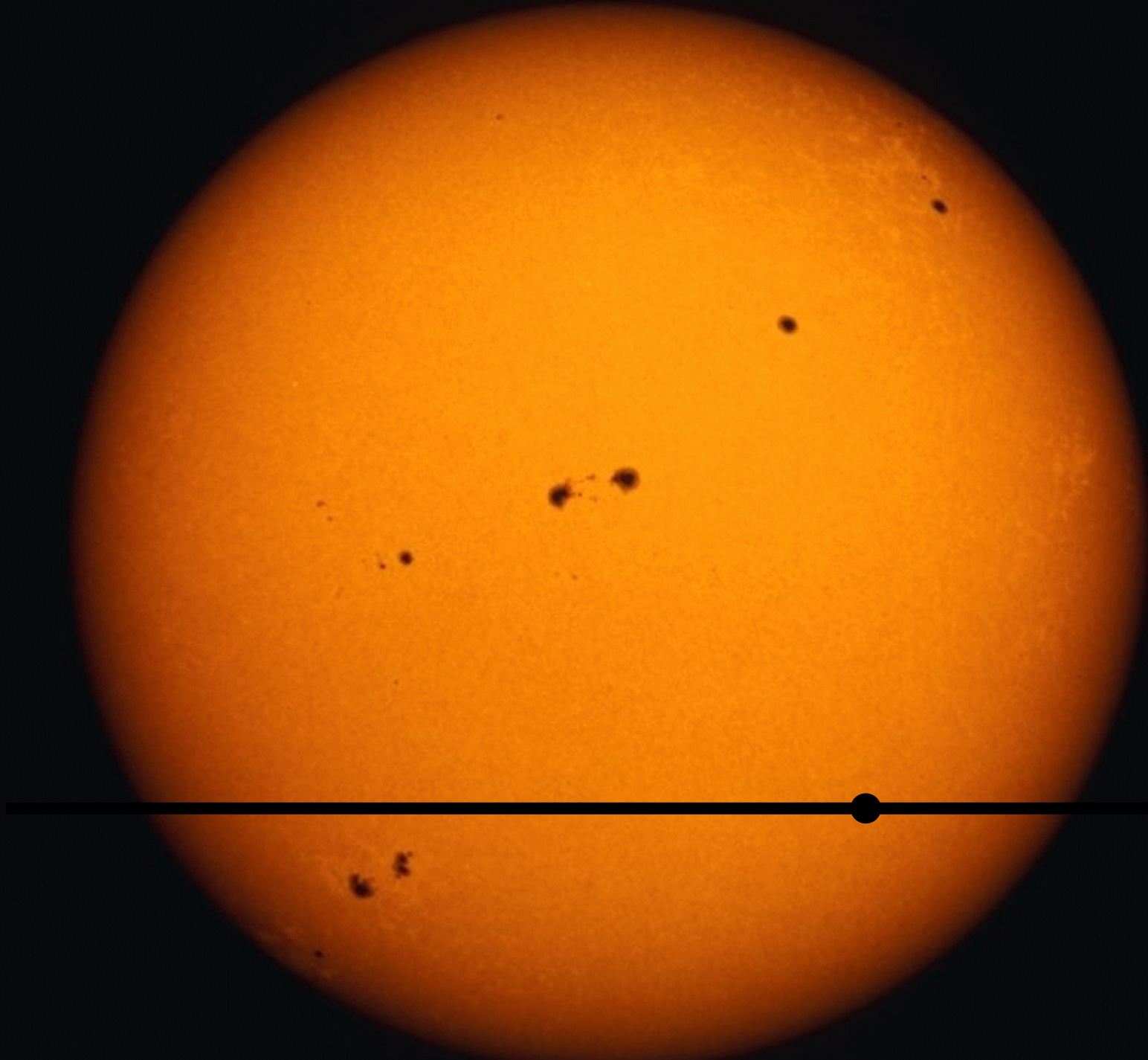
# Rosa Ursina



Christoph Scheiner (1611)









VOYAGE  
DANS  
LES MERS DE L'INDE,



FAIT PAR ORDRE DU ROI,

A l'occasion du PASSAGE DE VÉNUS,  
sur le Disque du Soleil, le 6 Juin 1761,  
& le 3 du même mois 1769.

Par M. LE GENTIL, de l'Académie Royale des Sciences.

Imprimé par ordre de Sa Majesté.

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TOME PREMIER.

---



A PARIS,  
DE L'IMPRIMERIE ROYALE.

---

M. DCCLXXIX.



## Guillaume Le Gentil

Destination captured

Waited 8 years

Clouded out

Contracted dysentery

Shipwrecked

Declared dead





John Goodricke



“A series of observations on, and a discovery of, the period of variation of the light of the Bright Star in the Head of Medusa, Called Algol,” *Phil. Trans.*, 73, 474 (1783)

# PROPOSAL FOR A PROJECT OF HIGH-PRECISION STELLAR RADIAL VELOCITY WORK

*By Otto Struve*

*The Observatory*, 72, 199

Berkeley Astronomical Department,  
University of California.

1952 July 24.



There seems to be at present no way to discover objects of the mass and size of Jupiter; nor is there much hope that we could discover objects **ten times as large in mass** as Jupiter, if they are at distances of one or more astronomical units from their parent stars.

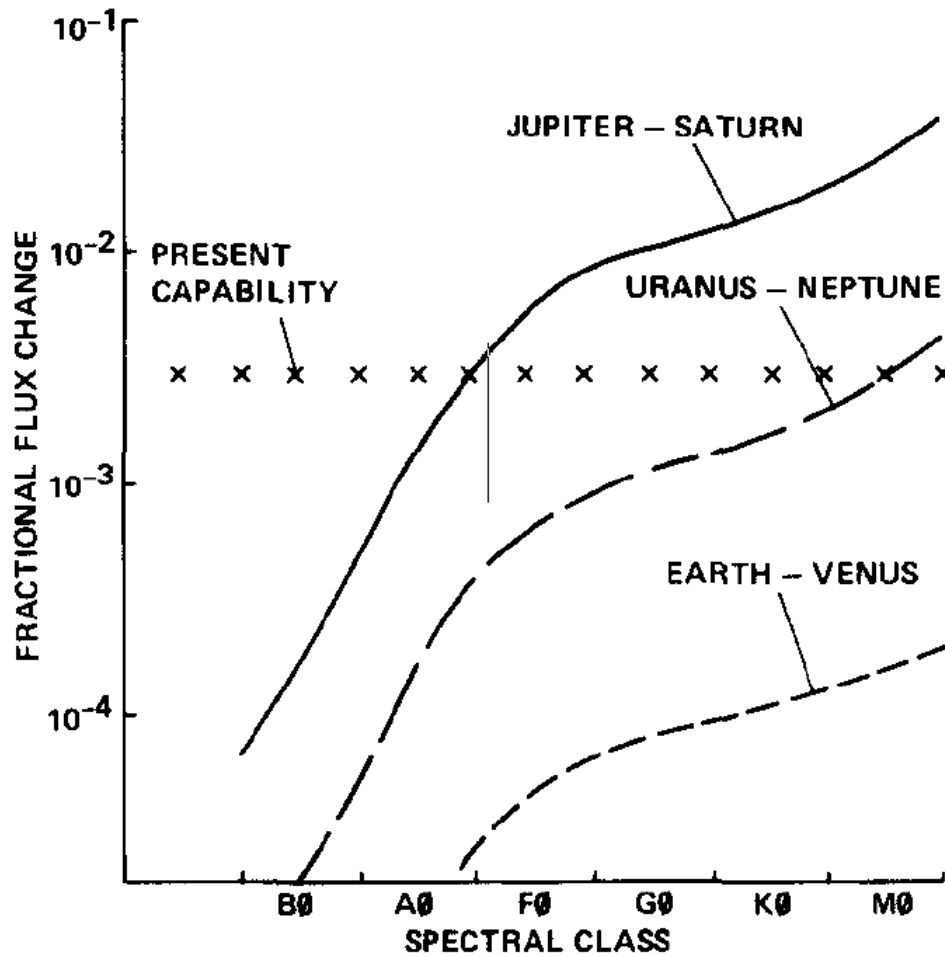
But there seems to be no compelling reason why the hypothetical stellar planets should not, in some instances, be **much closer to their parent stars** than is the case in the solar system.

...it should be possible, without much difficulty, to discover planets of 10 times the mass of Jupiter by the Doppler effect.

***There would, of course, also be eclipses.***

# Modern history of transits

- Borucki & Summers (1984) – survey design



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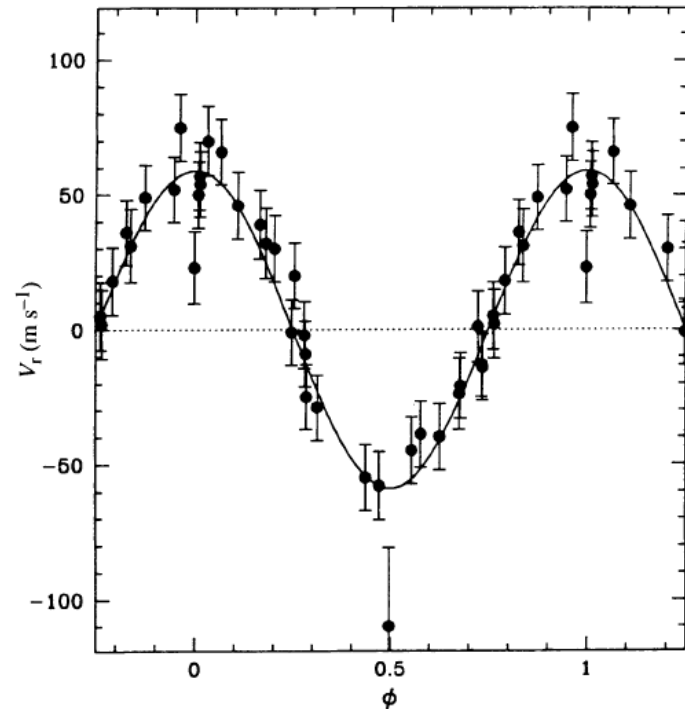
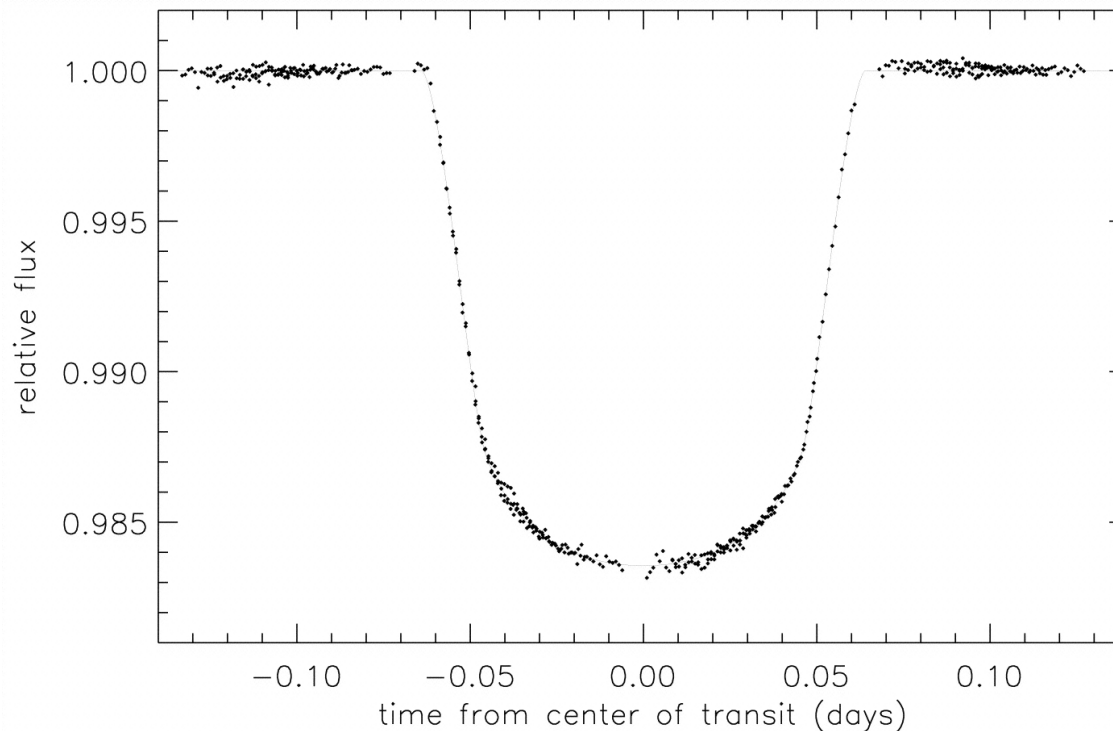


FIG. 4 Orbital motion of 51 Peg corrected from the long-term variation of the  $\gamma$ -velocity. The solid line represents the orbital motion computed from the parameters of Table 1.



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- Mayor & Queloz (1995) – hot Jupiters
- Charbonneau+ (2000), Henry+ (2000), Brown+ (2001)

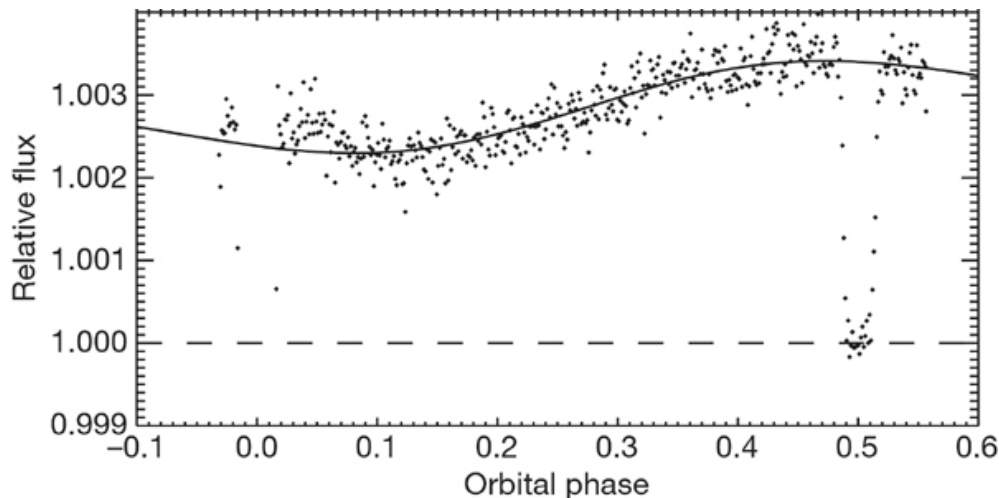


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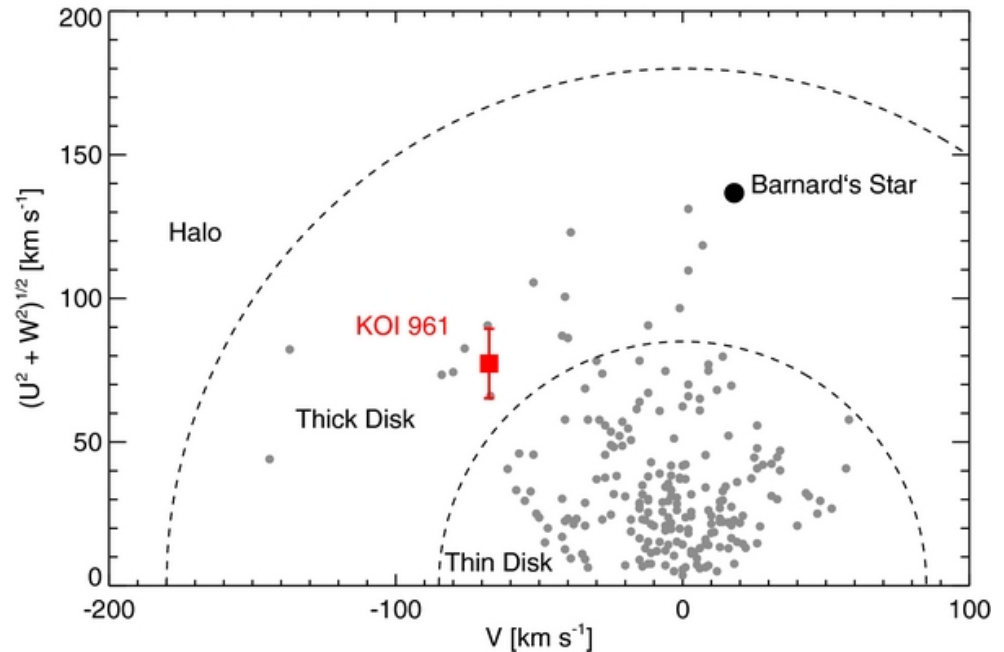
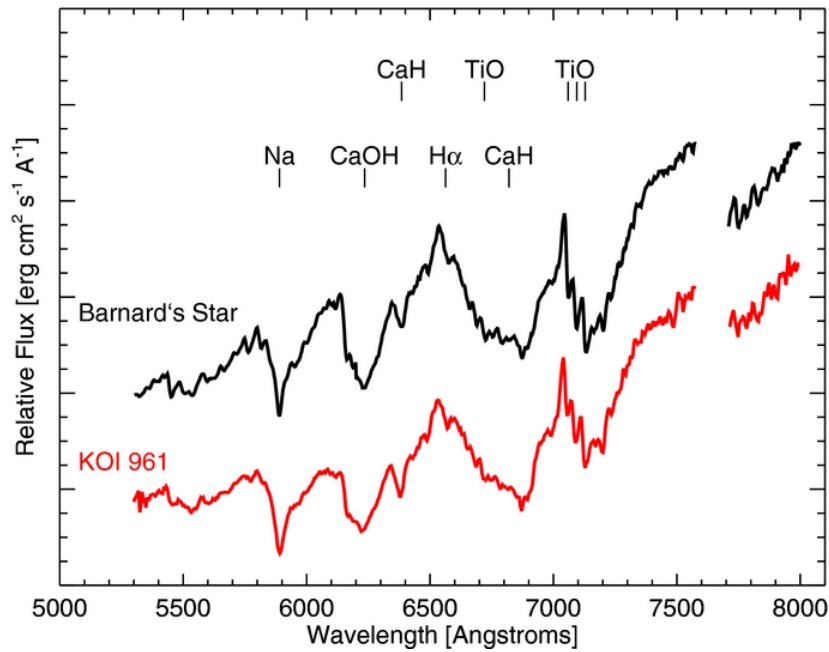


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- Konacki et al. (2003) – Discovery via transit
- Alonso et al. (2004) – Wide-field survey
- Knutson et al. (2007) – Phase curve
- Barge et al. (2008) – first *Corot* results
- Léger et al., Charbonneau et al (2009) – super-Earths
- Borucki et al. (2010) – first *Kepler* results

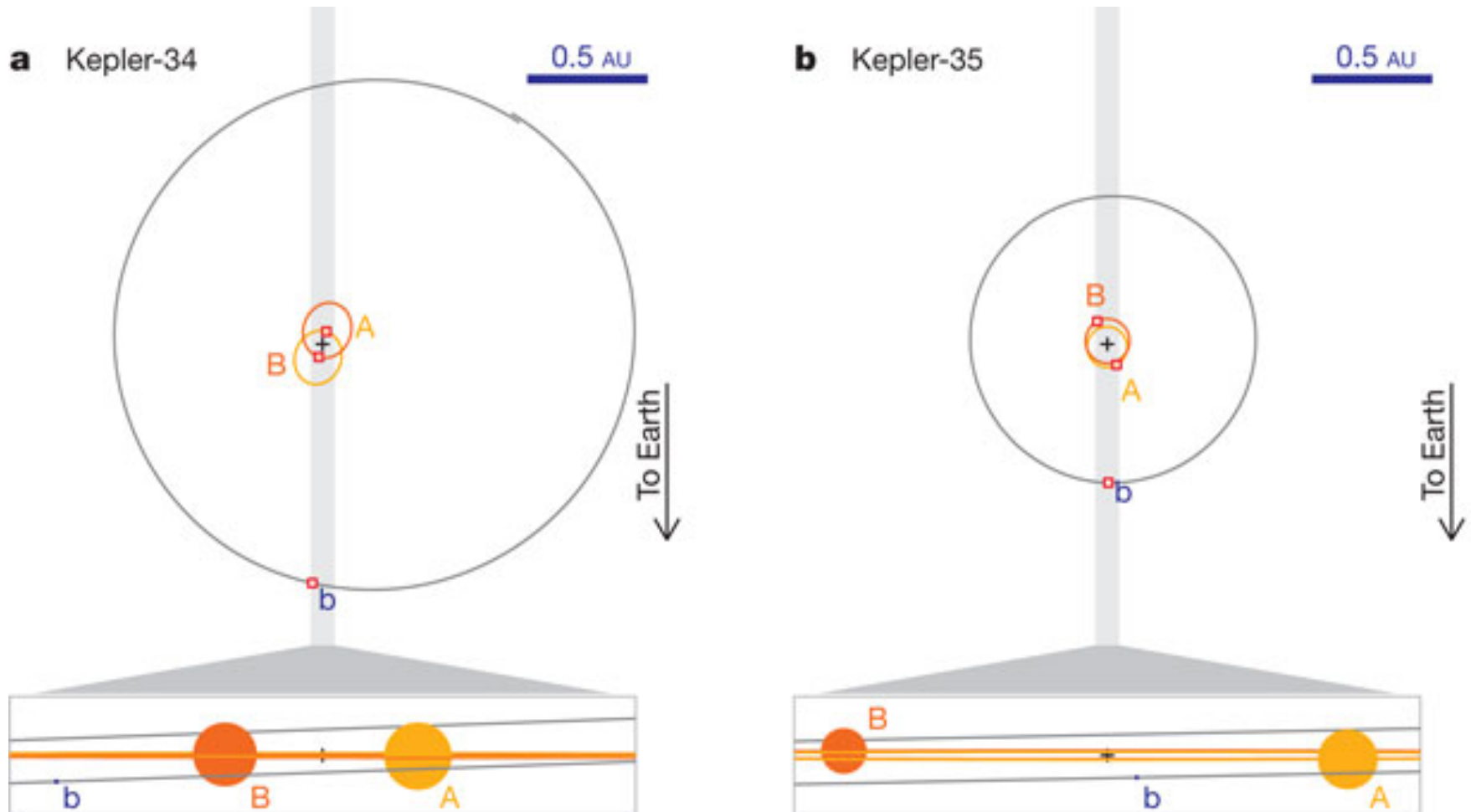
# Characterizing the cool KOIs.

## III. KOI 961: A small star with large proper motion and three small planets

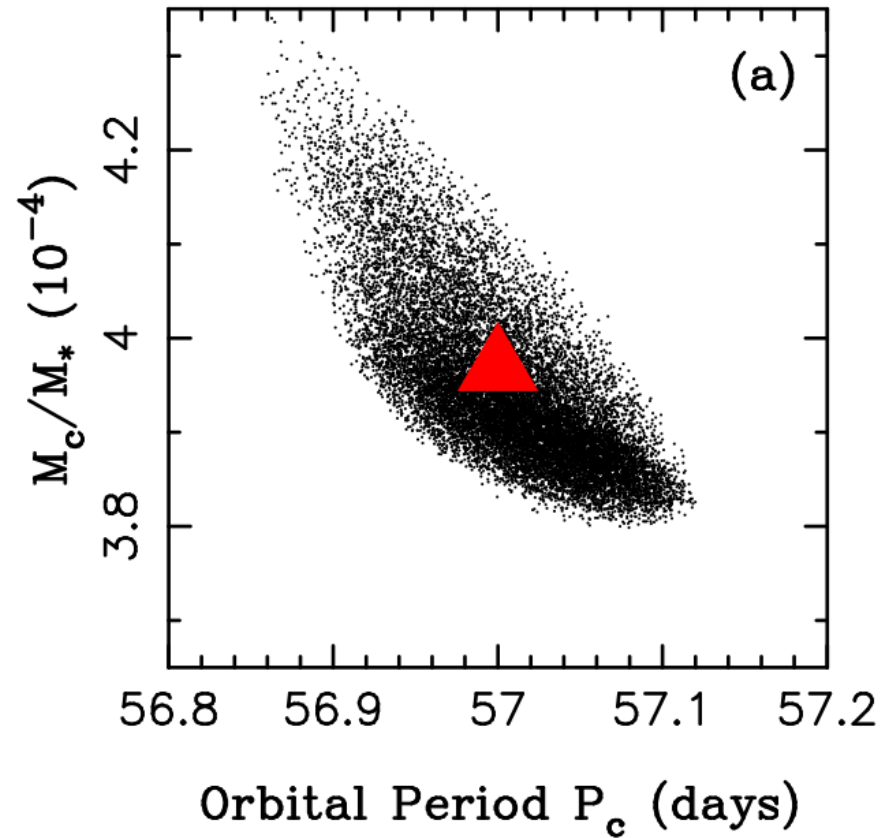
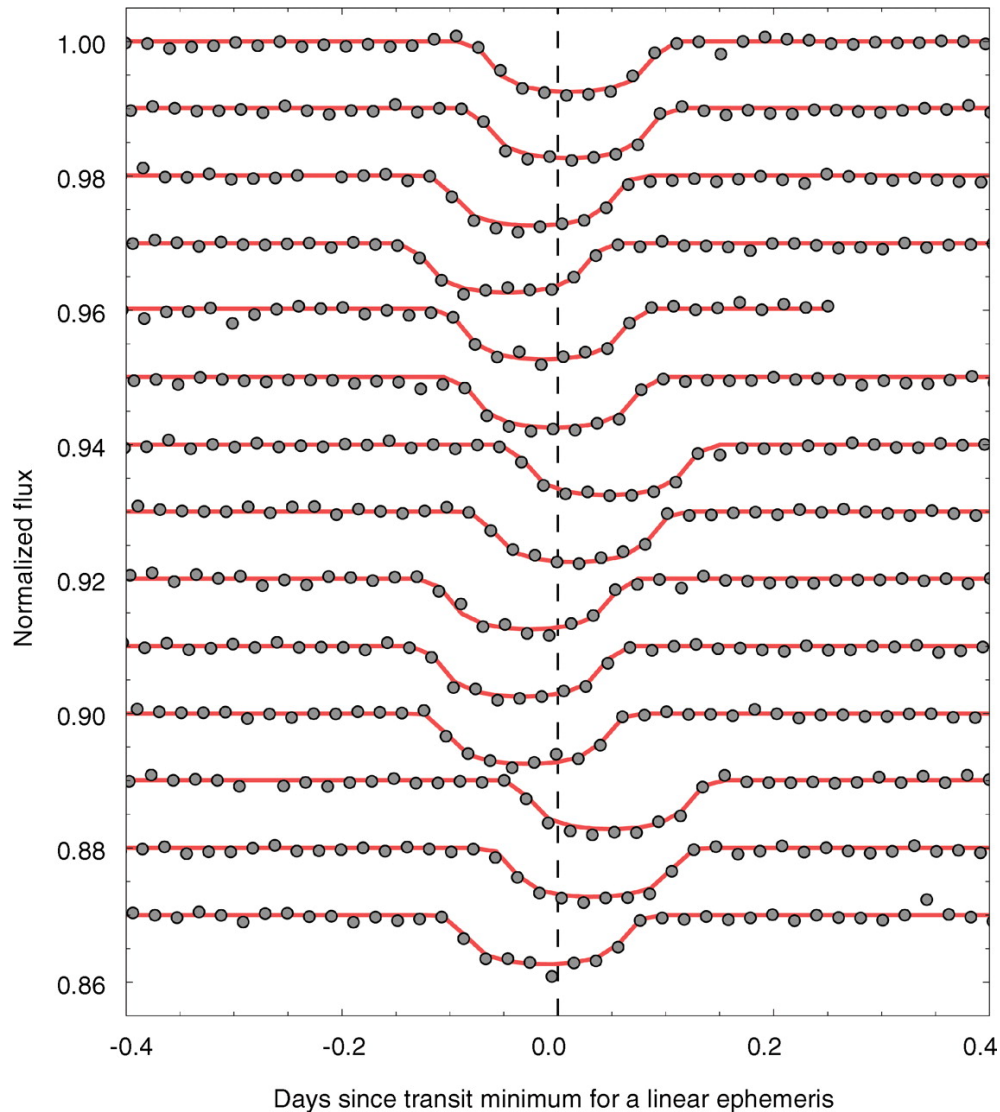


KOI	Period (days)	$R_P$ ( $R_{\oplus}$ )
961.01	$1.2137672 \pm 4.6 \times 10^{-6}$	$0.78 \pm 0.22$
961.02	$0.45328509 \pm 9.7 \times 10^{-7}$	$0.73 \pm 0.20$
961.03	$1.865169 \pm 1.4 \times 10^{-5}$	$0.57 \pm 0.18$

# Transiting circumbinary planets Kepler-34 b and Kepler-35 b



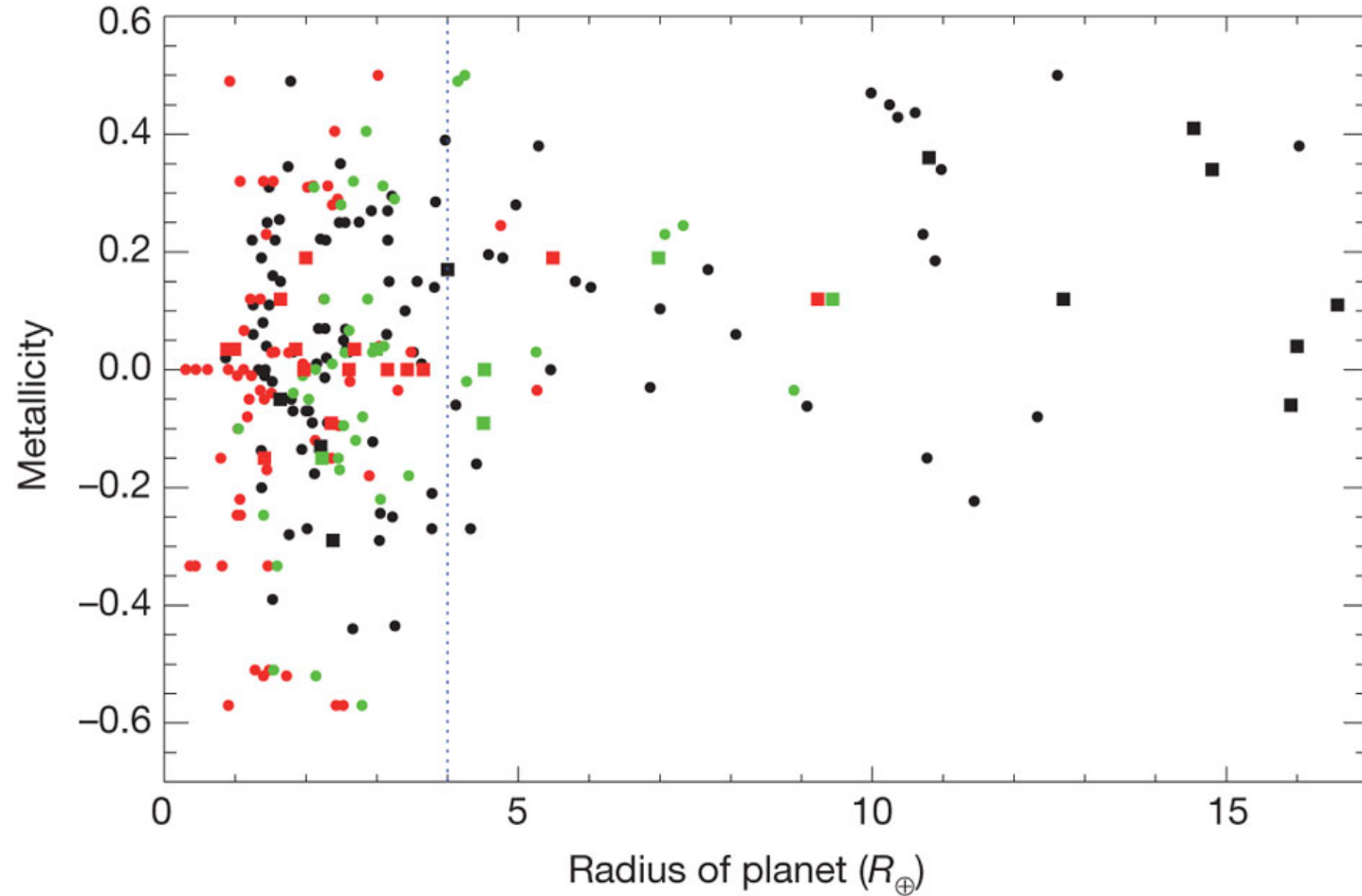
# The detection and characterization of a nontransiting planet by transit timing variations



Nesvorny et al. (2012)  
see also Ballard et al. (2011)



# An abundance of small exoplanets around stars with a wide range of metallicities



**Figure 3 | Individual host-star metallicity as a function of planet radius.**

The black dots represent single-planet systems, whereas the green dots represent the largest planet and the red dots represent all the smaller planets in multiple-planet systems. The confirmed, published Kepler planets in our samples are plotted as squares with the same colour code as the dots.

# Rapid dynamical chaos in an exoplanetary system

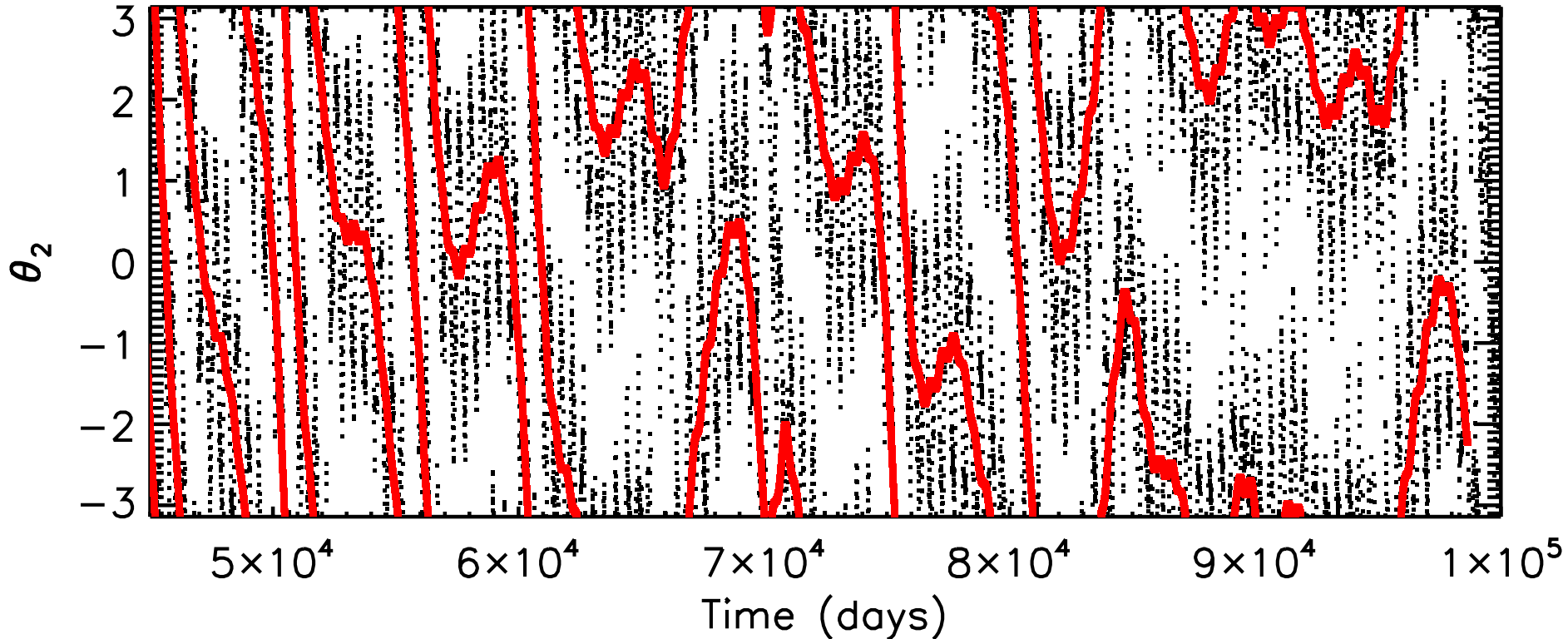
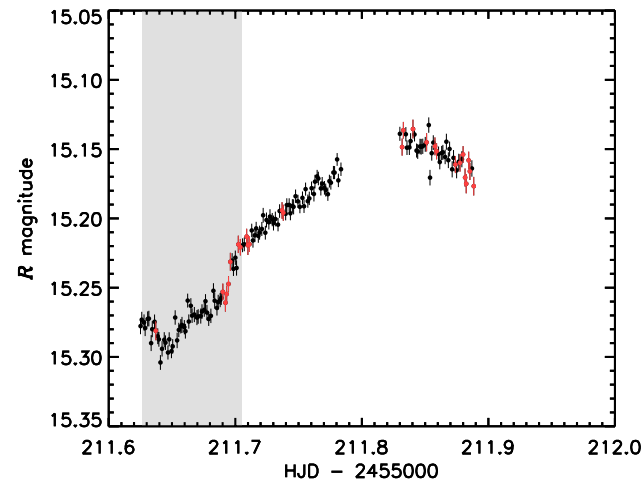
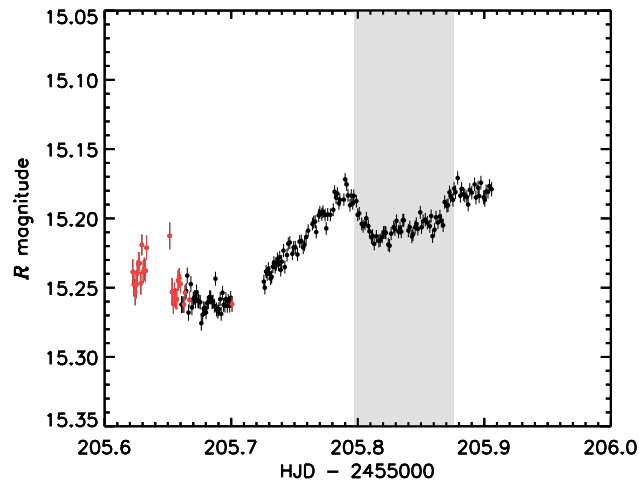
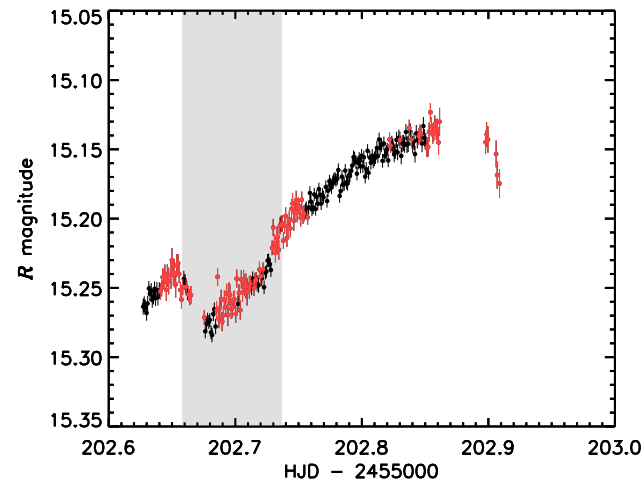
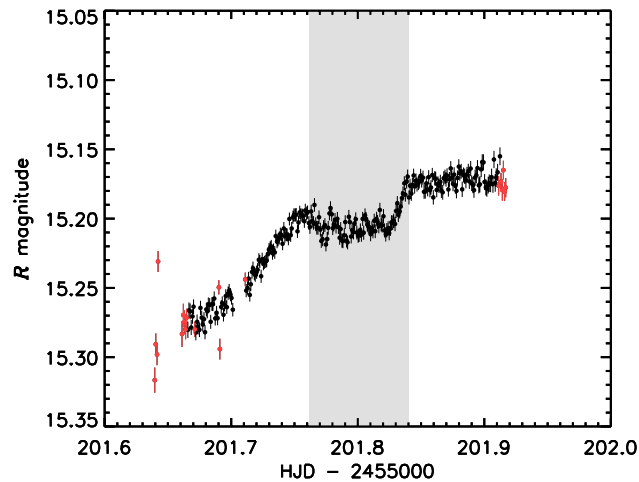


FIG. 2.— Chaotic evolution of the resonant angles  $\theta_1 = 34\lambda_c - 29\lambda_b - 5\varpi_b$  and  $\theta_2 = 34\lambda_c - 29\lambda_b - 5\varpi_c$  for a randomly chosen trajectory from the long lived region. The red overlaid points show a smoothed version of the black points to guide the eye.

# The PTF Orion Project: A possible planet transiting a T-Tauri star



Gold mine

# Gold mine – or gold rush?



“Every new field has its little run of luck, and the exoplanet game is going to be tougher at *<prestigious journal>* now that that’s over.”

— astronomy editor of *<prestigious journal>*

# Reasons for optimism

It always seems “too late” ●

The peak is probably broad

Both contenders for the next  
NASA Explorer mission are  
transit missions (TESS and  
FINESSE)

*Spitzer*: hot Jupiters :: *JWST* :  
small planets



# Reasons for optimism

Our most fundamental and precise knowledge of *stars* comes from eclipsing systems, even after more than a century since eclipses were first observed. The same is likely to be true for exoplanets.