



From Compact to Fuzzy: The New Interior Paradigm for Jupiter, Saturn, and Gas Giant Exoplanets

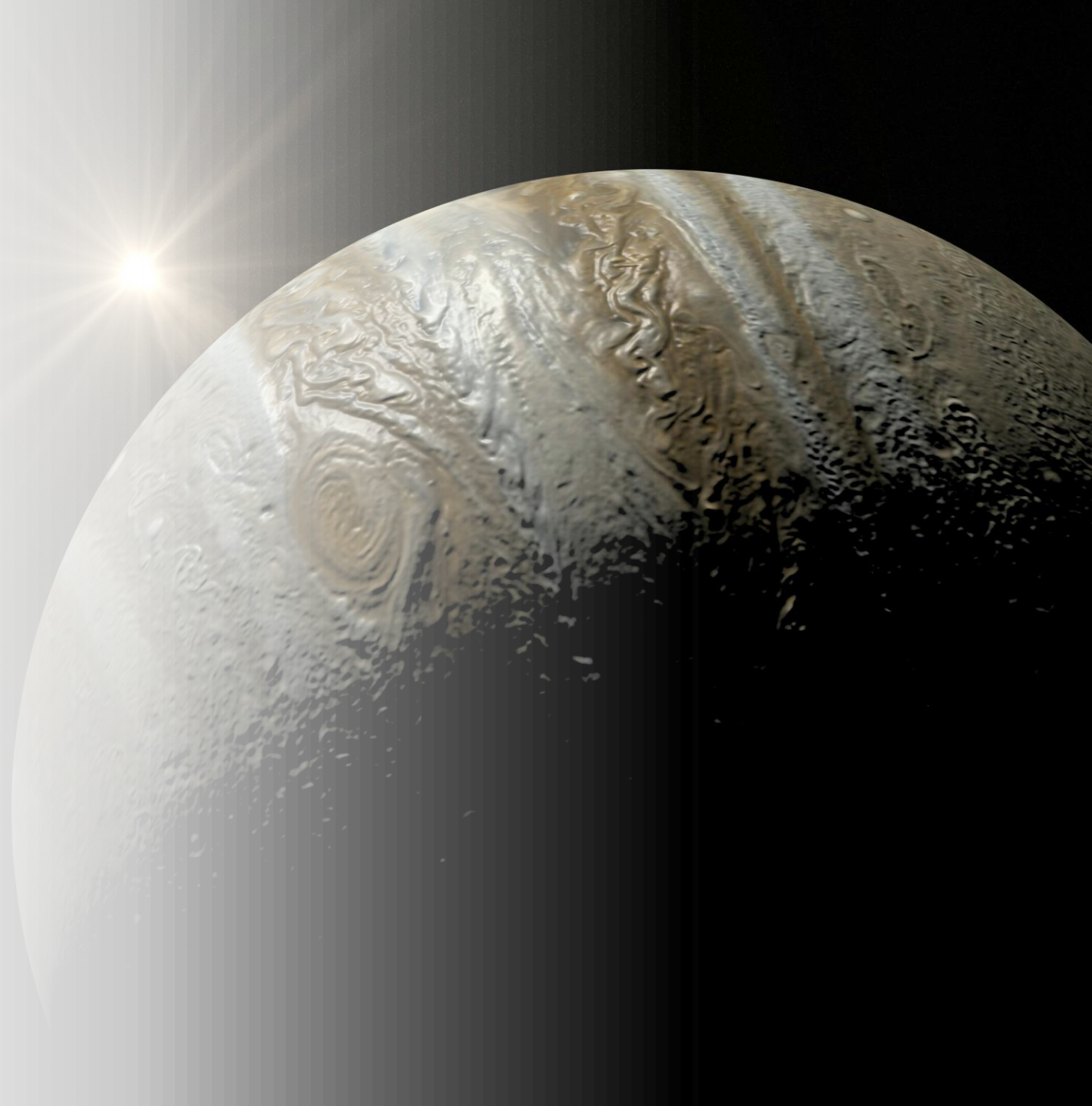
Ankan Sur

Earth, Planetary, and Space Sciences

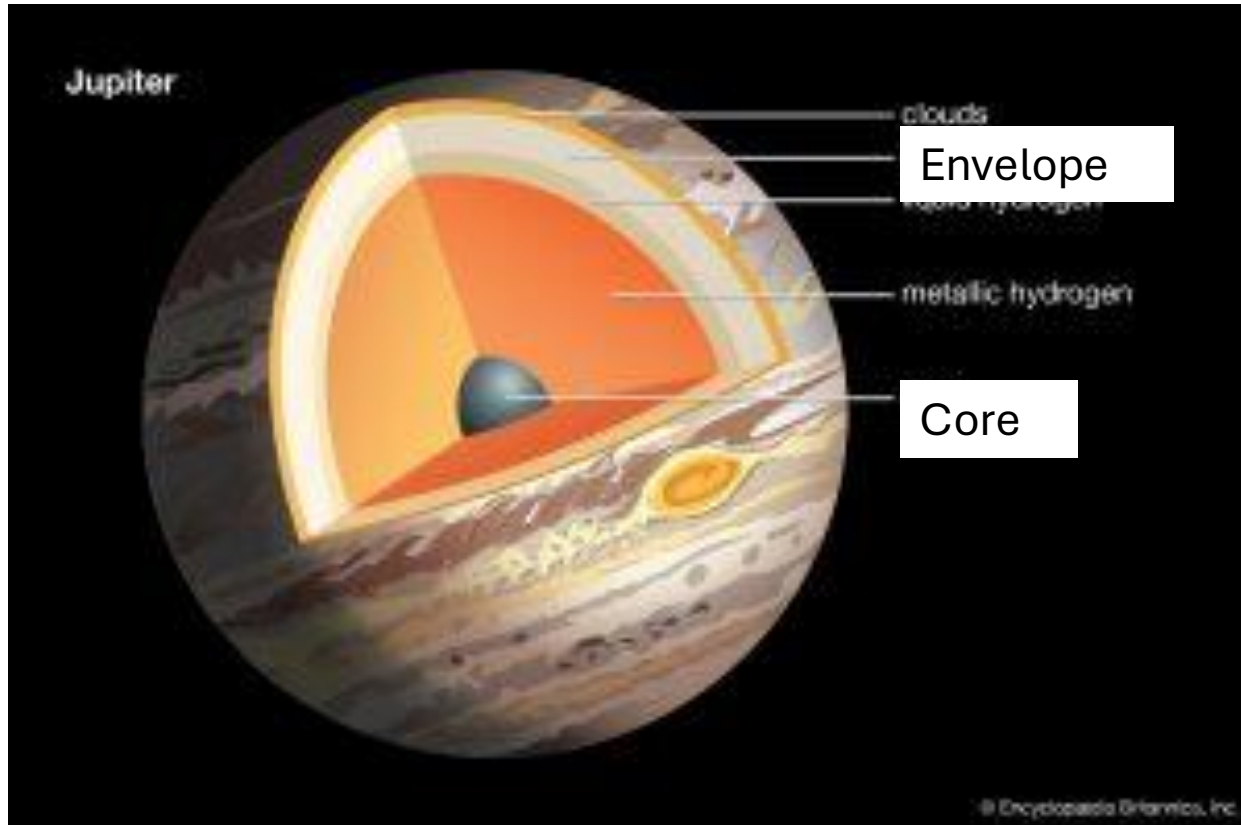
UCLA

ExSoCal 2025

Ucla



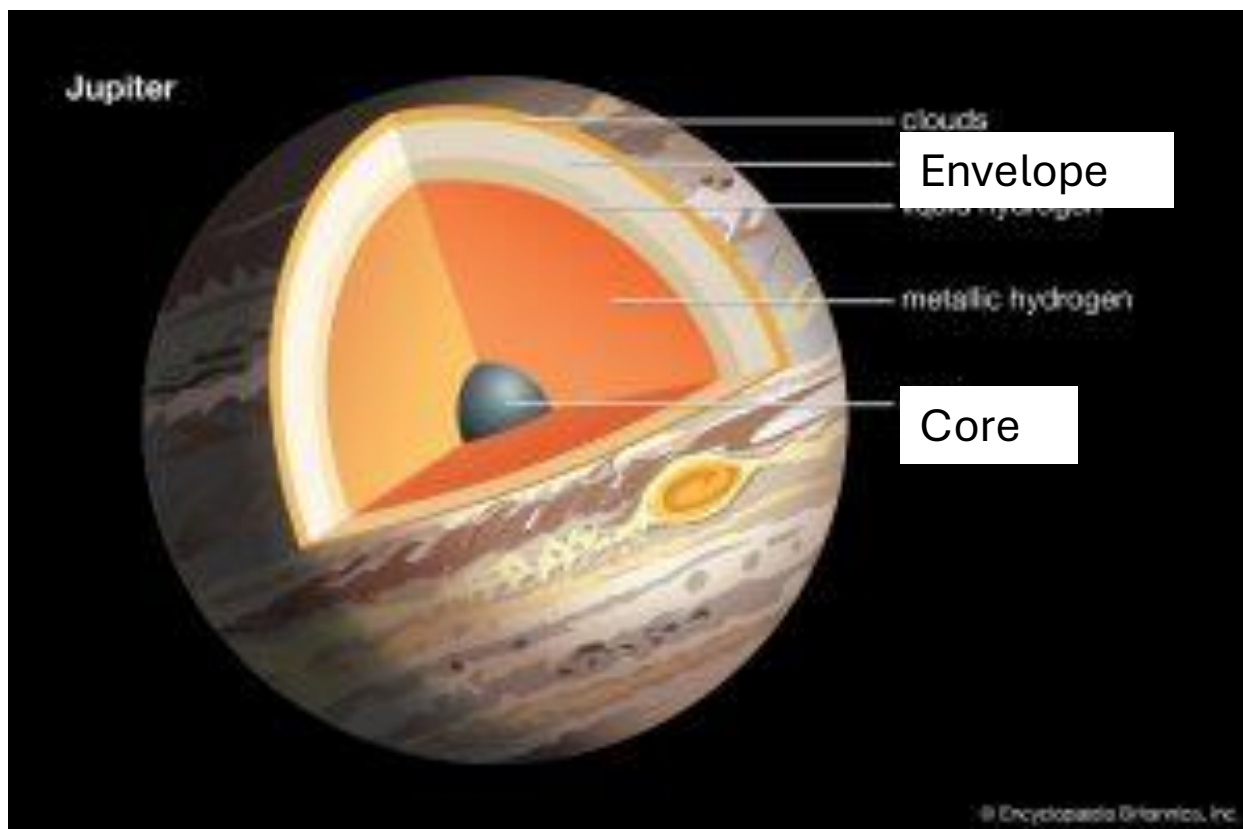
Planet Structure and Evolution: old picture



- Planet consists of a

Rocky Core + Hydrogen-Helium Envelope
- Gives the right radius

Planet Structure and Evolution: old picture



Planet is fully convective and homogeneous.
Able to explain the luminosity of Jupiter.

Adiabatic Models: $TdS/dt = -dL/dM$

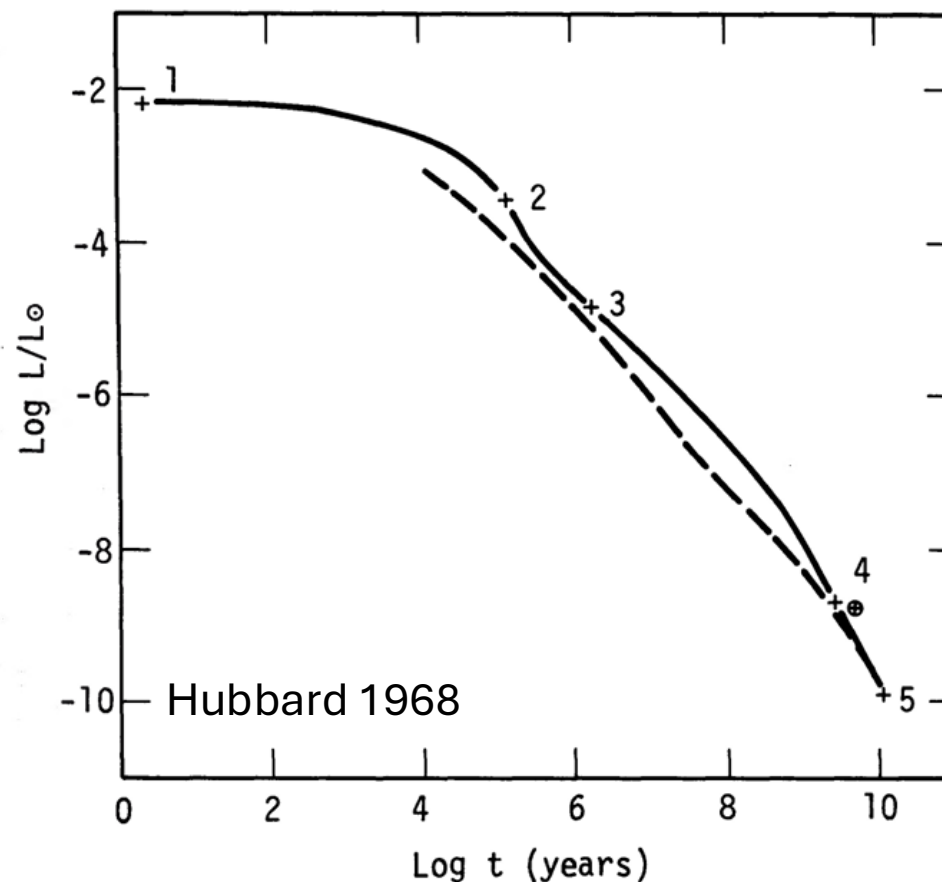
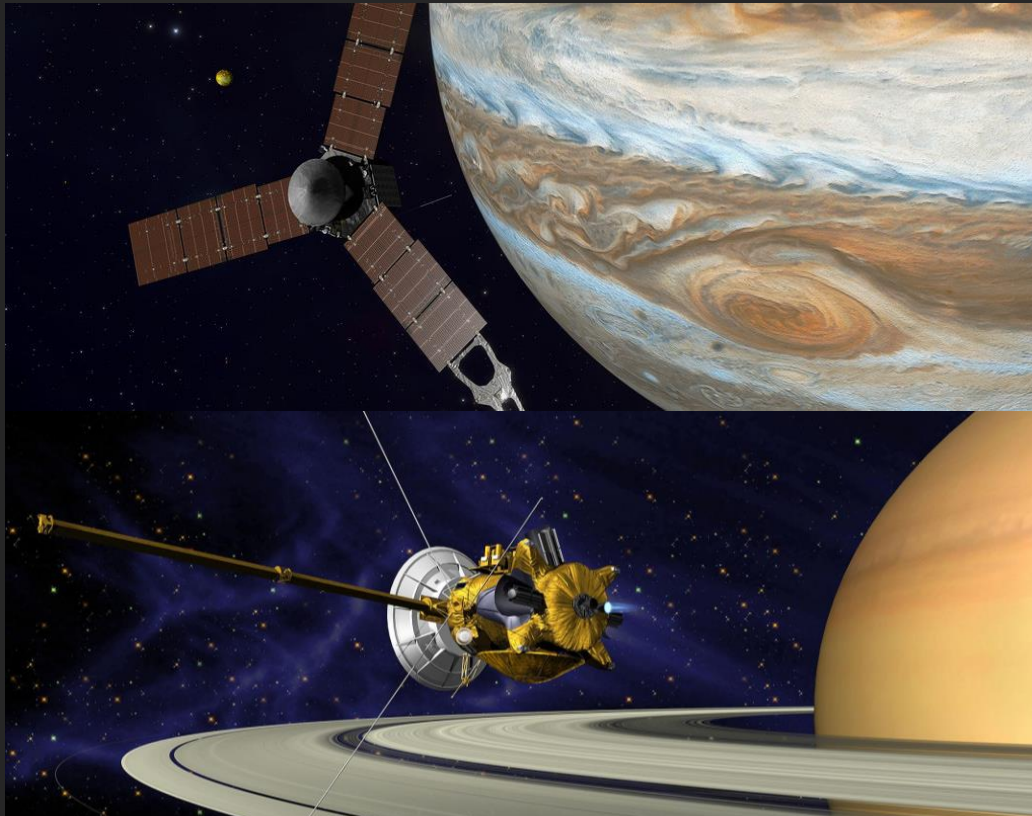
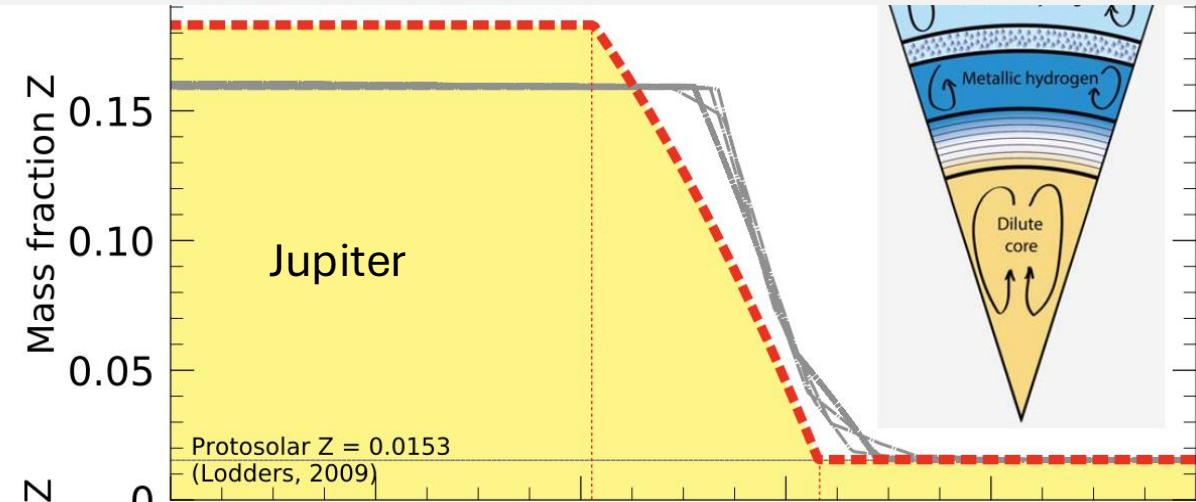


FIG. 3.—Time dependence of luminosity for the S1 standard sequence, with epochs 1 to 5, the Jovian luminosity and the hydrogen sequence shown as in Fig. 2.

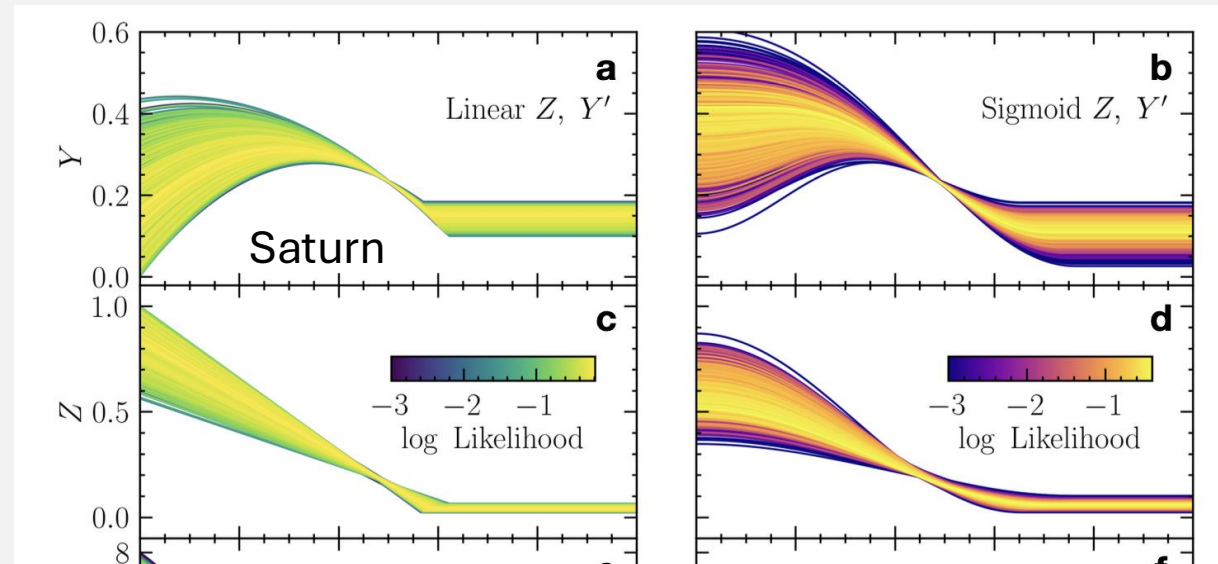
Juno and Cassini: Gravity data, Seismology



Fuzzy Cores



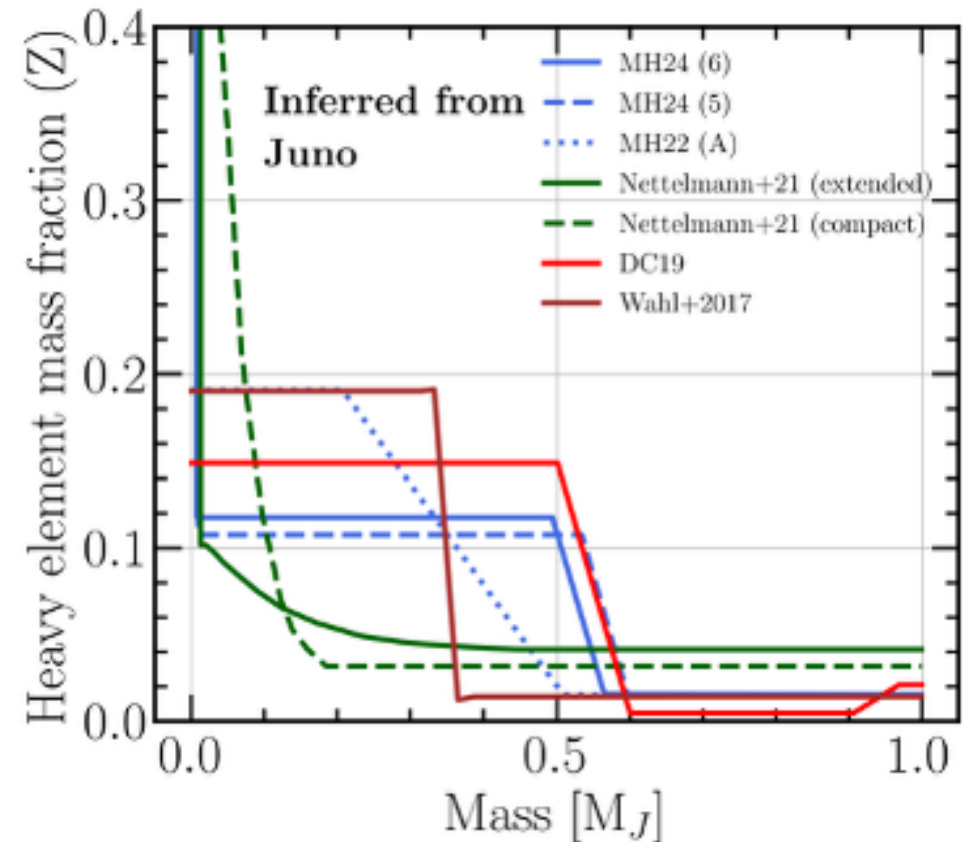
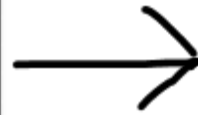
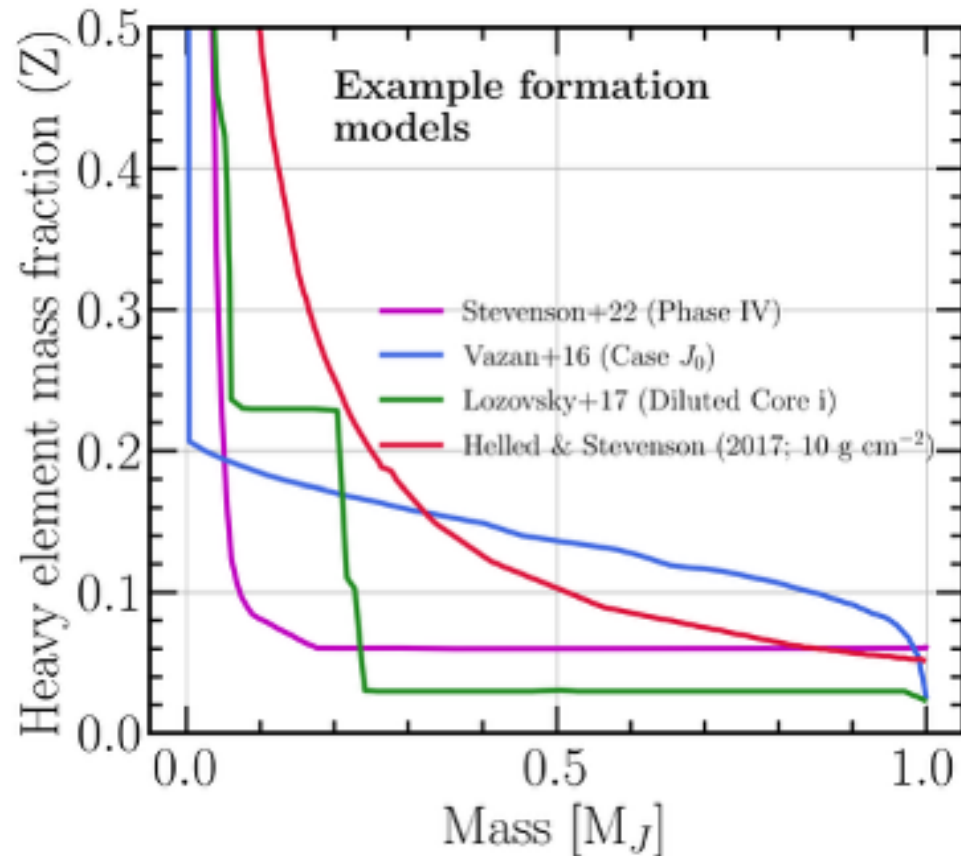
Militzer (2022)



Mankovich&Fuller 2021

From Formation to Current age

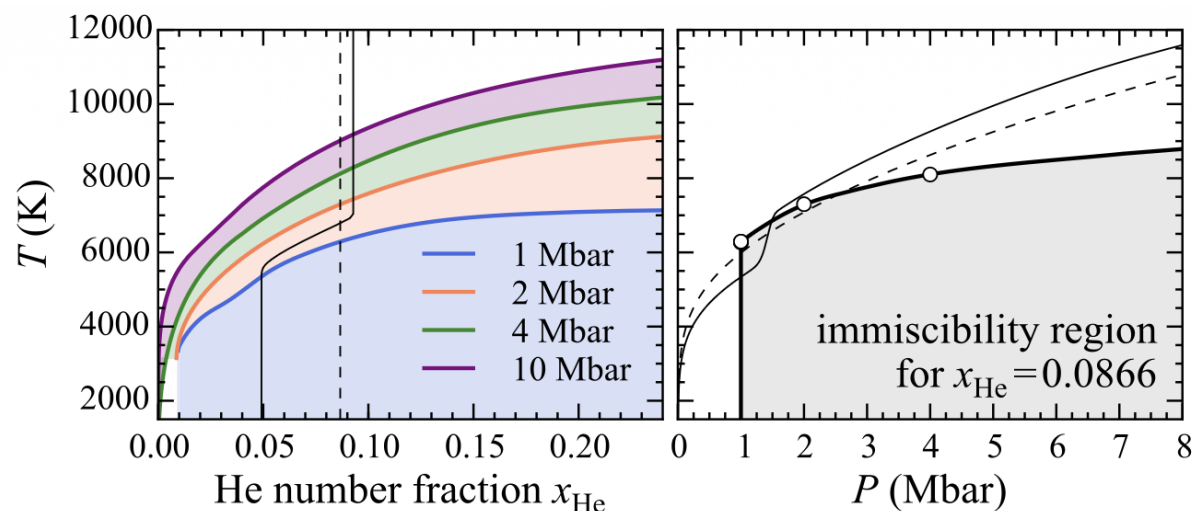
Jupiter



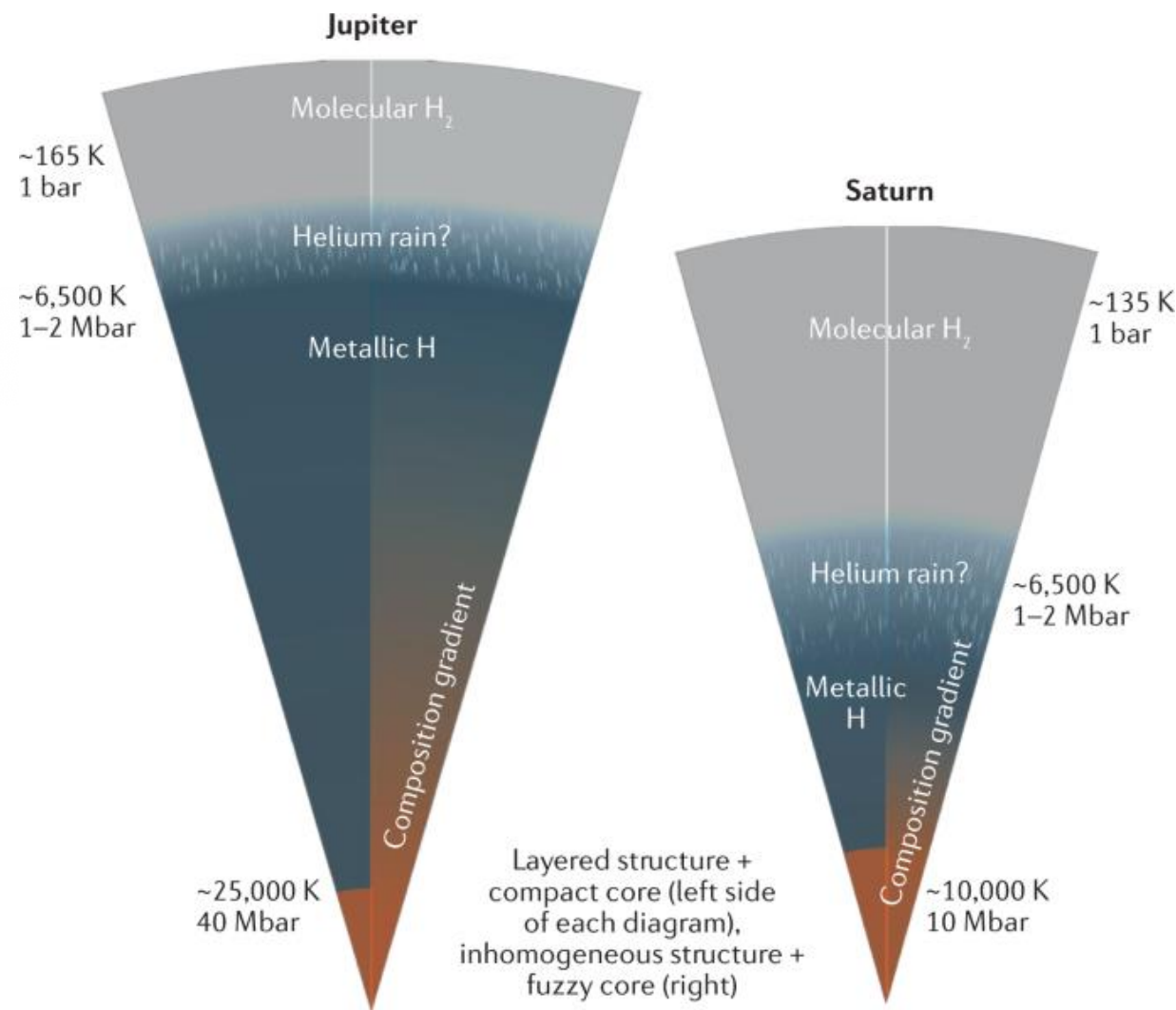
Compiled by Tejada Arevalo

Demixing in Giant Planets: Helium Rain

Hydrogen-Helium miscibility



Lorenzen et al (2009, 2011)

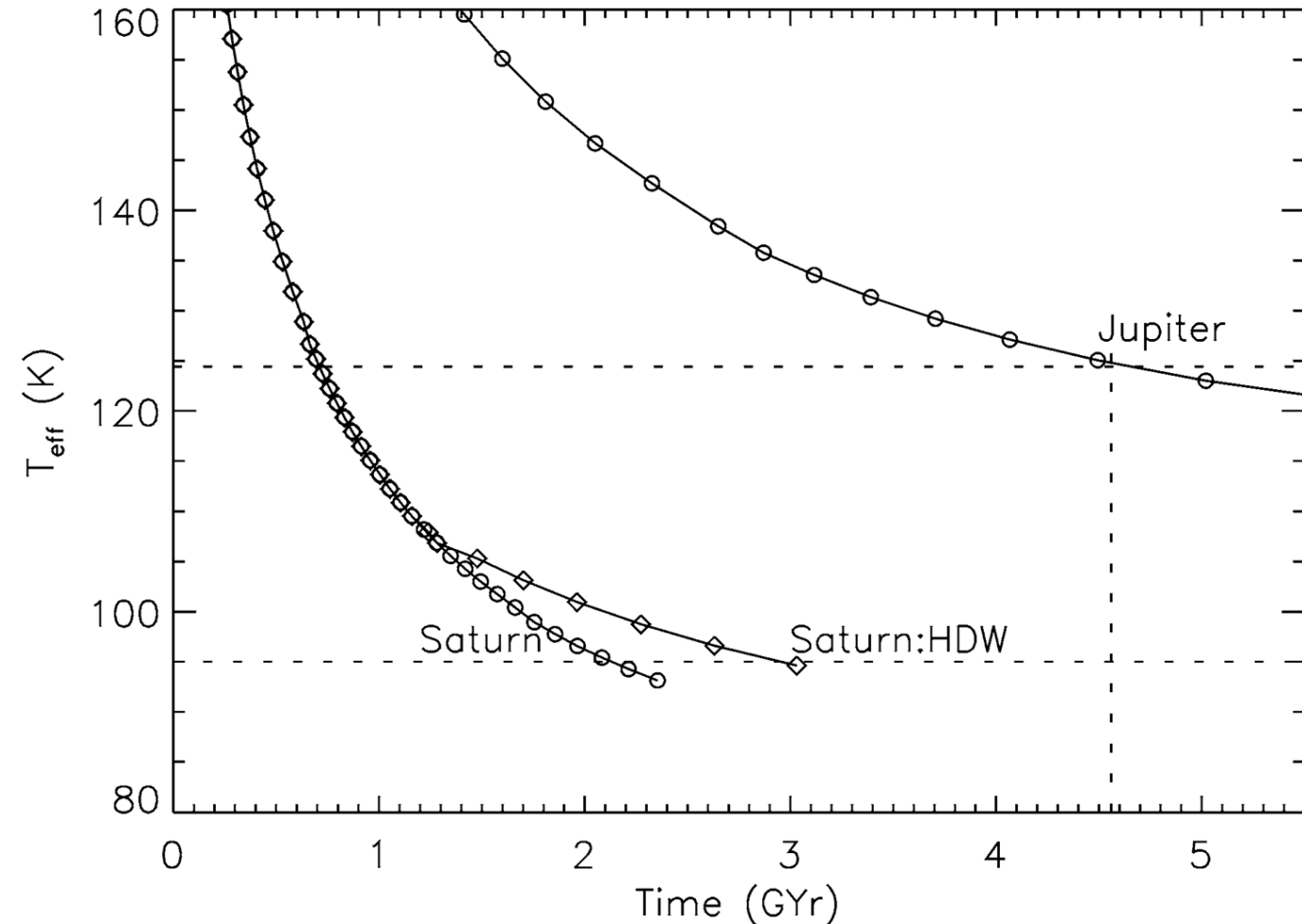


Credit: Ravit Helled

Planet Structure and Evolution: Adiabatic and Inhomogeneous Evolution

- Hydrogen-Helium demixing: helium rain
- Explain Saturn's luminosity (Stevenson & Salpeter 1977)
- Explain Galileo observation: $Y = 0.234 \pm 0.005$

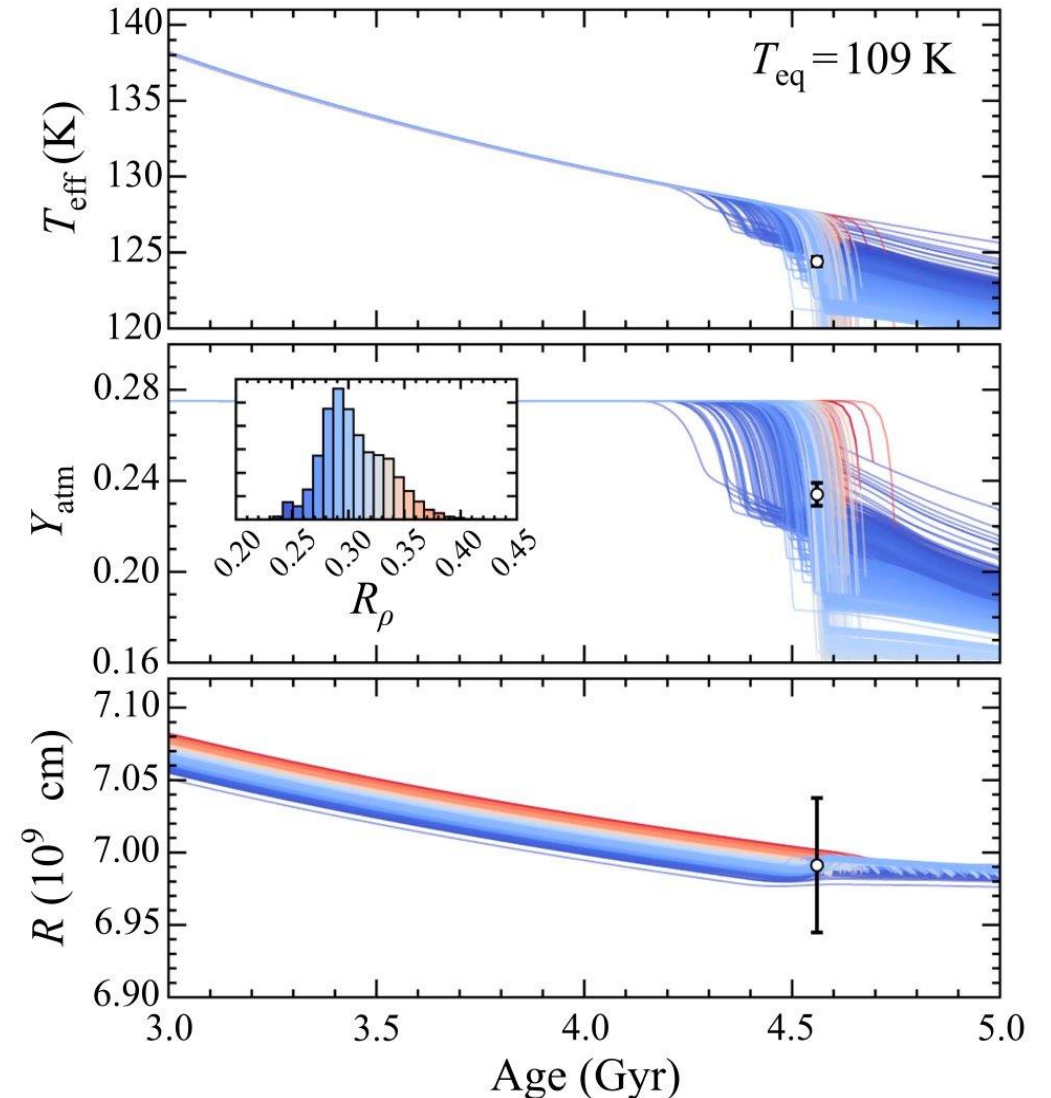
Fortney & Hubbard 2003



Jupiter and Saturn's Structures and Evolution: Adiabatic and Inhomogeneous Evolution

Mankovich & Fortney 2016

- Evolution with helium rain
- Superadiabatic temperature gradients
- Used Bond Albedo=0.5
- Saturn's $Y_{\text{atm}}=0.07$

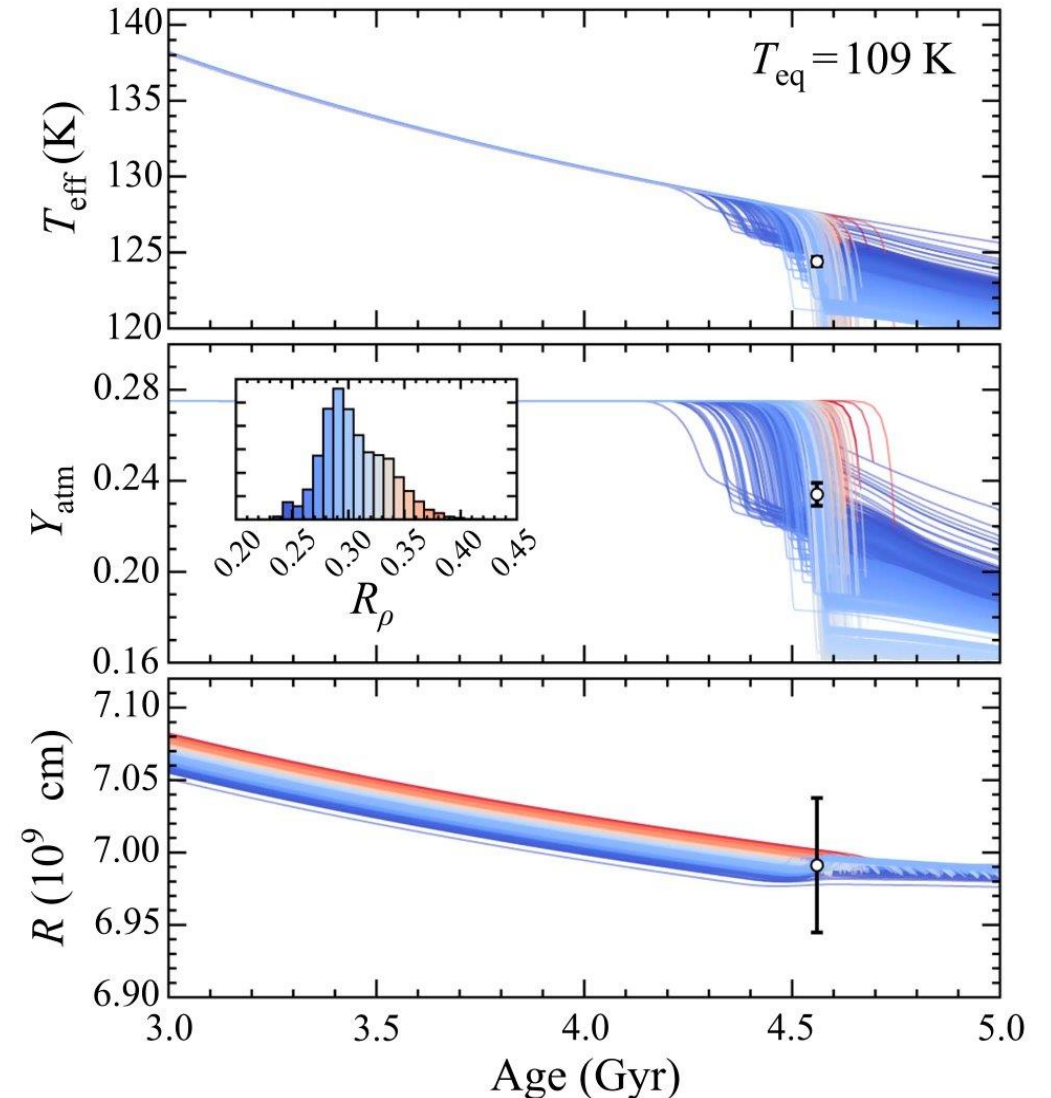


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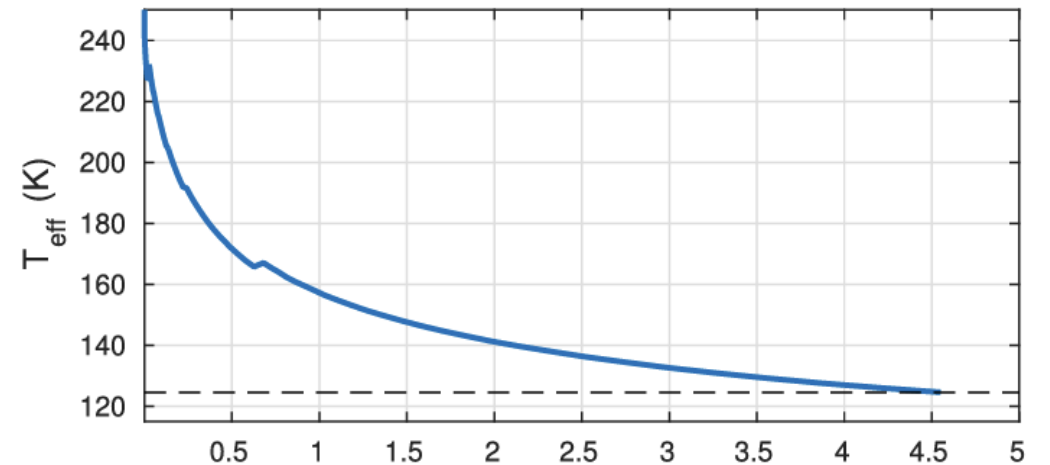
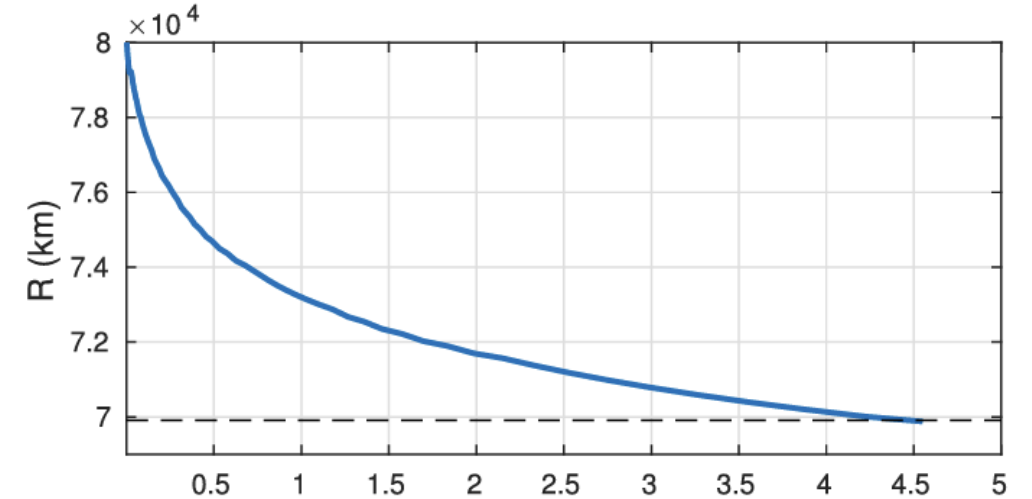
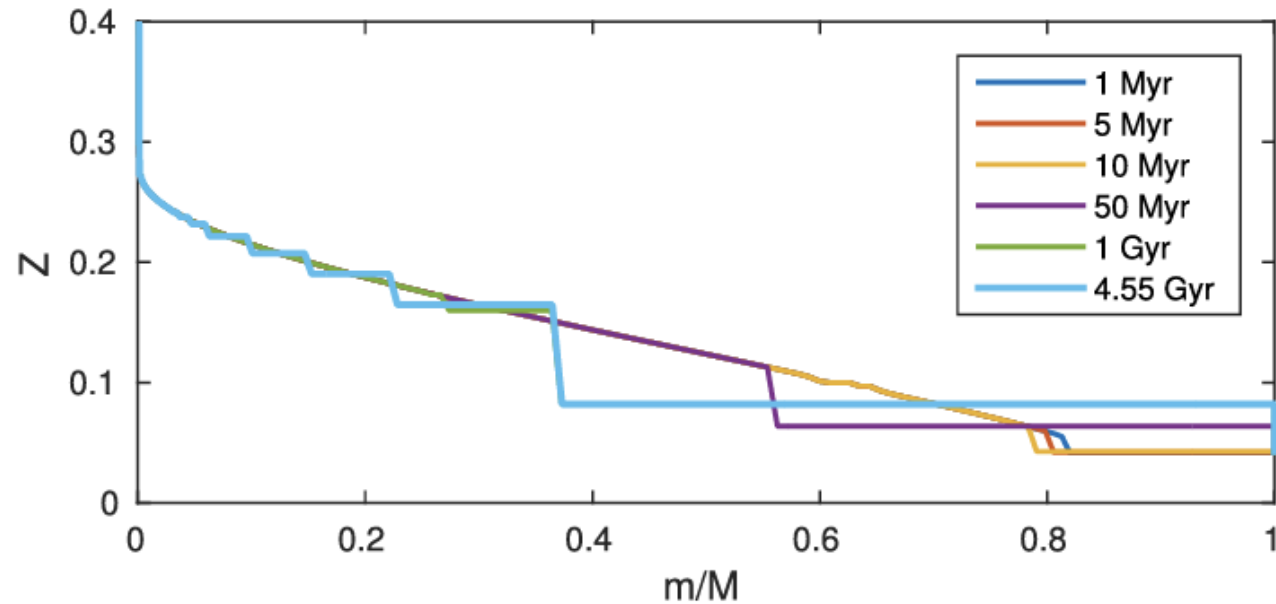
Problem: Did not include Fuzzy Cores



Jupiter Structure and Evolution: Non-adiabatic and Inhomogeneous Evolution

Vazan, Helled, Guillot (2018)

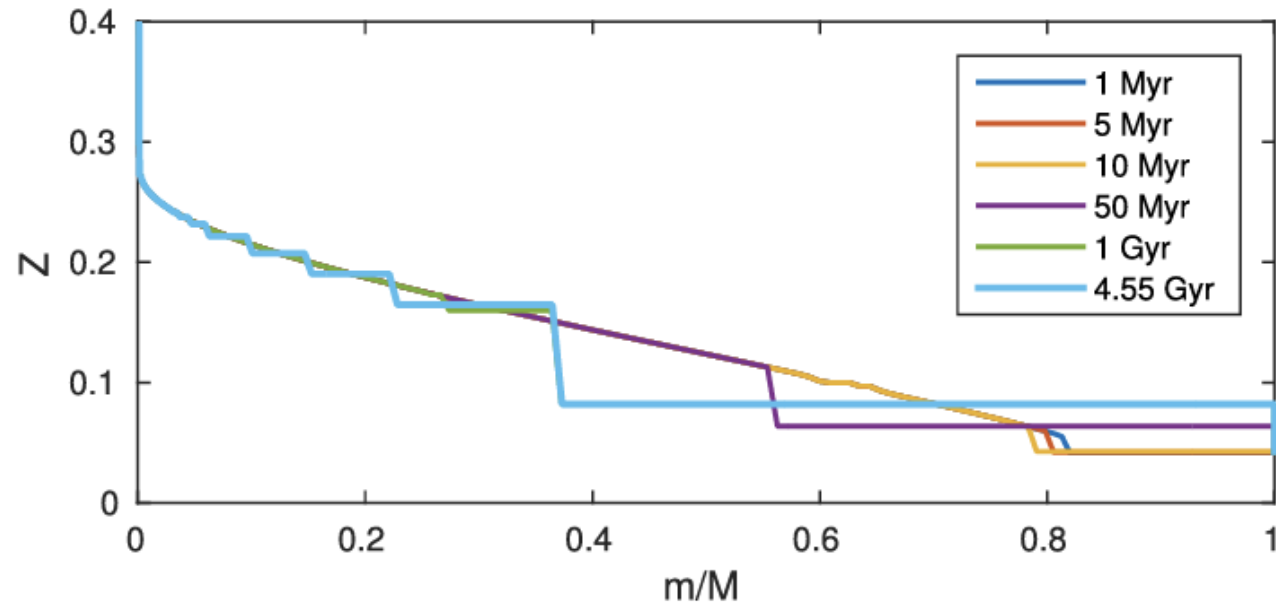
Evolution of Jupiter with composition gradients



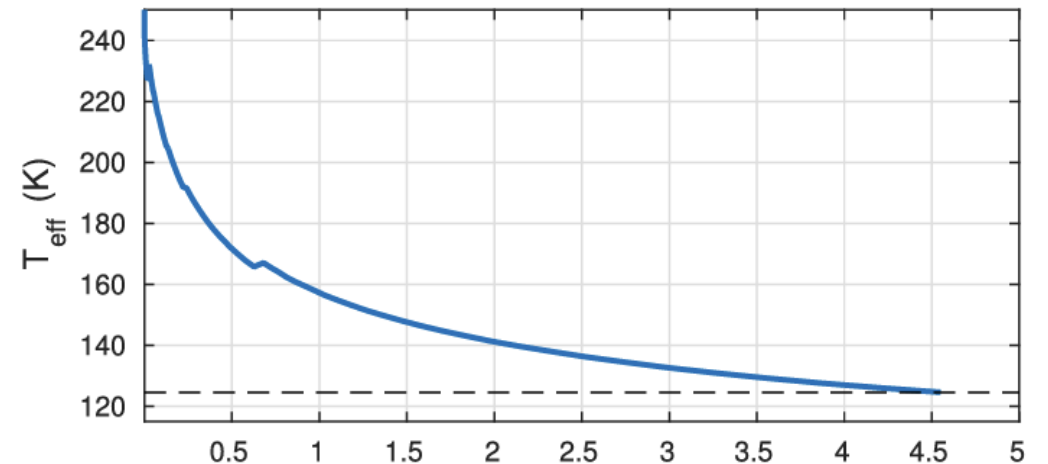
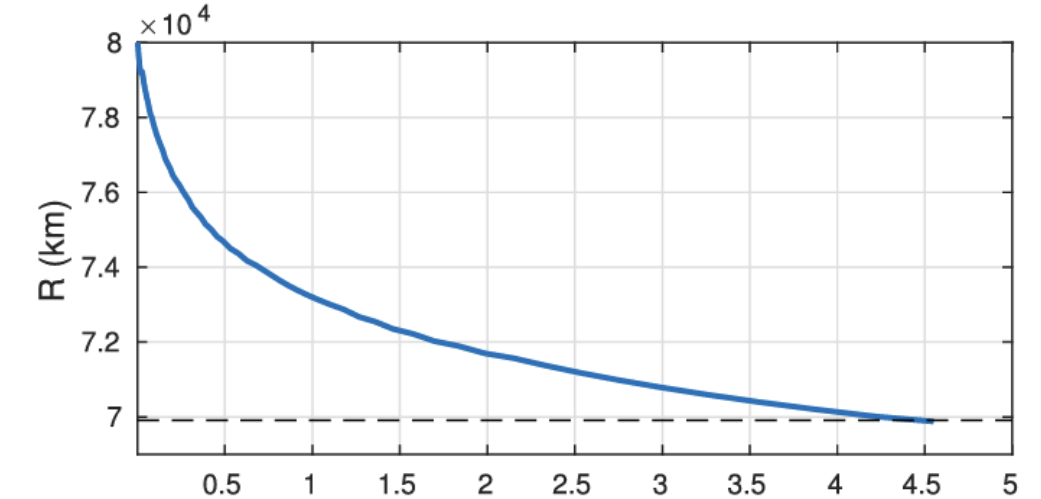
Jupiter Structure and Evolution: Non-adiabatic and Inhomogeneous Evolution

Vazan, Helled, Guillot (2018)

Evolution of Jupiter with composition gradients



Problem: No Helium Rain



Old Evolutionary Models

- Homogenous and Adiabatic
- Non-homogenous (helium rain) and Adiabatic
- Non-homogeneous (extended cores), non-Adiabatic, but no helium rain
- Old Equations of State (EOS)
- Old Atmospheric boundaries

Old Evolutionary Models

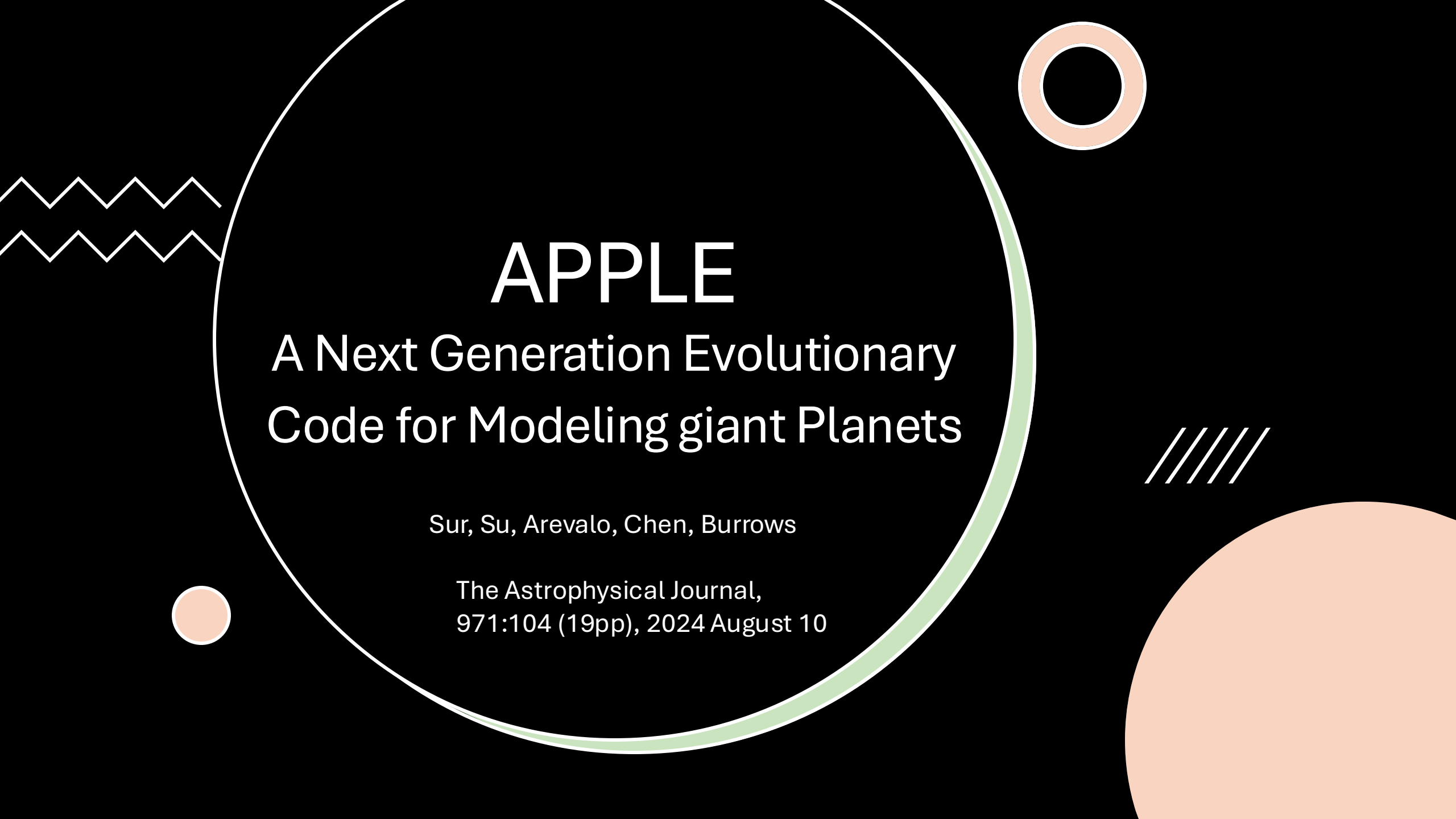


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New Evolutionary Models



- Non-homogenous: Constraints from Juno, and Cassini (fuzzy cores, Brunt), and helium rain,
- Heat Transport: Non-adiabatic (convection, radiation, and semi-convection)
- New H-He and heavy element EOSes
- New atmospheric boundary conditions



APPLE

A Next Generation Evolutionary Code for Modeling giant Planets

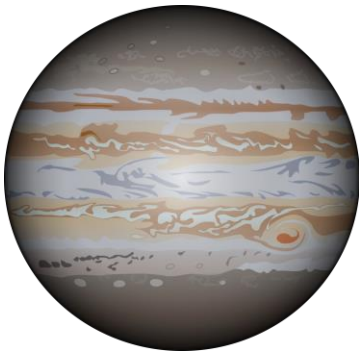
Sur, Su, Arevalo, Chen, Burrows

The Astrophysical Journal,
971:104 (19pp), 2024 August 10

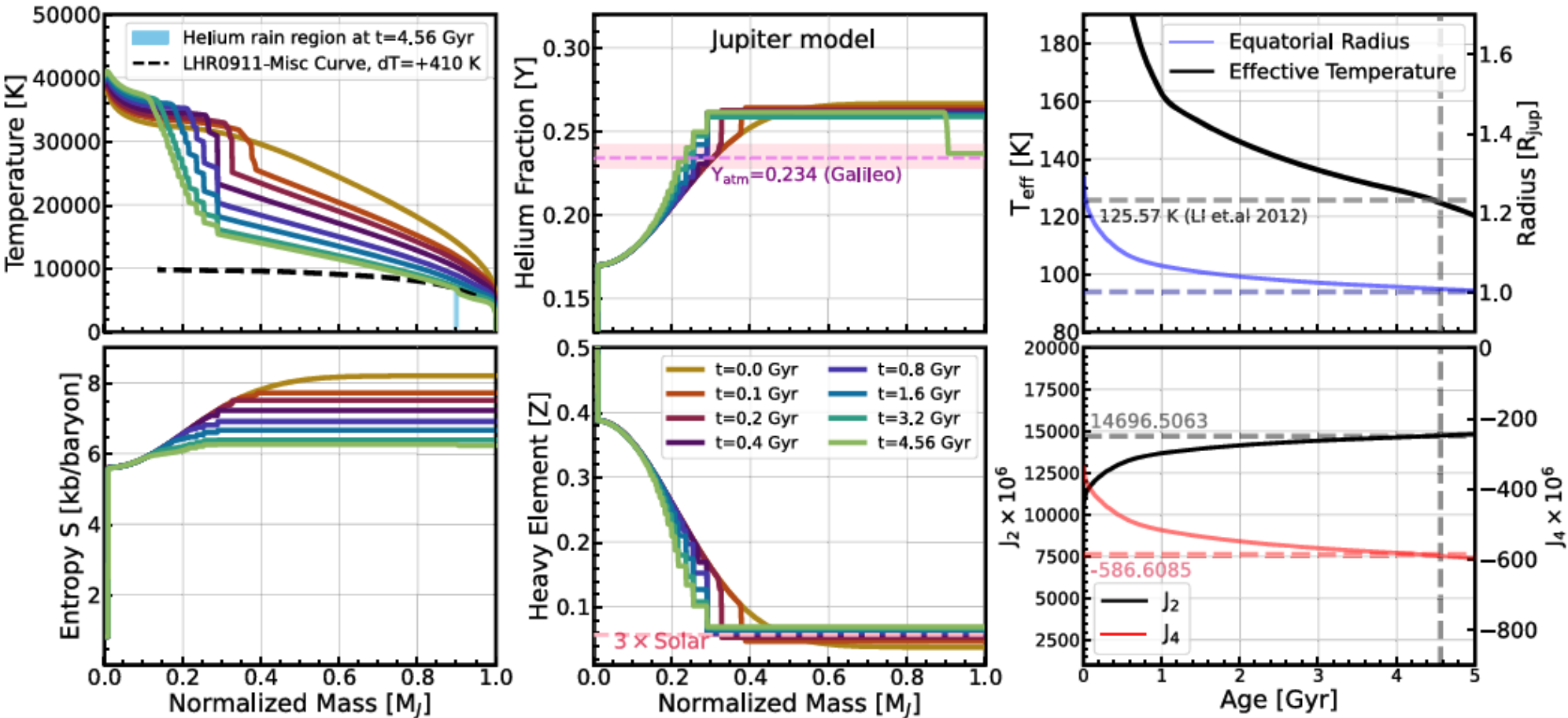
New Evolutionary Models for Jupiter and Saturn

Sur et.al ApJ 2025

Total heavy element mass
~ 42 Mearth (13%)



Quantity	Jupiter	
	Measured	Our Model (4.56 Gyr)
T_{eff} (K)	125.57 ± 0.07 (1)	124.6
Equatorial Radius (km)	71492 ± 4 (3)	72019.5
Atmospheric Helium Fraction (Y_{atm})	0.234 ± 0.005 (4)	0.236
Atmospheric Metallicity (Z/Z_{\odot})	1.5–5 (6)	3.6
$J_2 \times 10^{-6}$	14696.572 (7)	14731.6
$J_4 \times 10^{-6}$	-586.609 (7)	-591.46



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Matches T_{eff}

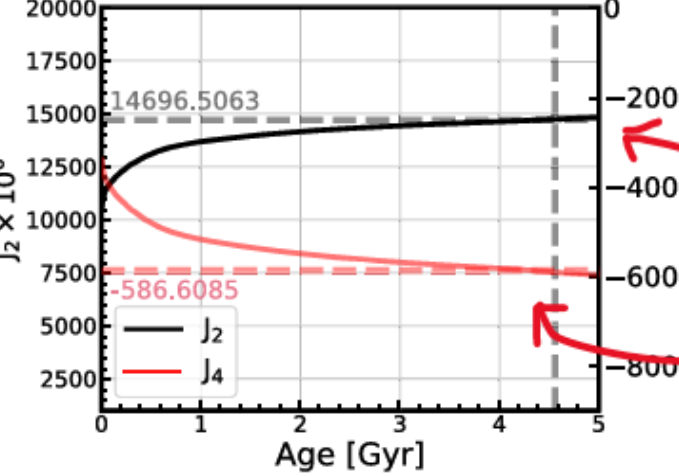
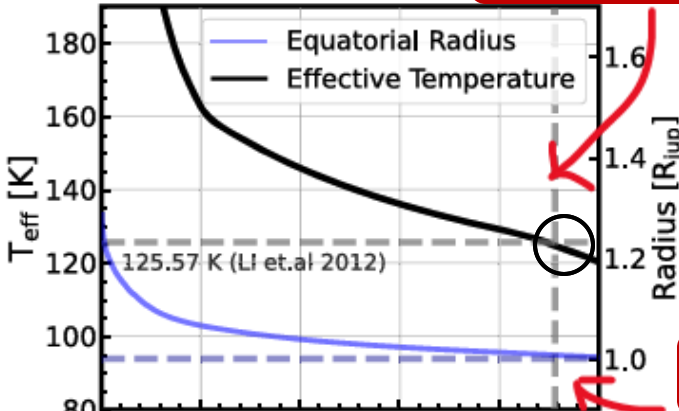
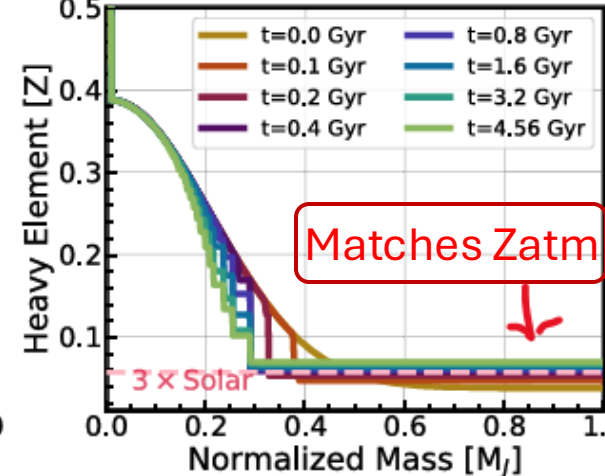
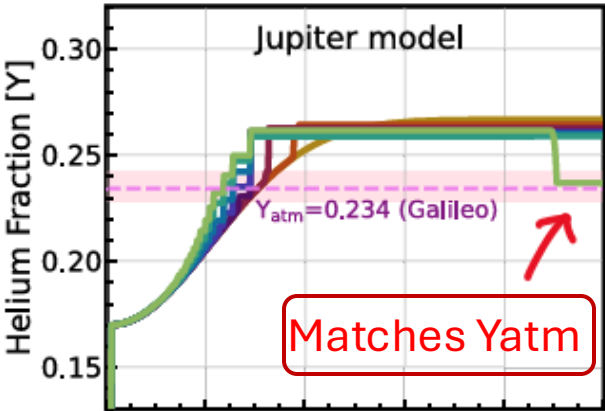
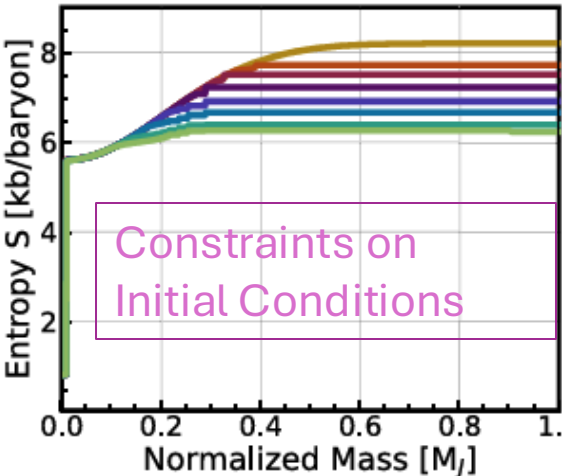
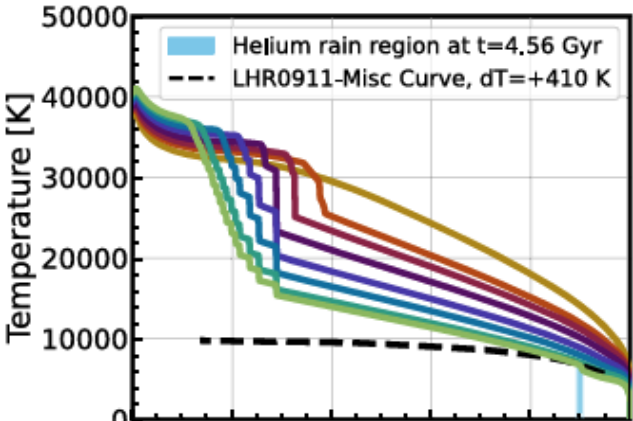
Matches R_{eq}

Matches Y_{atm}

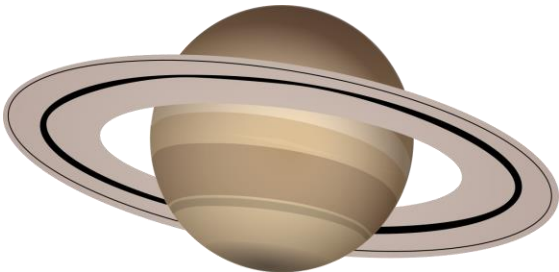
Matches Z_{atm}

Matches J_2 and J_4

Cold-Start



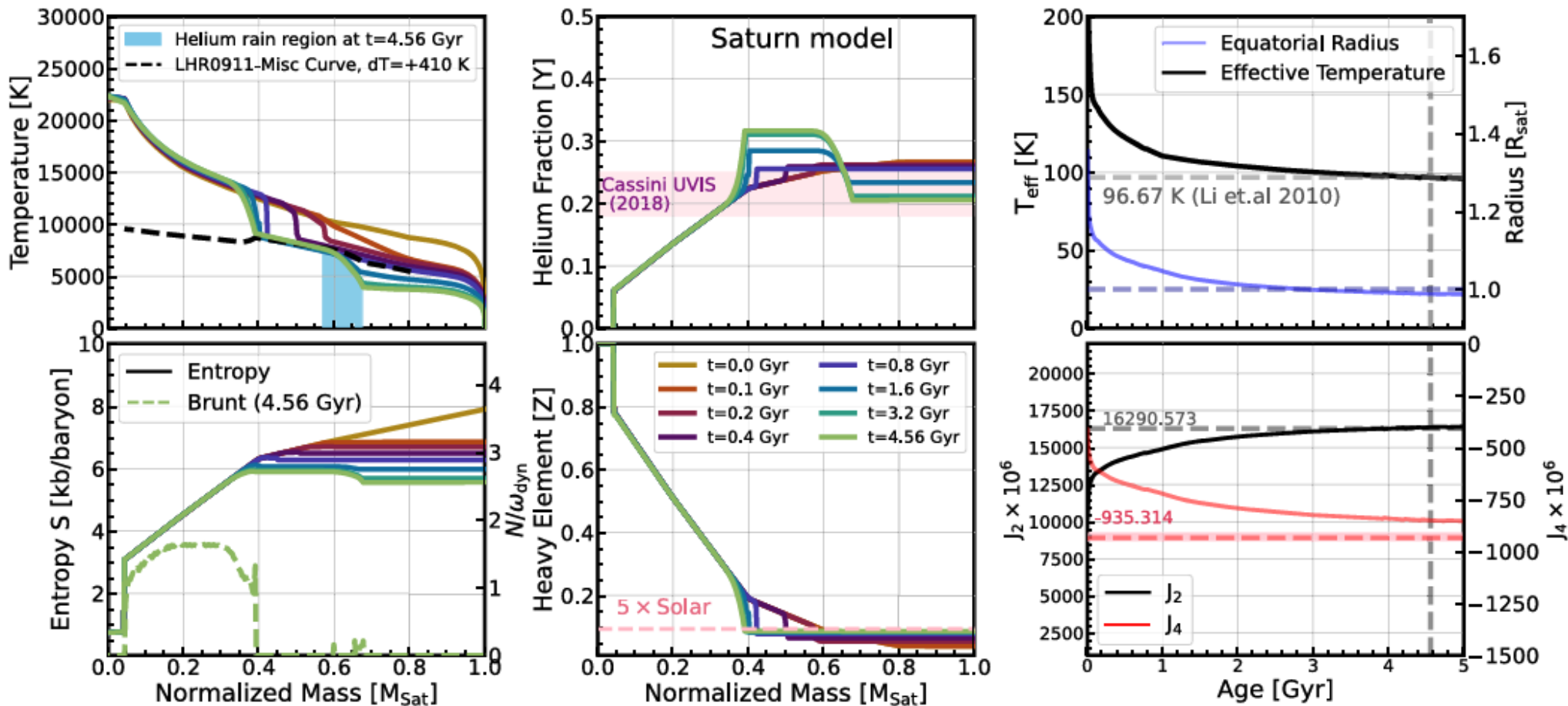
New Evolutionary Models for Jupiter and Saturn



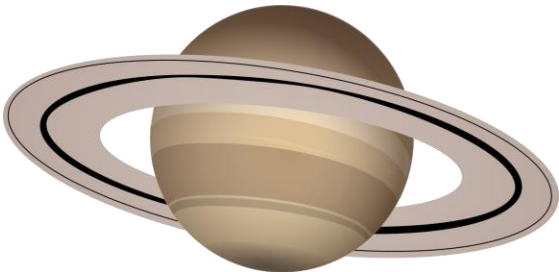
Sur et.al ApJ 2025

Quantity	Saturn	
	Measured	Our Model (4.56 Gyr)
T_{eff} (K)	96.67 (2)	96.54
Equatorial Radius (km)	60268 ± 4 (3)	59,551.8
Atmospheric Helium Fraction (Y_{atm})	0.075–0.22 (5)	0.205
Atmospheric Metallicity (Z/Z_{\odot})	5.0–10.0 (6)	4.6
$J_2 \times 10^{-6}$	16290.573 (8)	16365.7
$J_4 \times 10^{-6}$	−935.314 (8)	−850.11

Total heavy element mass
~ 25 Mearth (26%)



New Evolutionary Models for Jupiter and Saturn

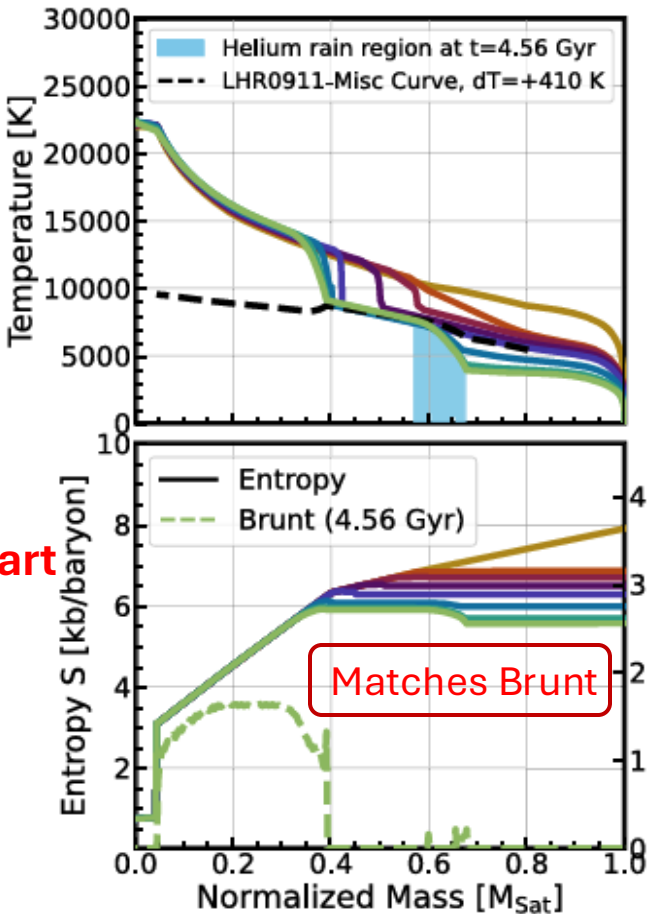


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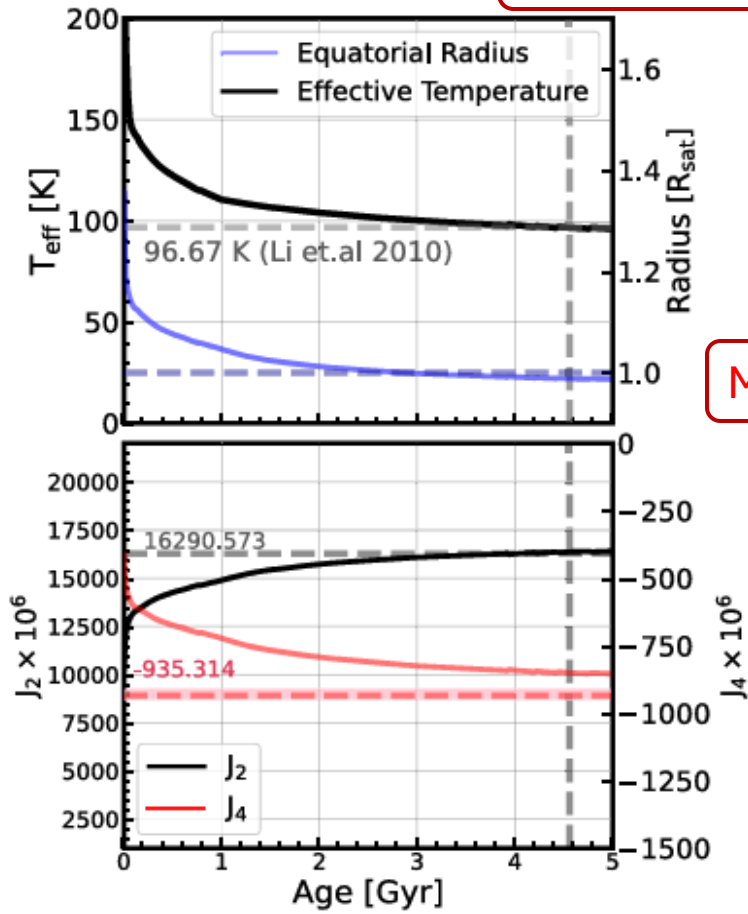
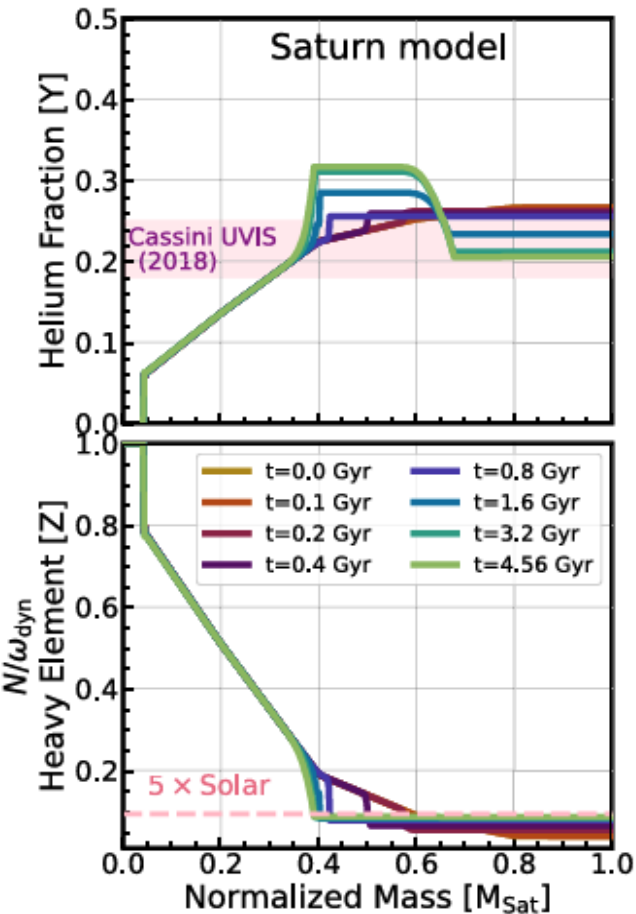
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Matches Teff



Cold-Start

Matches Brunt



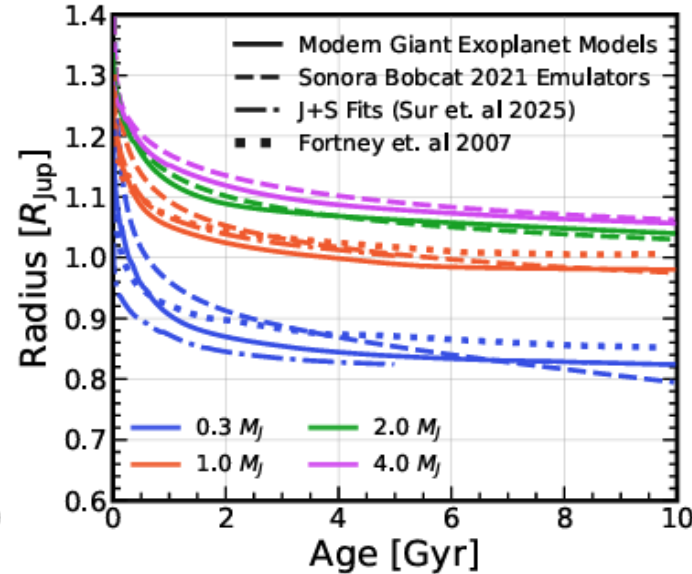
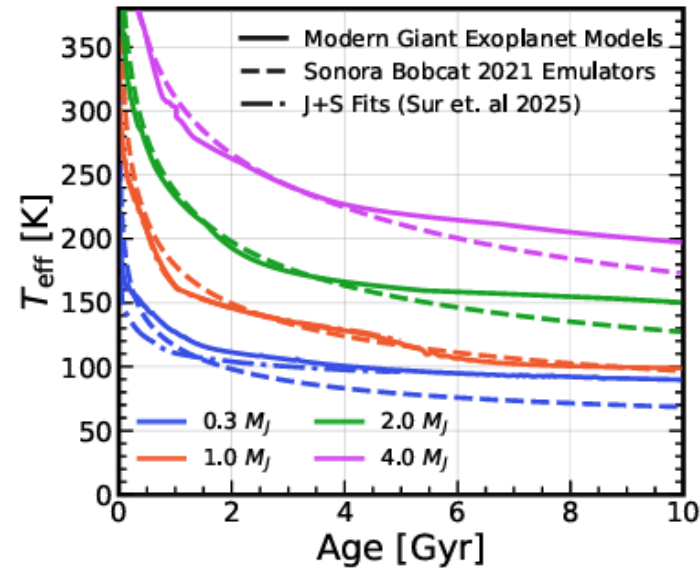
Matches R_{eq}

Matches J_2

Not J_4

New Evolutionary Models for Giant Exoplanets

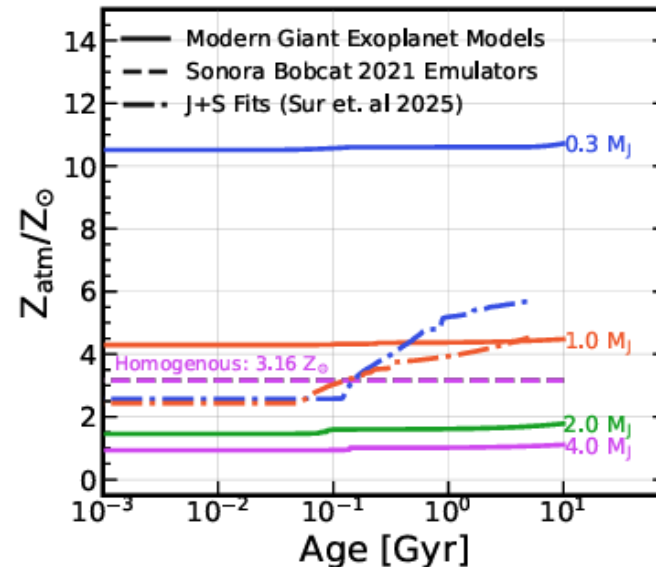
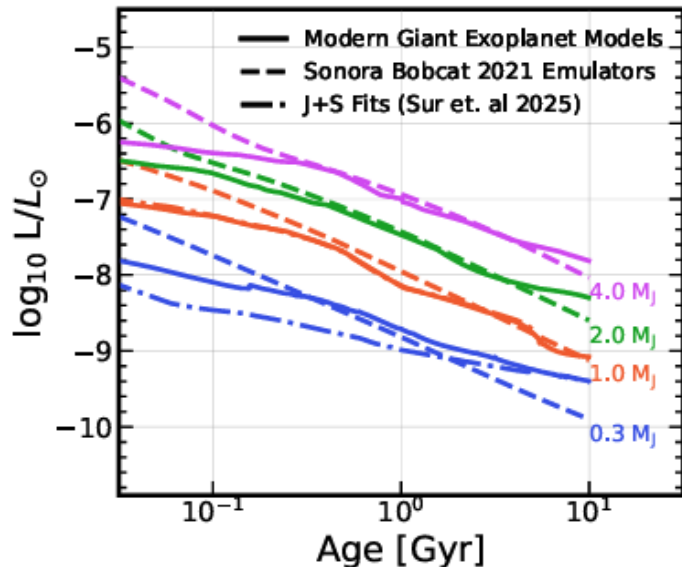
Sur et.al (under review ApJ, 2025)



Comparing with **Sonora Bobcat Models**

Improvements:

- Use real heavy element EOS
- Helium rain
- Non-adiabatic evolution
- Fuzzy cores



Results

- Predictions for Z_{atm}
- T_{eff} higher by 10-20%
- Radii differ by 5-10%



Conclusions and Summary

- Non-adiabatic models are required.
- Need improved EOSes and miscibility diagrams.
- Fuzzy cores vs Homogeneous.
- Cold-start initial conditions
- Saturn's $Y_{\text{atm}} \sim 0.2$