



Caltech



ExSoCal Conference  
September 18-19th 2017, Pasadena



# HIGH-CONTRAST IMAGING OF YOUNG PLANETS WITH JWST

Marie Ygouf



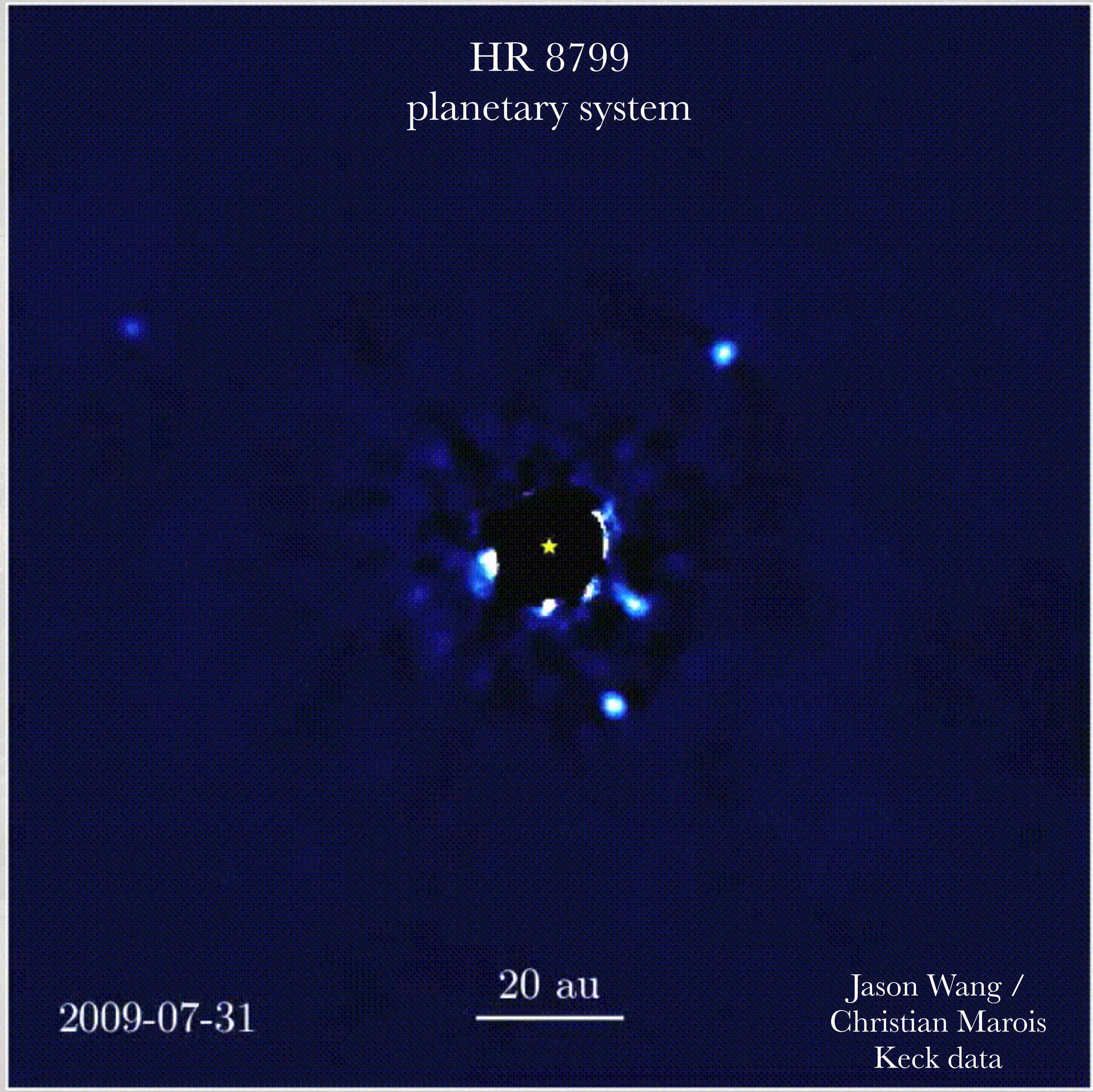
HR 8799  
planetary system

2009-07-31

20 au

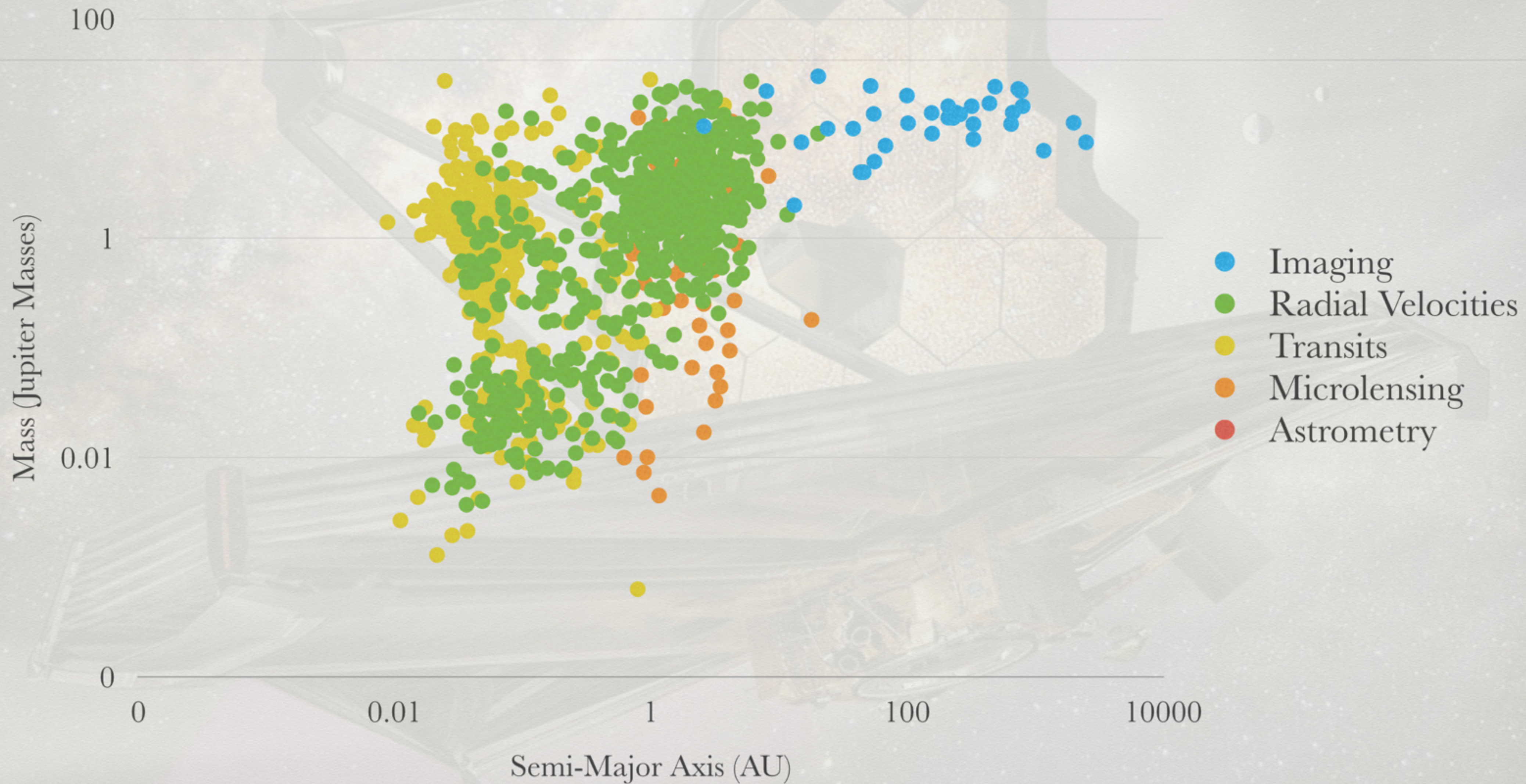
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Jason Wang /  
Christian Marois  
Keck data





# Confirmed planets\*

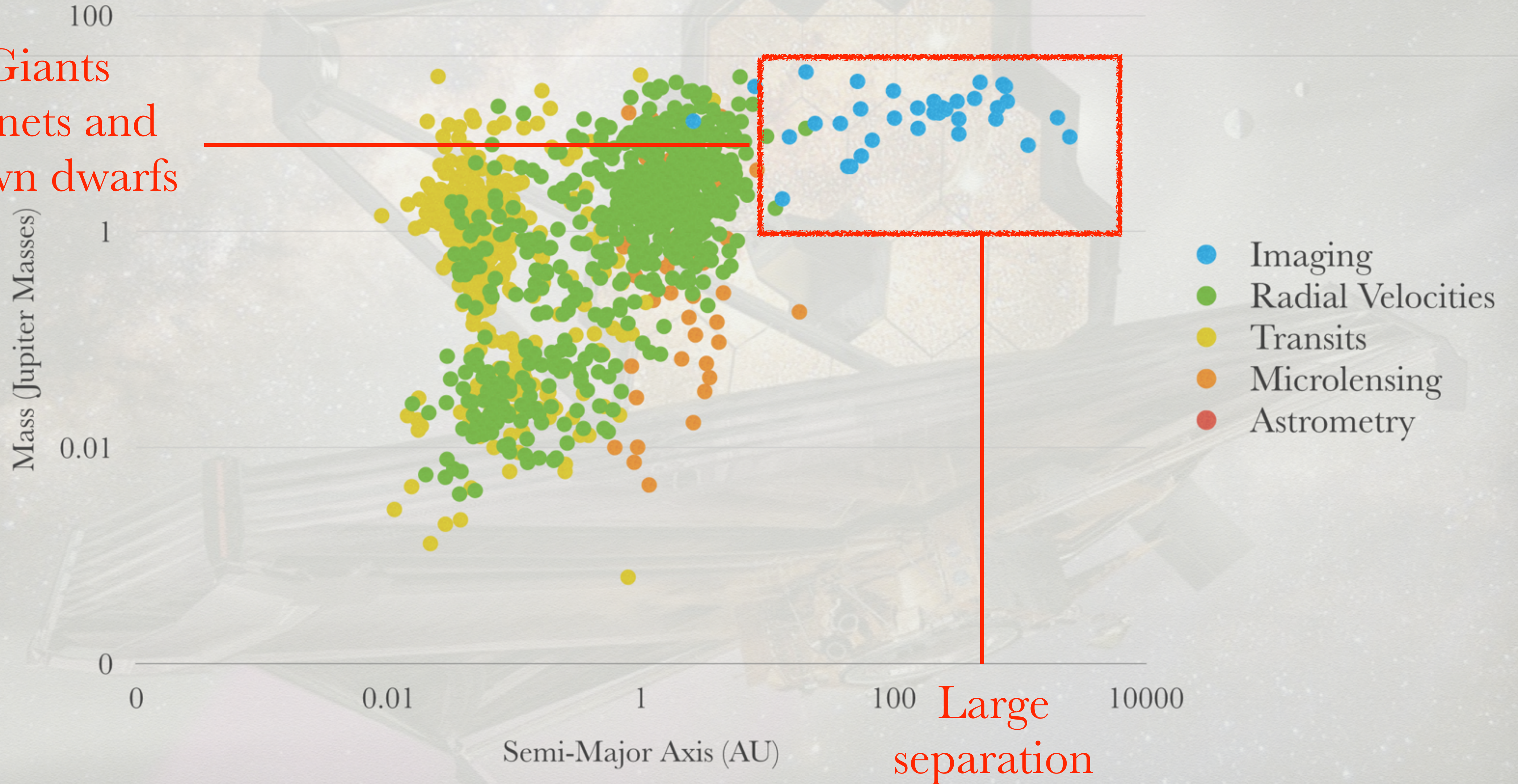


\*As 14 June 2017 from [exoplanetarchive.ipac.caltech.edu](http://exoplanetarchive.ipac.caltech.edu)



# Confirmed planets\*

Giants  
planets and  
brown dwarfs



\*As 14 June 2017 from [exoplanetarchive.ipac.caltech.edu](http://exoplanetarchive.ipac.caltech.edu)



# High contrast imaging science goals



## Characterization of known exoplanetary systems

- What is the occurrence of wide separated young exoplanets?
- What do they look like?
- How do they form?
- What is their evolutionary path?
- What are their interactions with the circumstellar disk?
- How can we put our solar system into context?



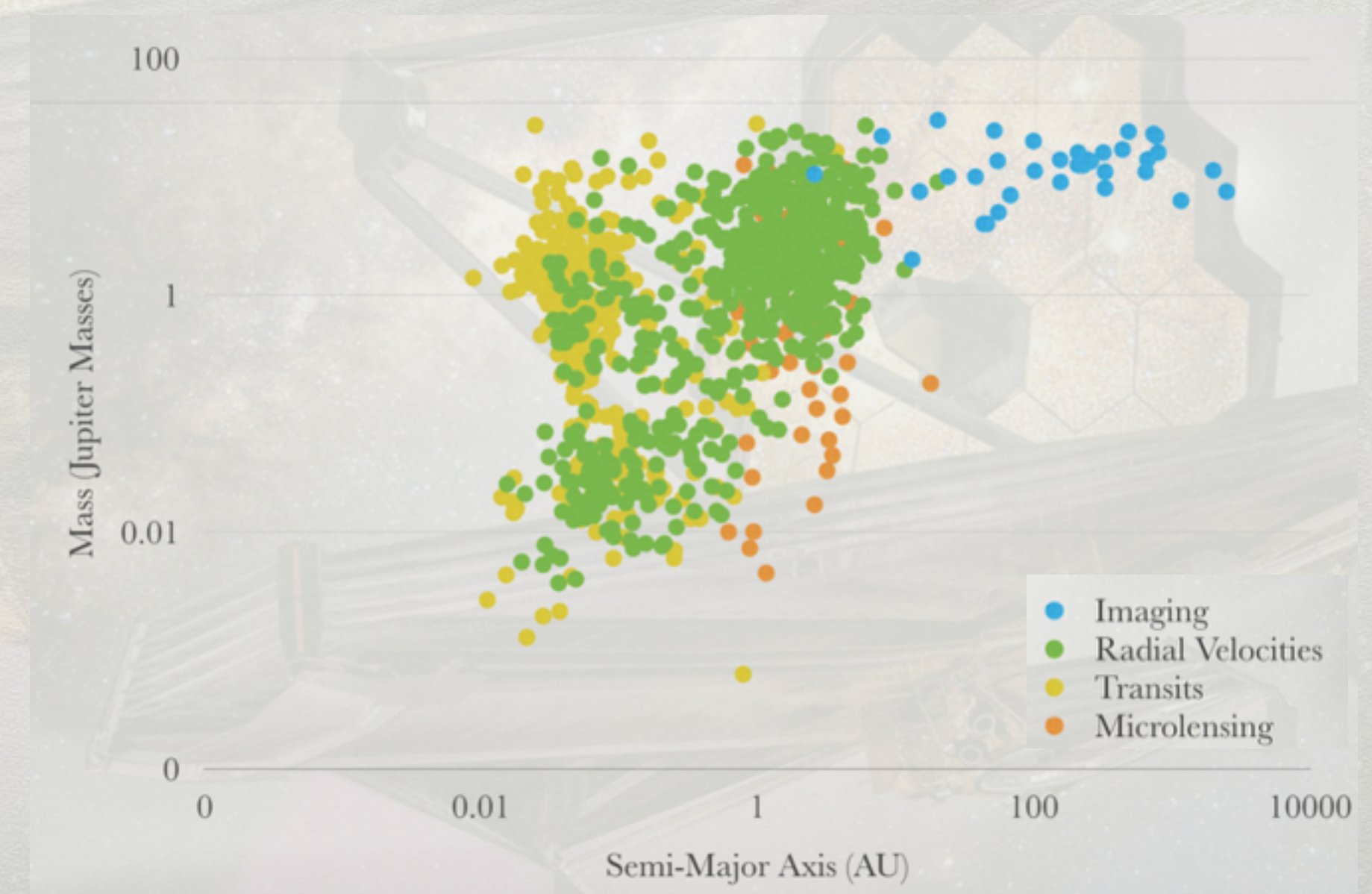
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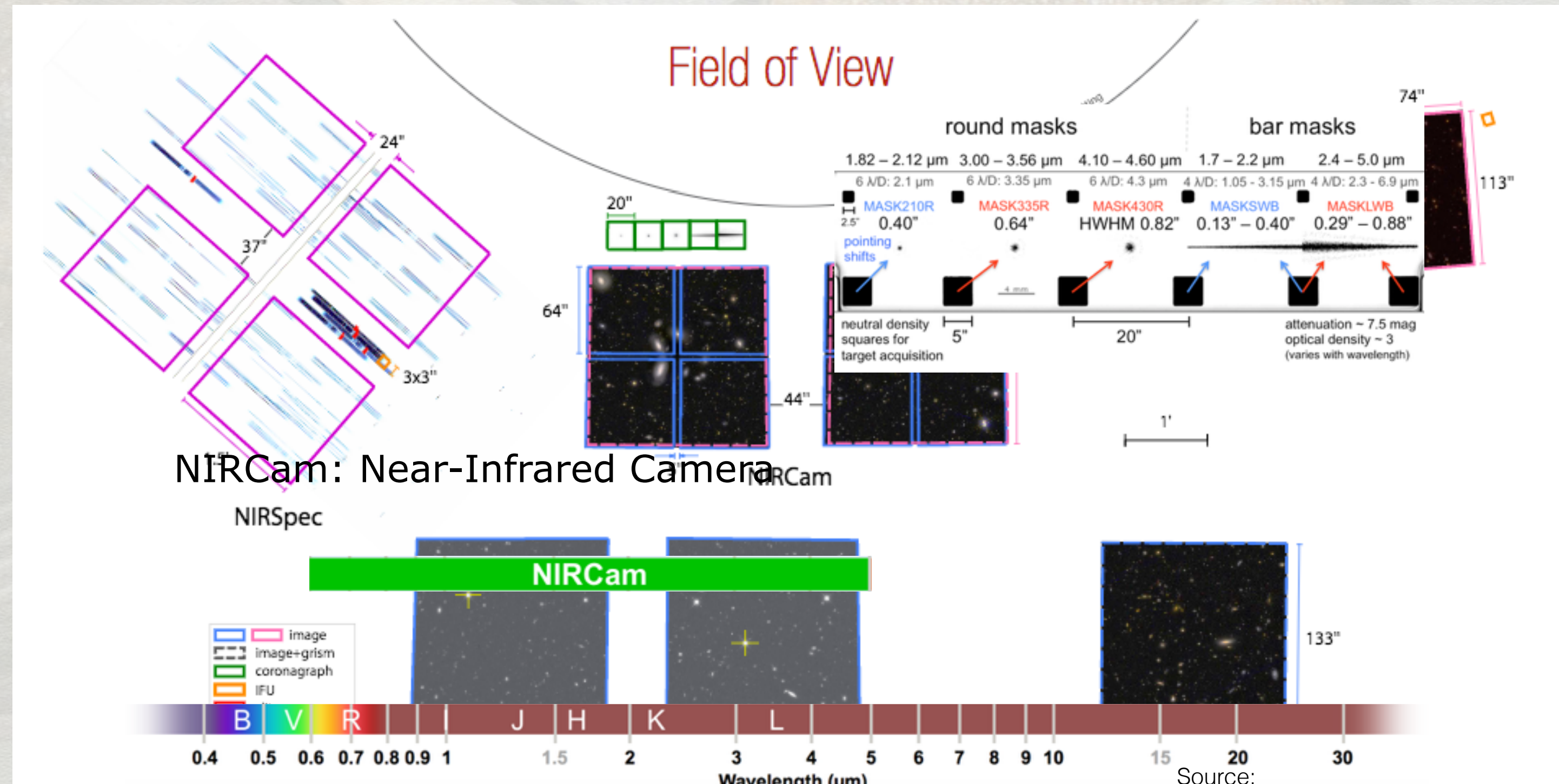
Search for new worlds

- How can we detect more of them?



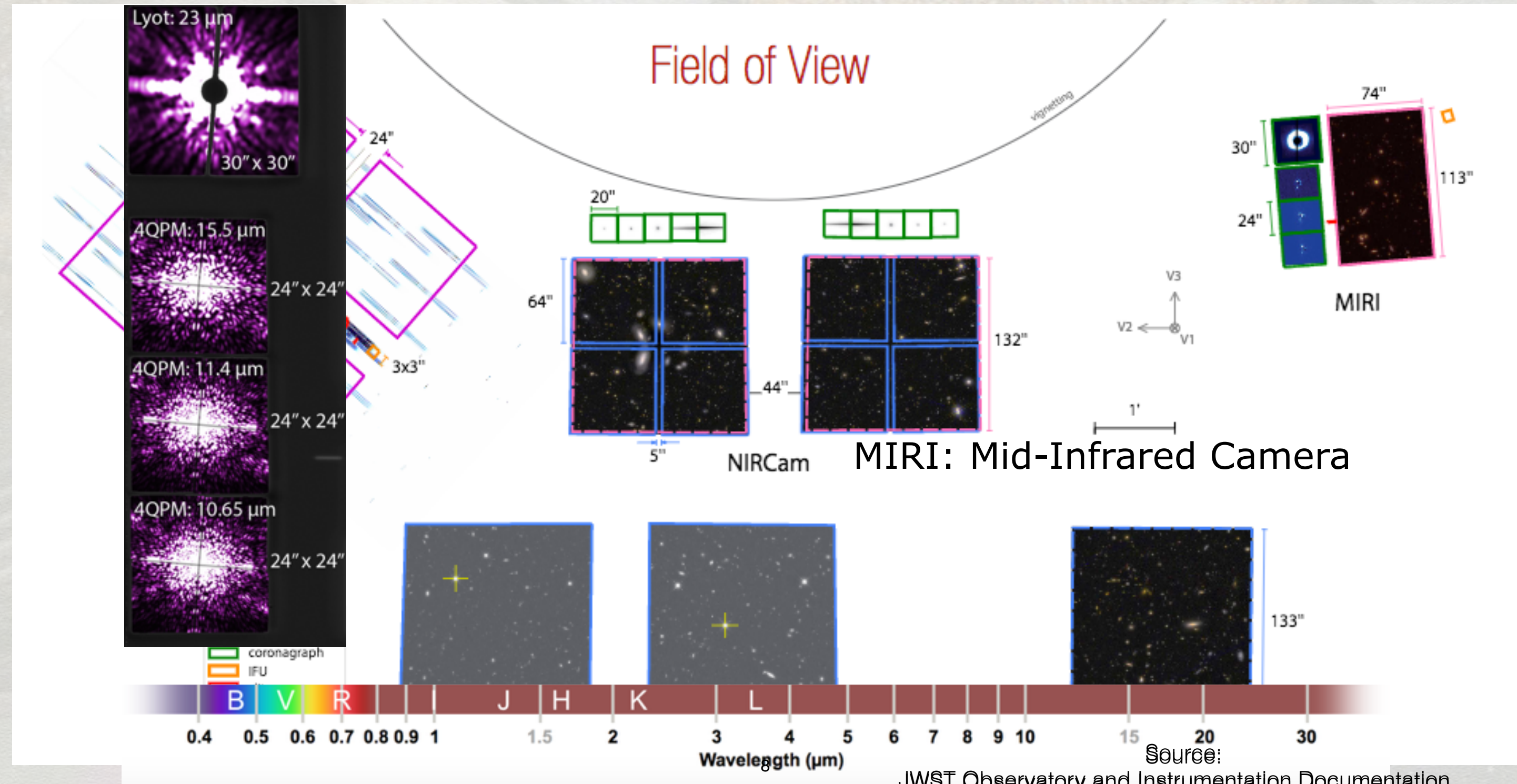


# JWST Coronagraphs





# JWST Coronagraphs





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Mass (Jupiter Masses)

100

1

0.01

0

0

0.01

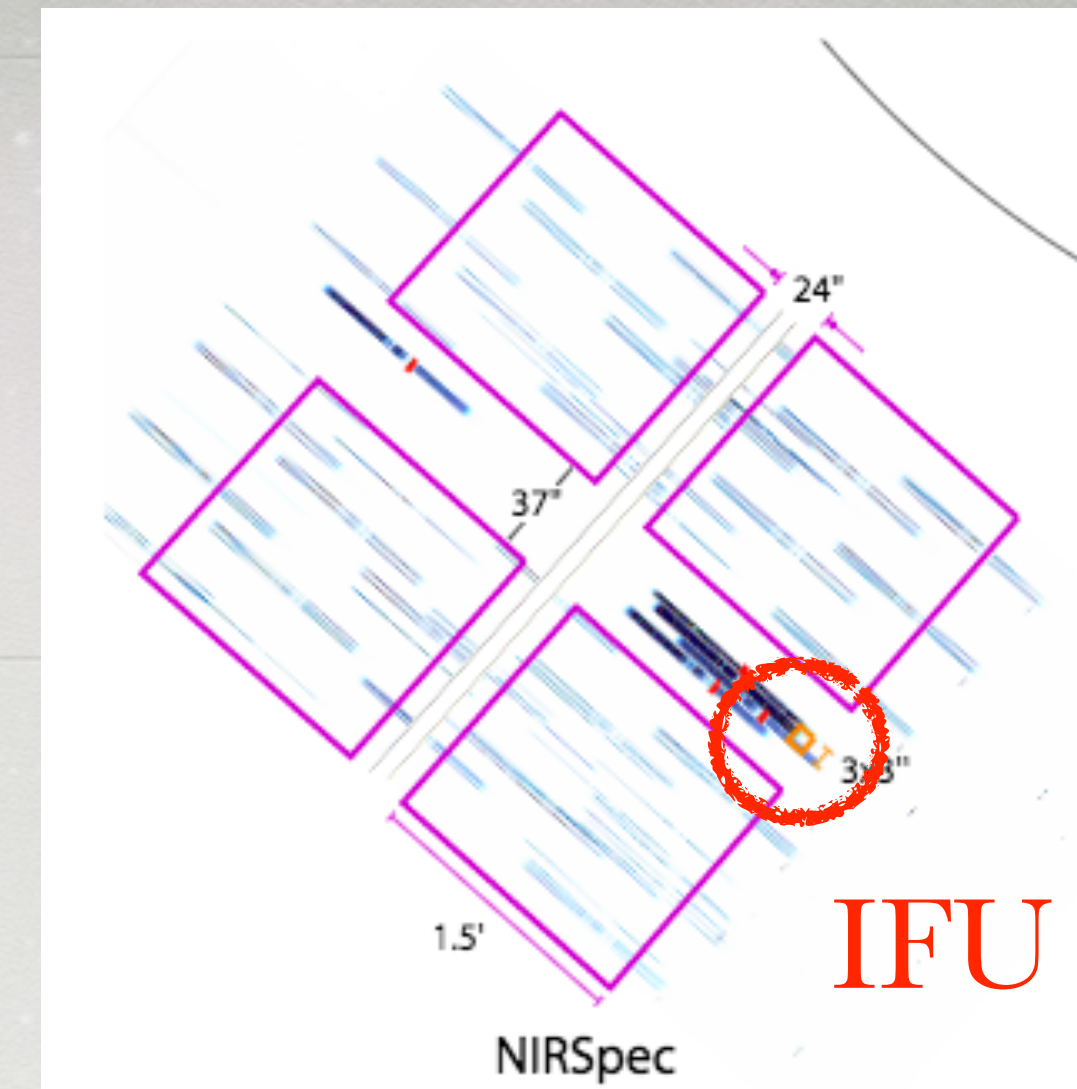
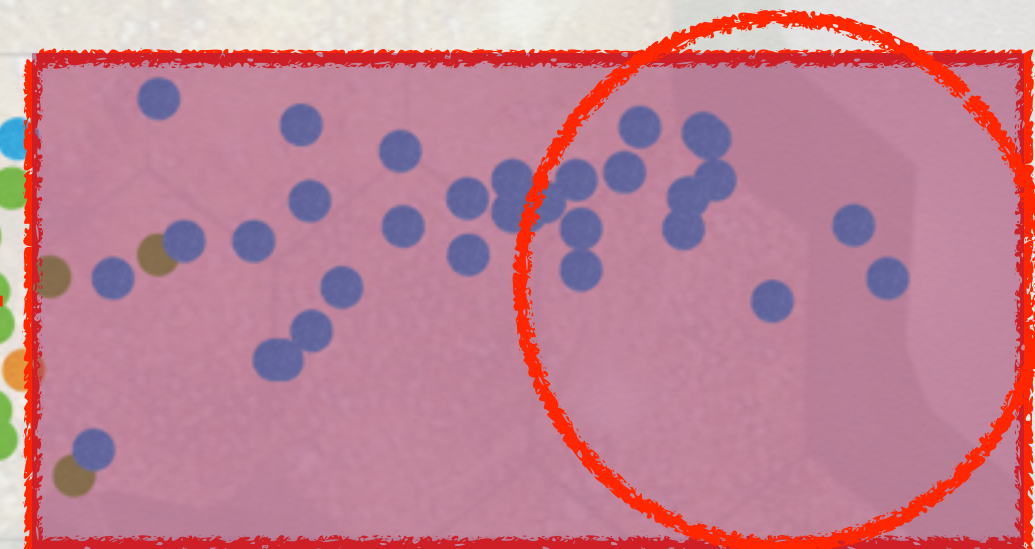
1

100

10000

Semi-Major Axis (AU)

Large  
separation



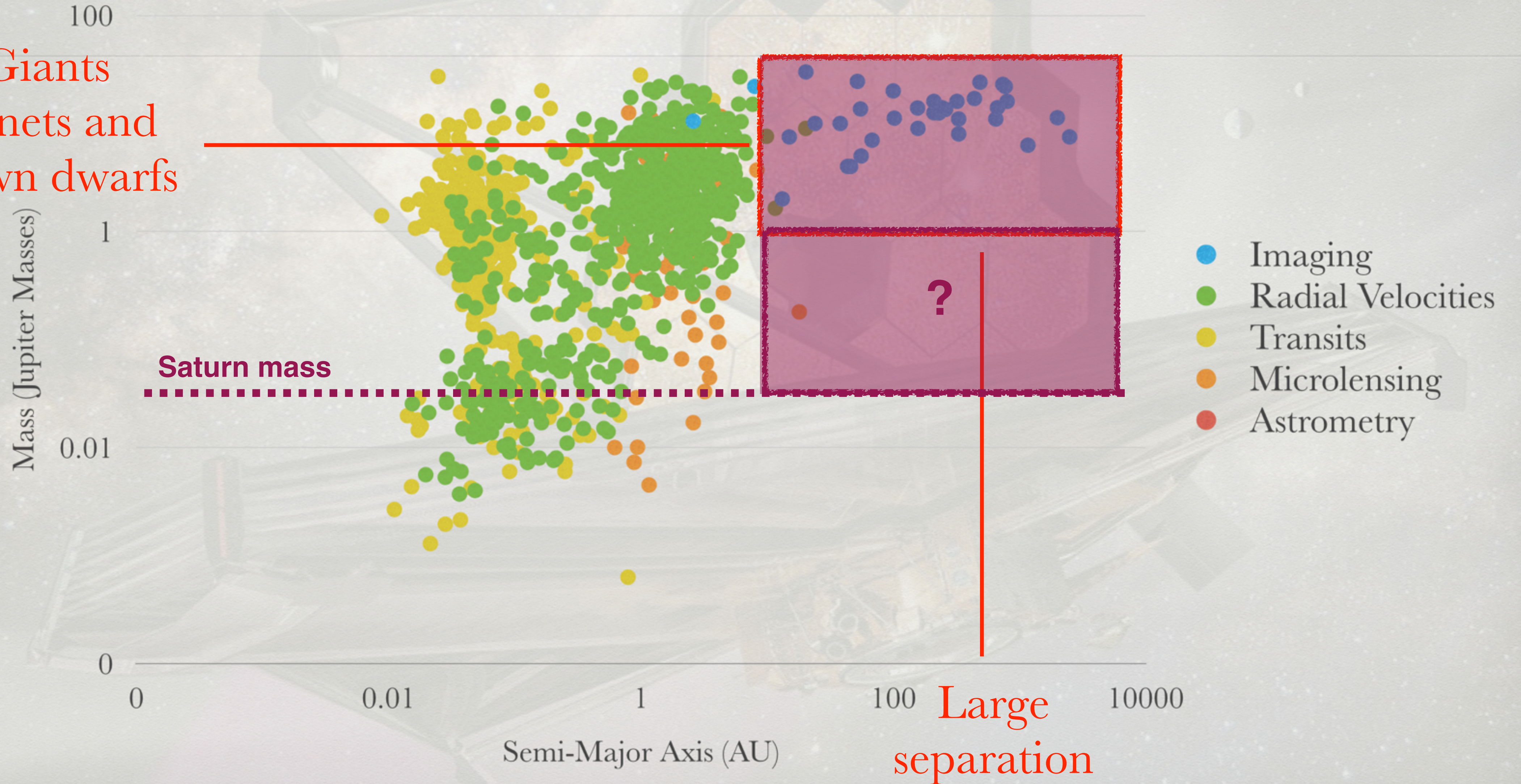
- Imaging
- Radial Velocities
- Transits
- Microlensing
- Astrometry

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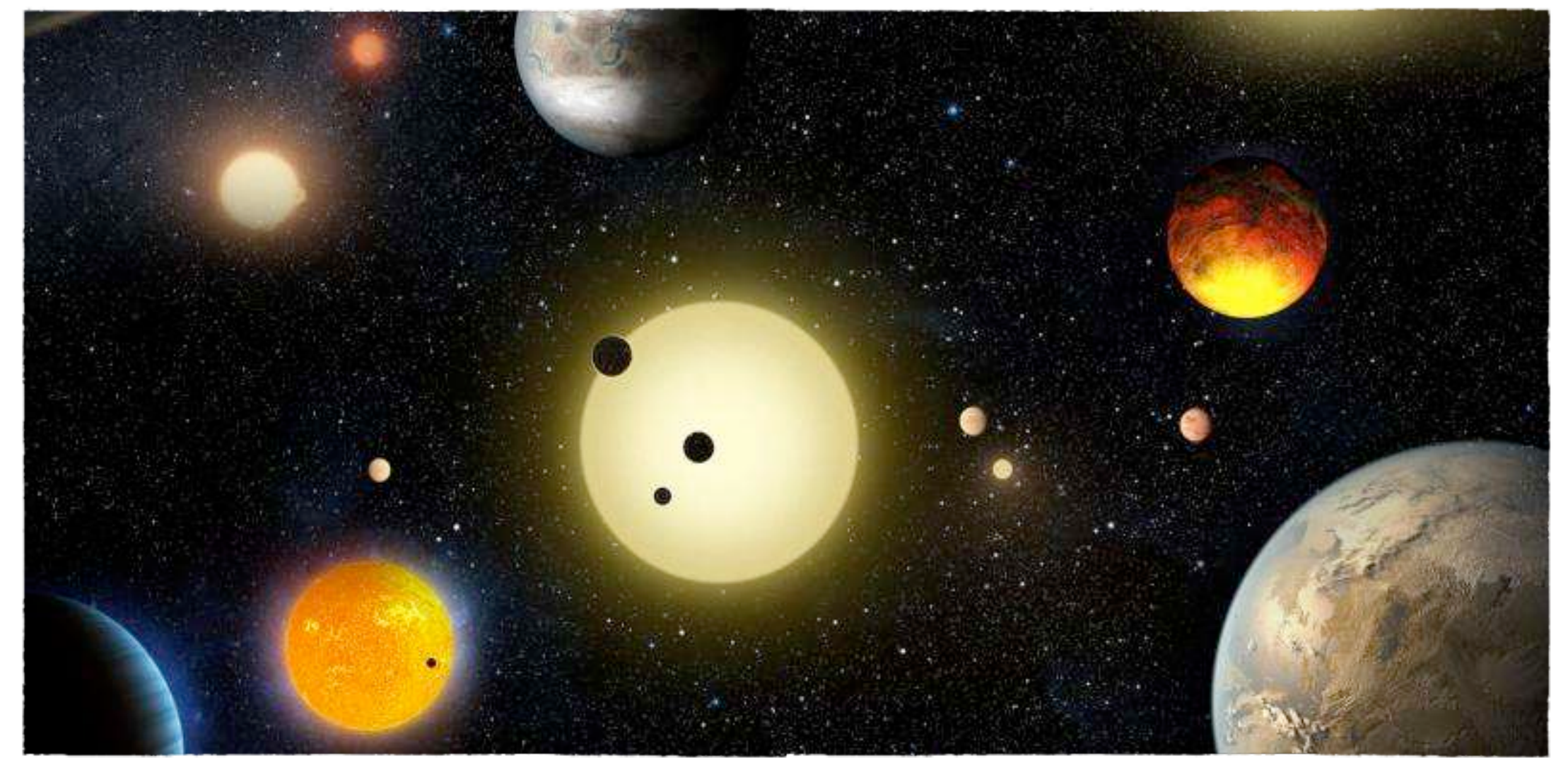
# Coronagraphic Imaging of Young Planets

## NIRCam GTO Program (PI: Chas Beichman)

Characterization of known  
exoplanetary systems



Search for new  
worlds





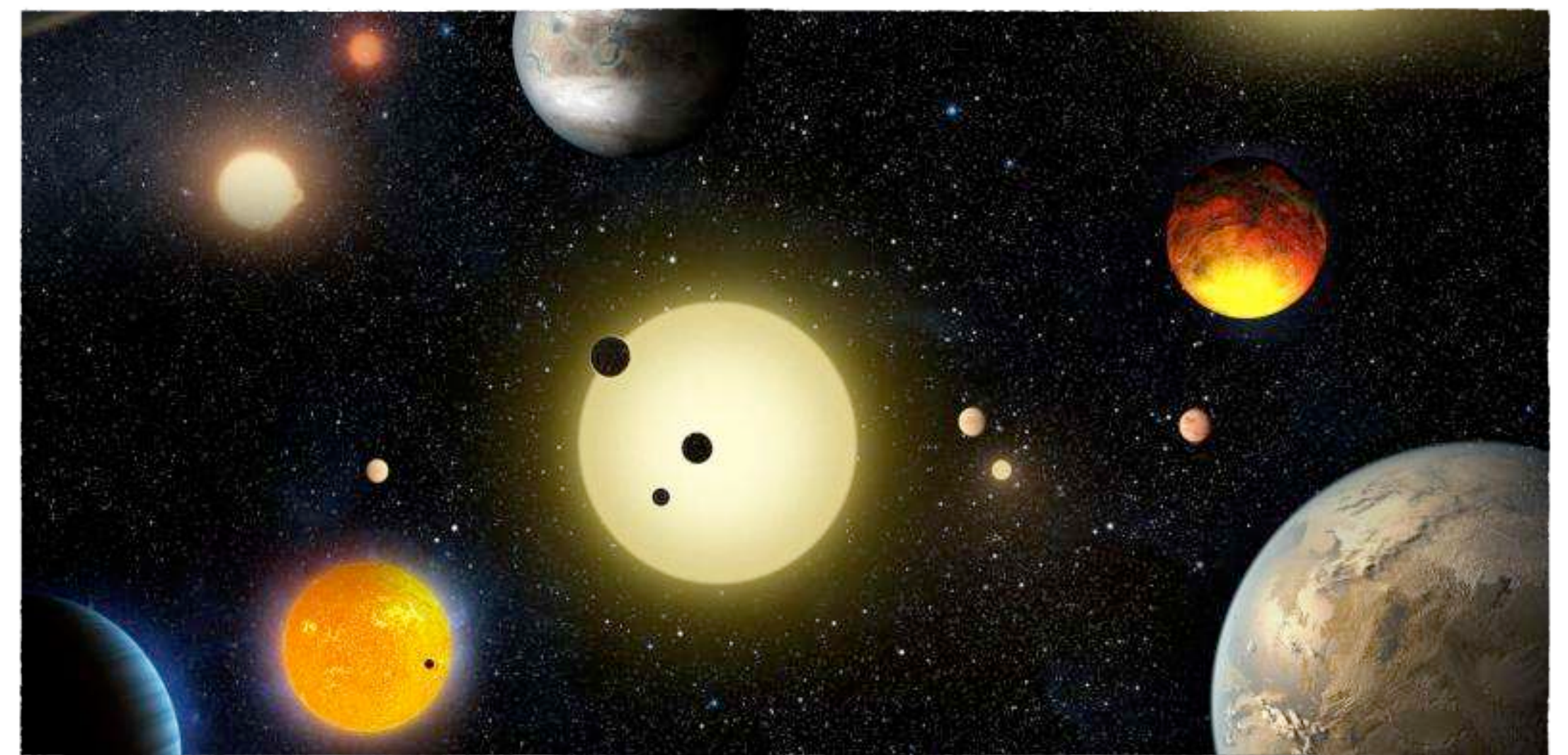
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# Characterization of known exoplanetary systems

- Determine atmospheric composition and physical properties to understand formation mechanism of exterior planets





# Characterization of known exoplanetary systems

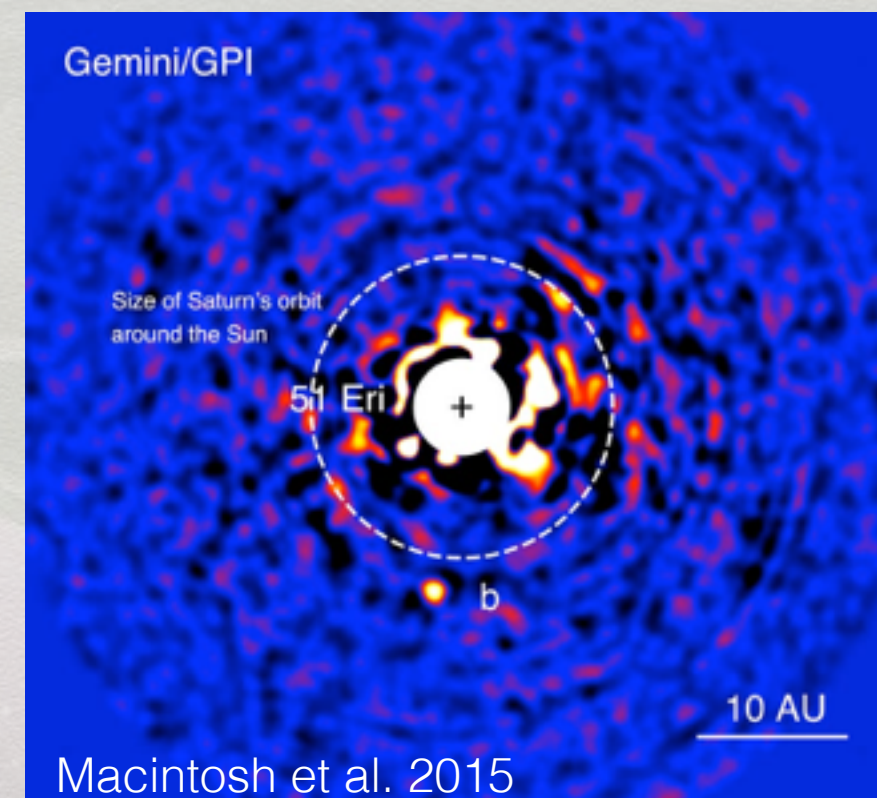
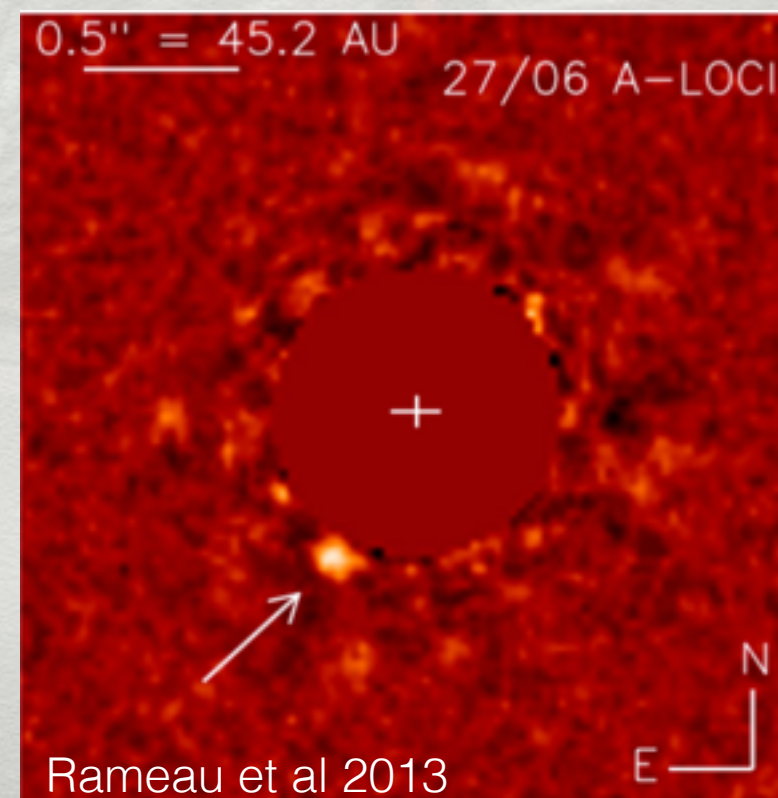
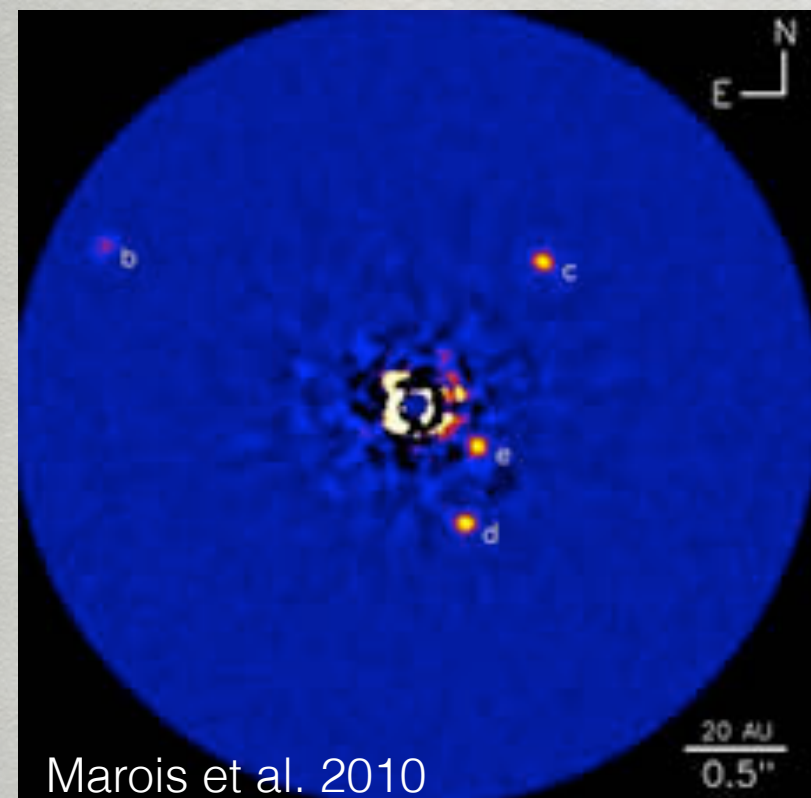
- Final Target List (as of 15 June 2017)

- HD 95086 b
- HR 8799 b, c & d
- HR 8799 e
- 51 Eri b

NIRCam & MIRI  
GTOs

Telescope Team GTO

- Determine atmospheric composition and physical properties to understand formation mechanism of exterior planets





# Characterization of known exoplanetary systems

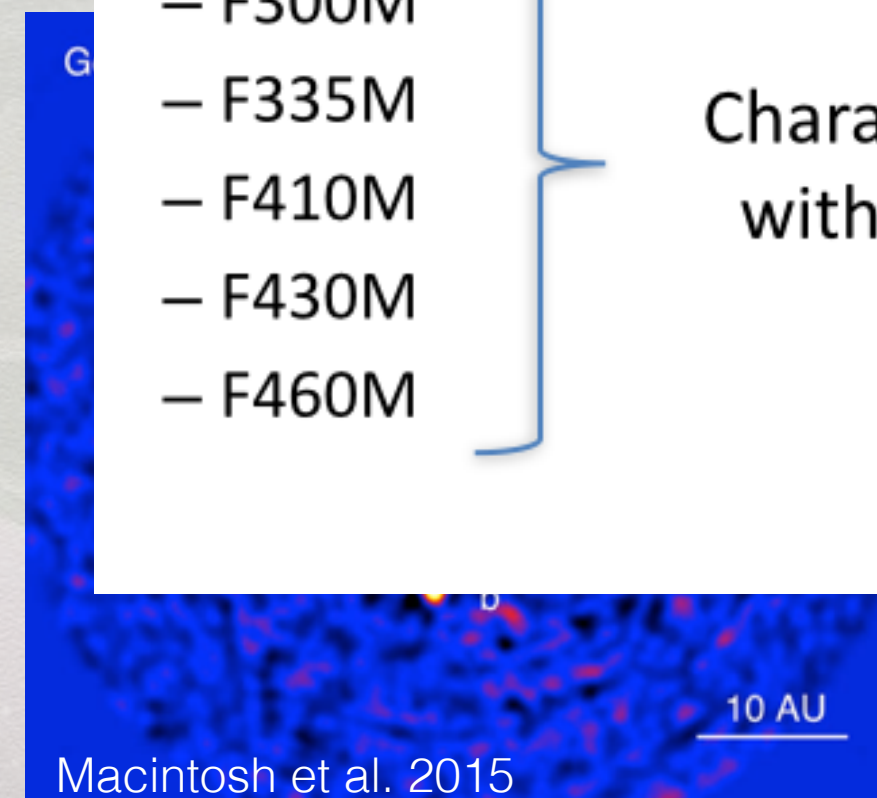
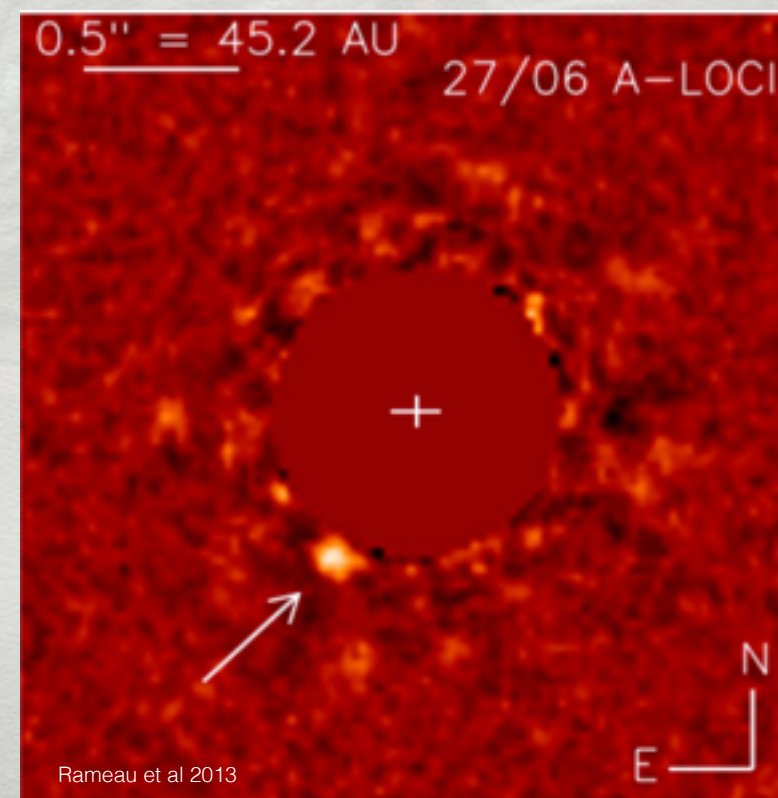
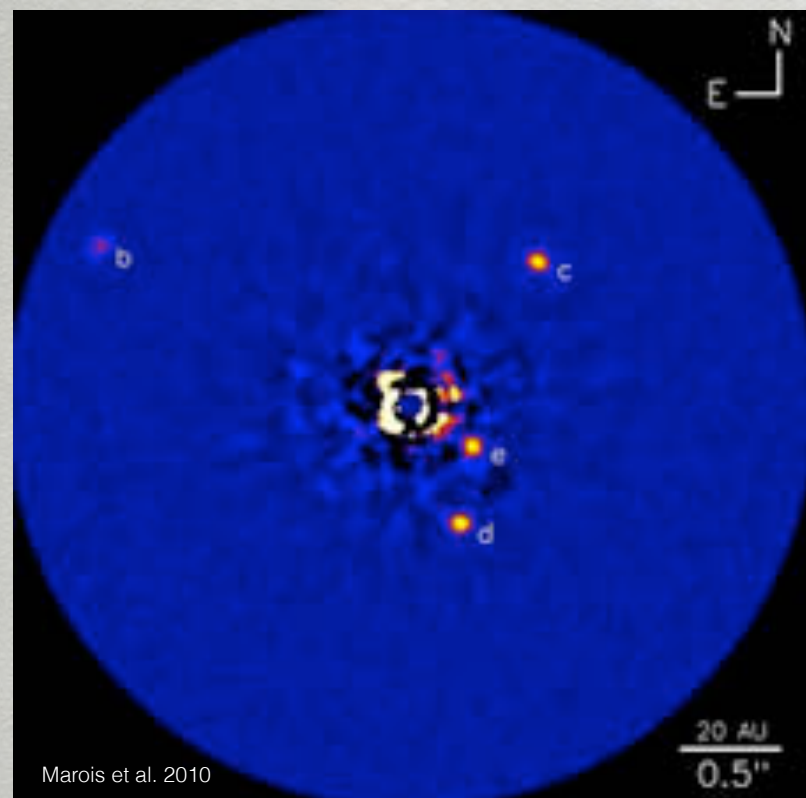
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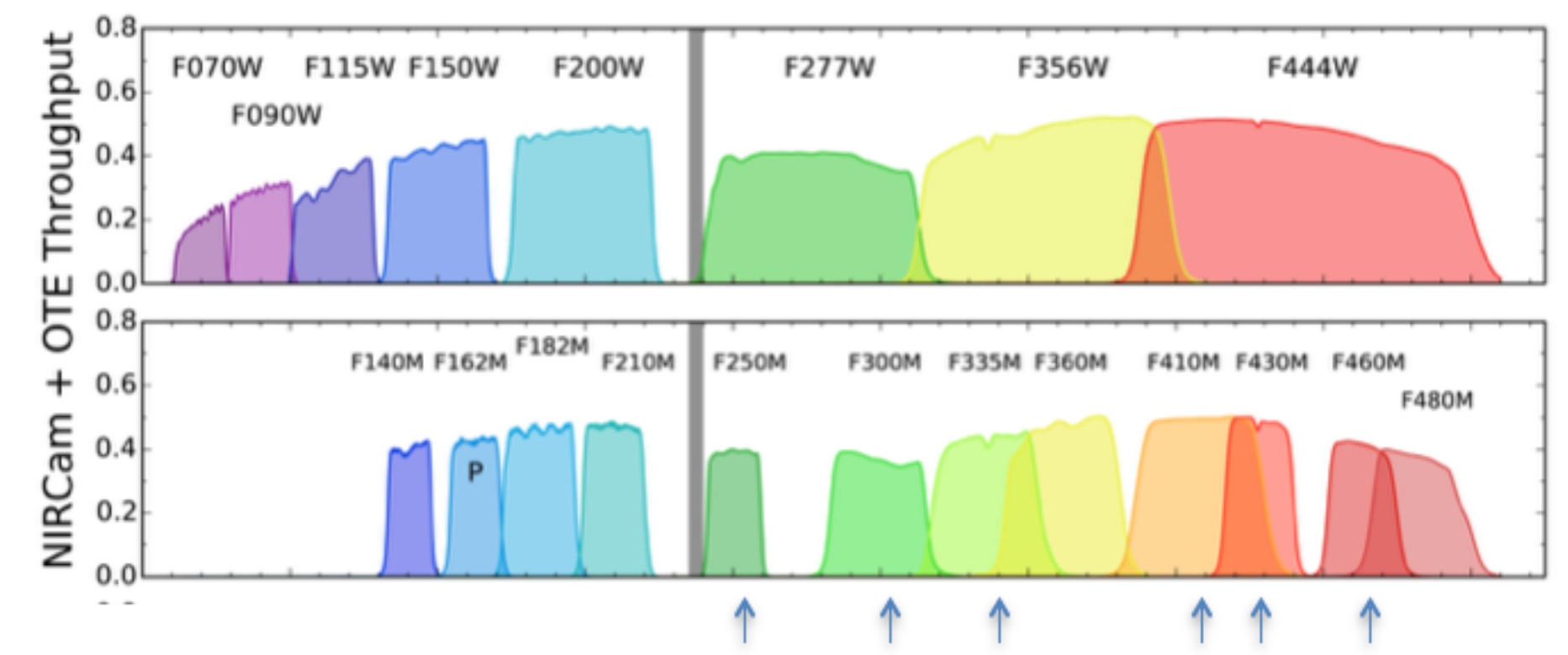
Telescope Team GTO

- Determine atmospheric composition and physical properties to understand formation mechanism of exterior planets
- Observing Strategy: Long Wavelength Mask, 6 filters, 1 roll, 1 reference star for speckle calibration



- F250M
- F300M
- F335M
- F410M
- F430M
- F460M

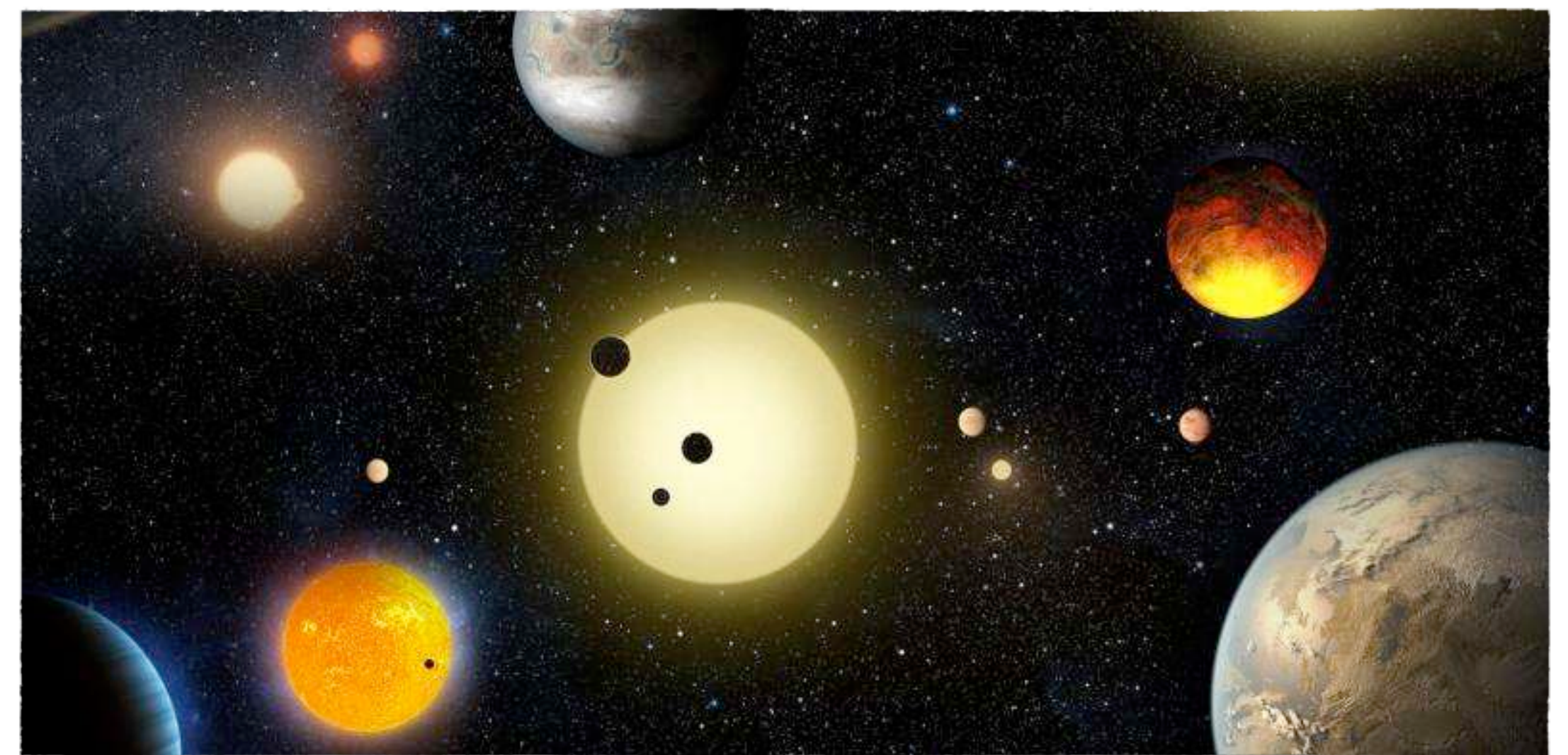
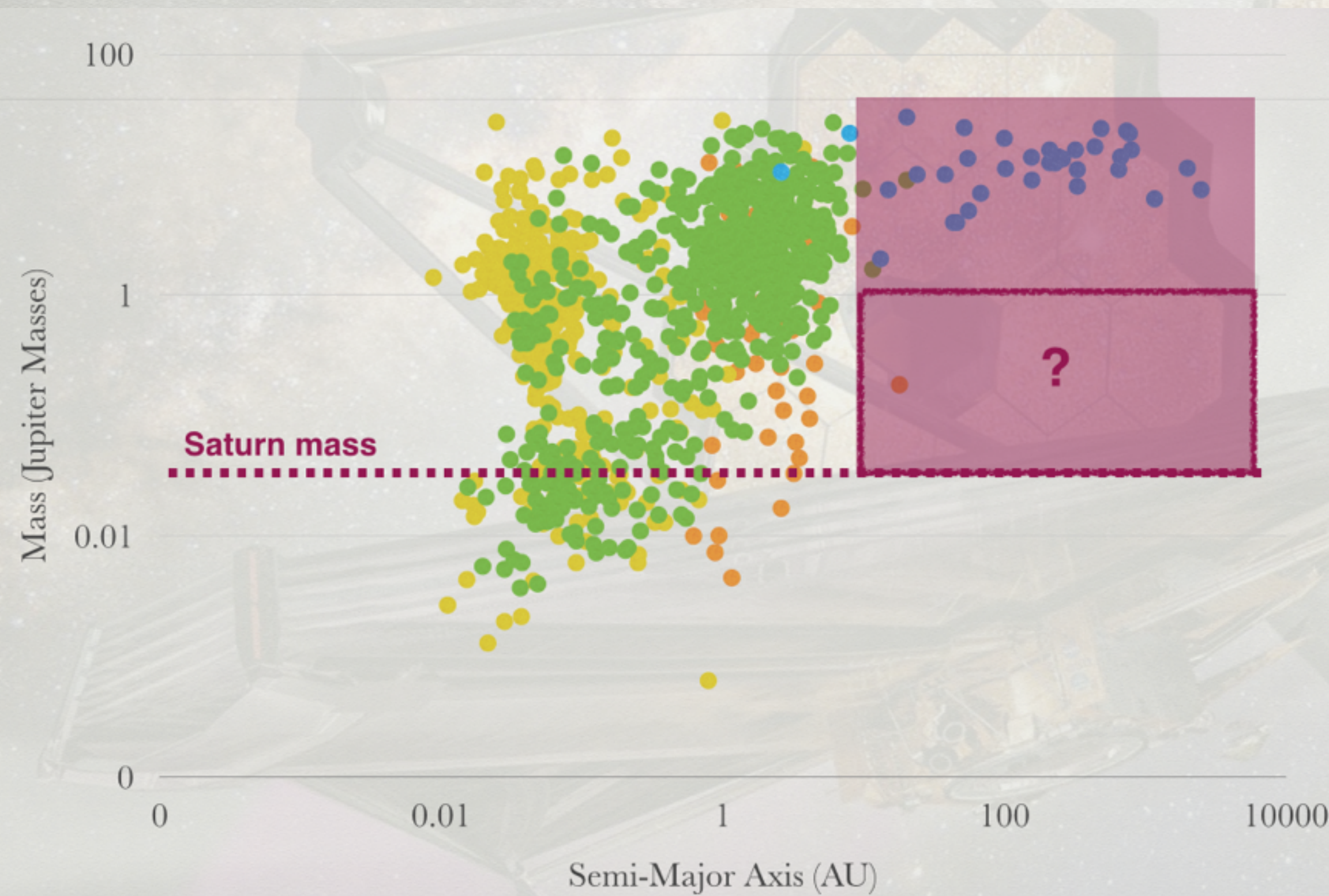
Characterization  
with MASKLW





# Search for new worlds

- Extend to lower mass limits compared to ground-based observations



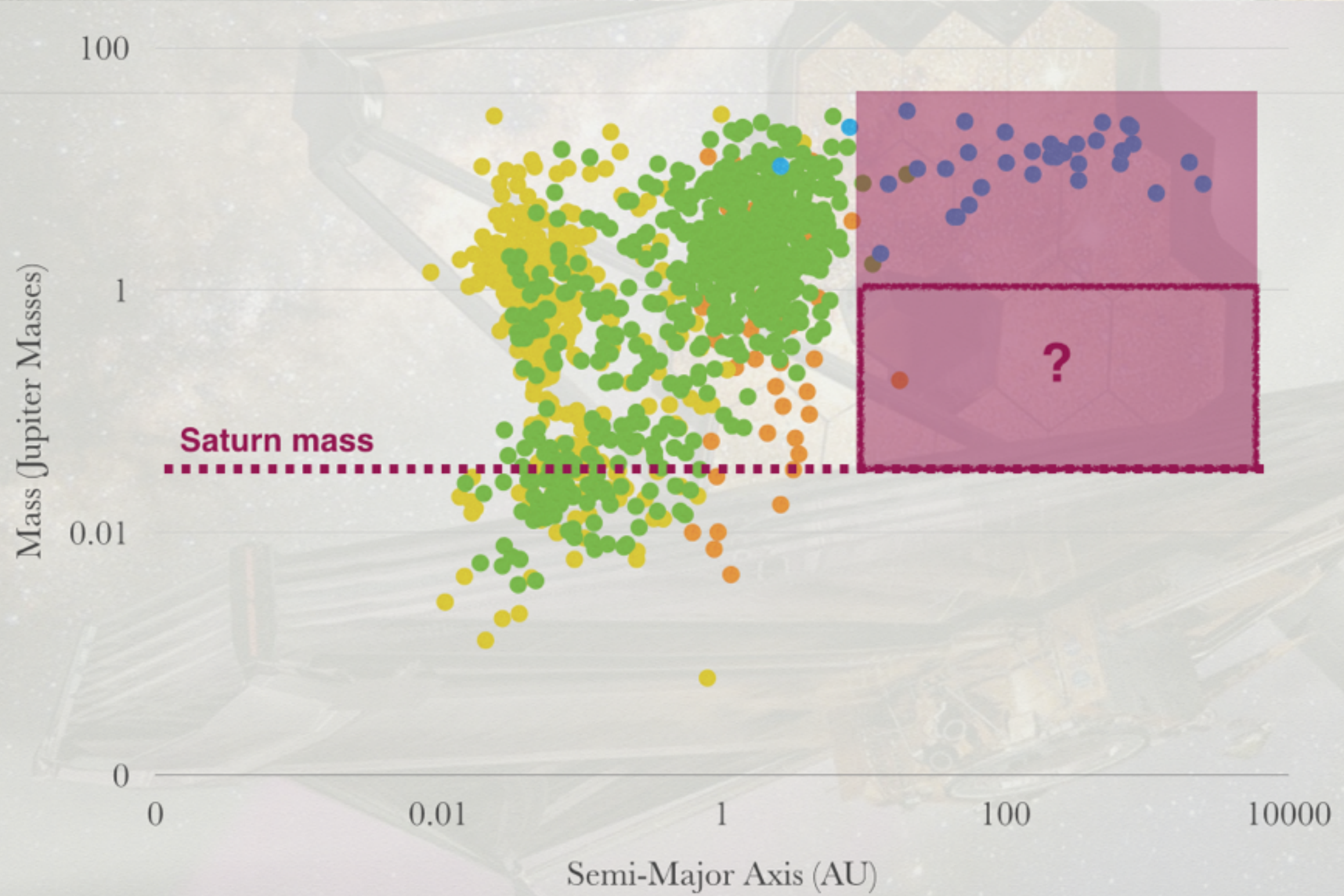
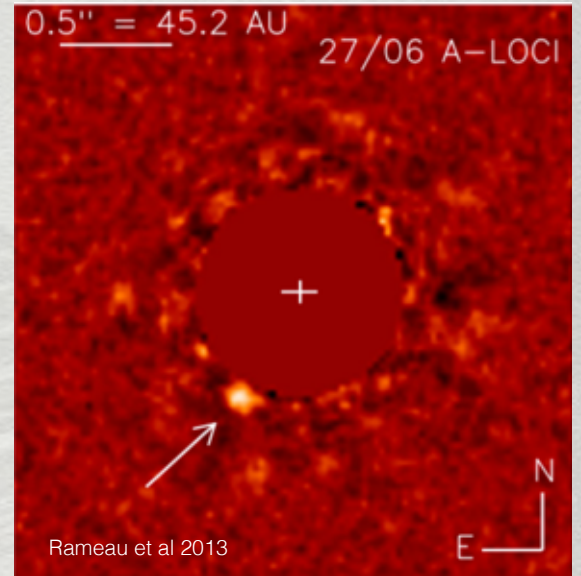
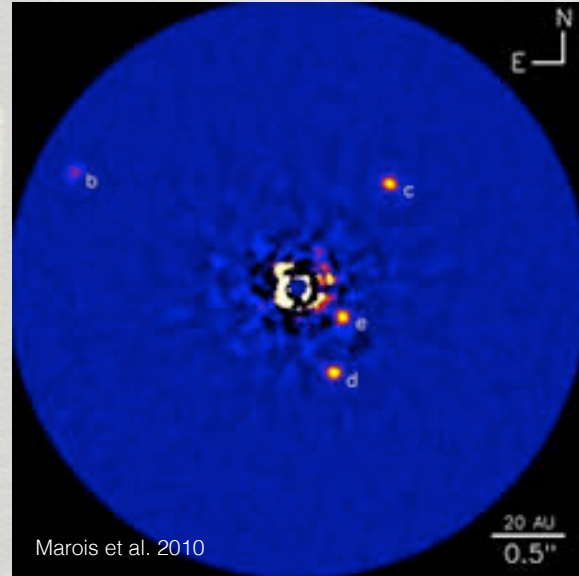


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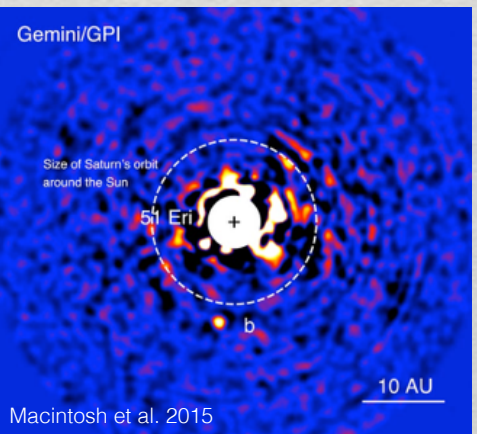
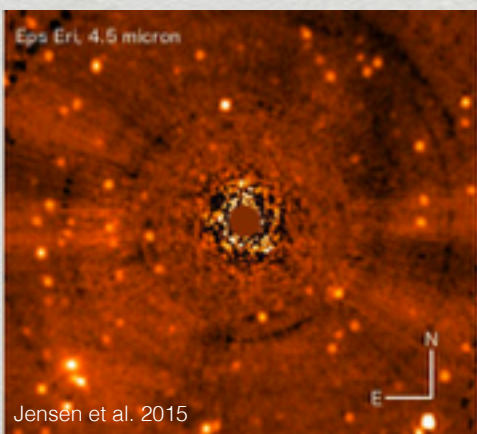
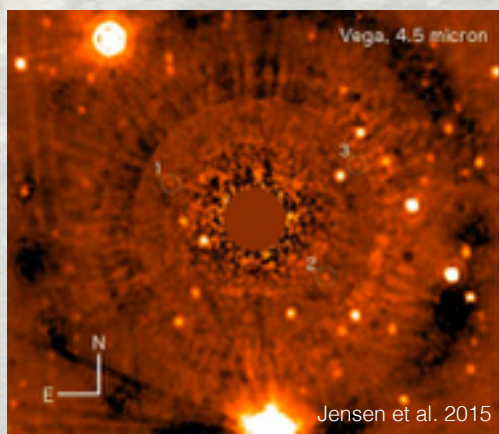
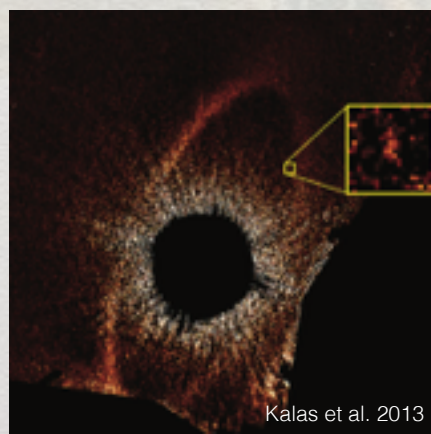
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- HR 8799
- 51 Eri
- Vega
- Fomalhaut
- Eps Eri



## Moonshots



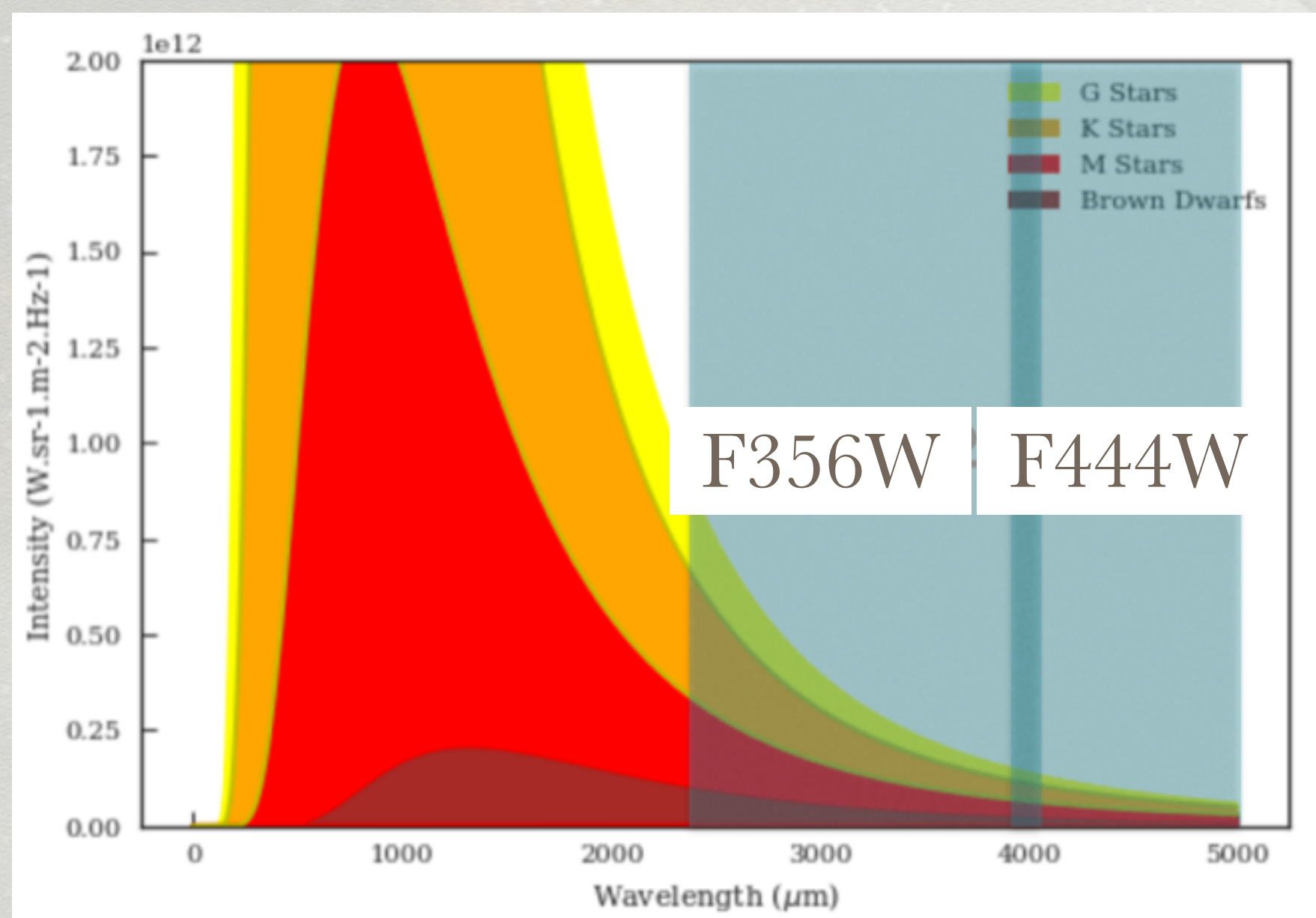
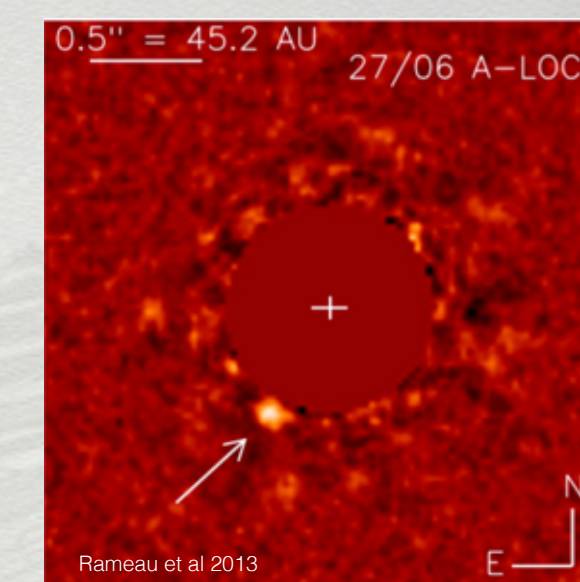
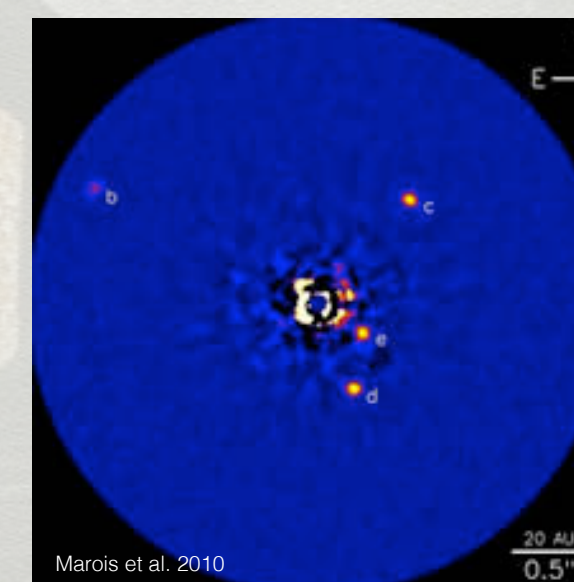


# Search for new worlds

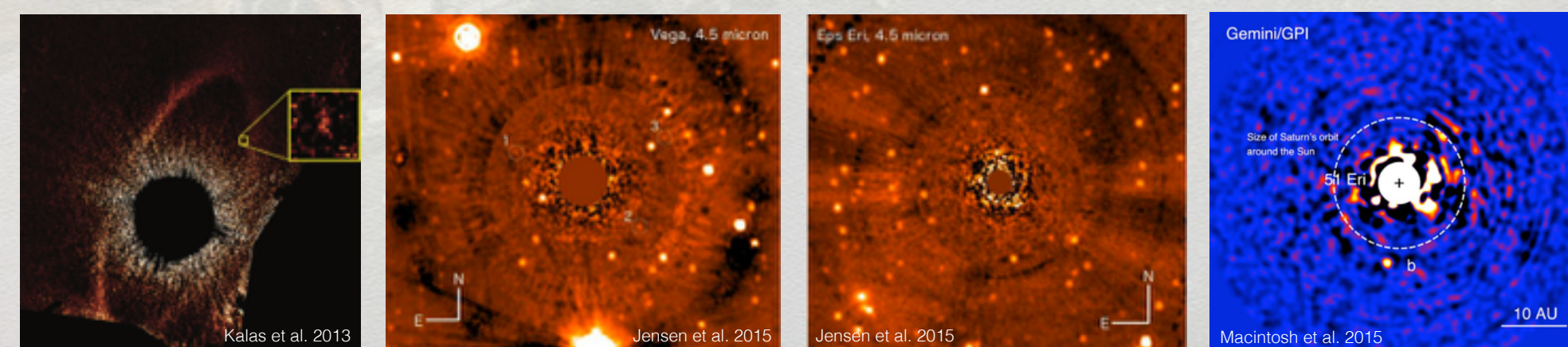
- Extend to lower mass limits compared to ground-based observations
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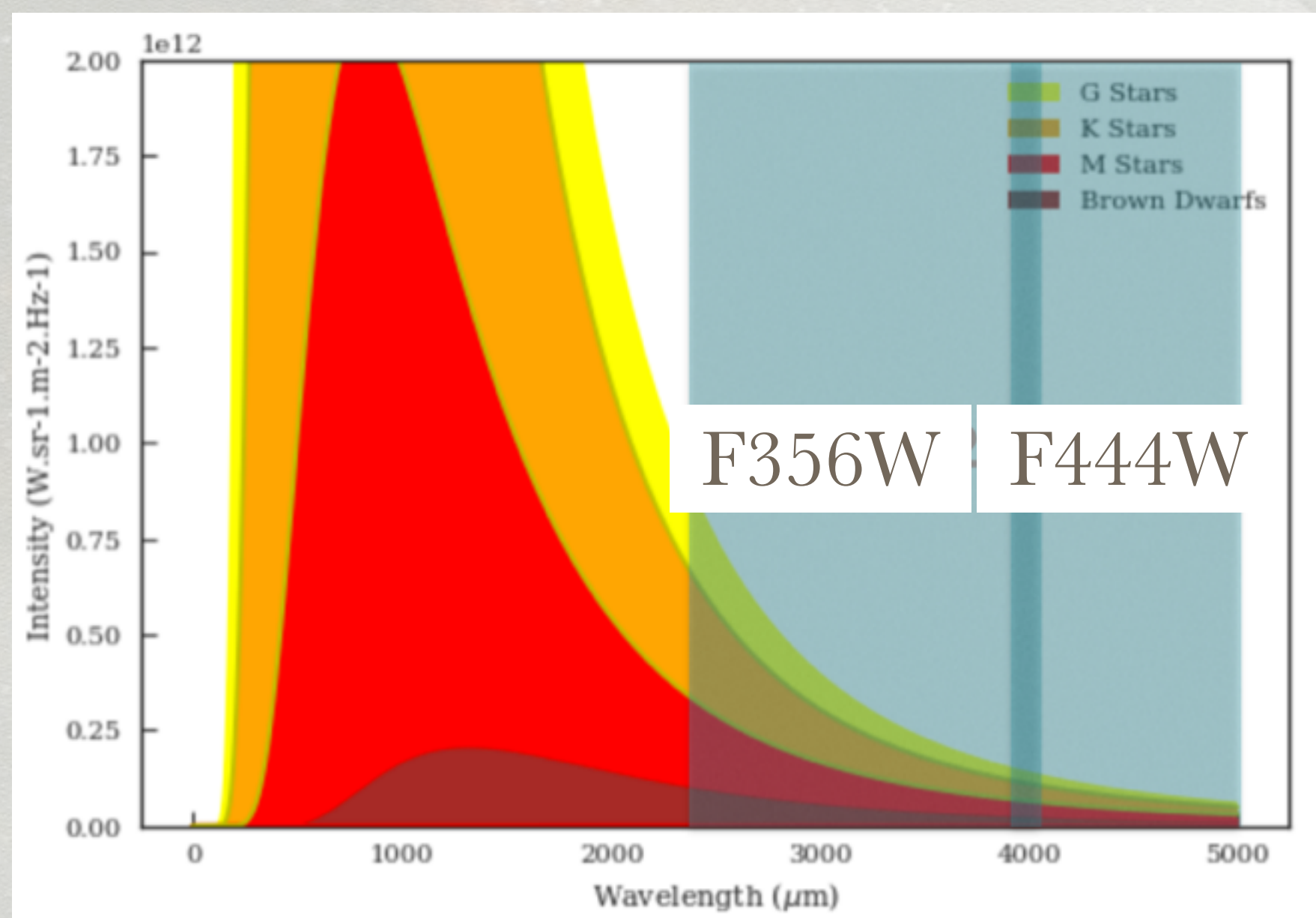
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CALCULATED PLANET MAGNITUDES (10 pc) FOR 50 MYR

$M_{\text{pl}}$ ( $M_{\odot}$ )	$T_{\text{eff}}$ (K)	$L_{\text{pl}}$ ( $L_{\odot}$ )	$\log(g)$ (cgs)	F356W (mag)	F444W (mag)
0.0001	139	-8.453	2.41	20.33	17.29
0.0002	196	-7.794	2.65	19.65	16.7
0.0003	232	-7.465	2.79	19.23	16.32
0.0004	262	-7.232	2.89	18.9	16.04
0.0005	285	-7.069	2.97	18.68	15.85
0.001	375	-6.587	3.27	17.88	15.16
0.002	491	-6.1	3.55	16.51	13.97
0.003	585	-5.789	3.72	15.75	13.36
0.004	676	-5.531	3.84	15.01	12.82
0.005	756	-5.336	3.93	14.44	12.44
0.006	840	-5.152	4.01	13.91	12.11
0.007	928	-4.978	4.08	13.39	11.8
0.008	1010	-4.832	4.14	12.96	11.55
0.009	1085	-4.706	4.19	12.59	11.36
0.010	1171	-4.569	4.23	12.21	11.17

Adapted from Beichuan et al. 2010: Extension of COND03 models (Baraffe et al. 2003) for planetary masses down to 0.1  $M_{\text{Jup}}$



# Search for new worlds

- Exposure time computations for searches up to Saturn Masses at 4'' with a 10 nm wavefront drift using the extension of COND03 models (Baraffe et al. 2003) for planetary masses down to 0.1 MJup (Beichman et al. 2010)

Target	Mass	Exposure Time	SNR >
HD 95086	Jupiter	~2000	6
HR 8799	Saturn	~3200	4
51 Eri	Saturn	~2000	12
Vega	Saturn	~3800	9
Fomalhaut	Saturn	~3800	18
Eps Eri	Saturn	~3800	136



# Take Home Messages



- Direct imaging is essential to further understand exoplanetology



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- Direct imaging is hard, not a lot of detected planets



# Take Home Messages

- Direct imaging is essential to further understand exoplanetology
- Direct imaging is hard, not a lot of detected planets
- The "Coronagraphic Imaging of Young Planets" program has been designed to do what NIRCcam will do best: characterization of known exoplanets and search for new worlds at lower mass limits down to at least Saturn masses

**First results in about 20 months, ... stay tuned**